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the first two years of life. The first year of life is characterized by rapid growth and development, and the second year by continued growth and the beginning of walking and talking.

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NAVIGATION
AND
NAUTICAL ASTRONOMY:

CONTAINING
PRACTICAL RULES, NOTES, AND EXAMPLES.

BY
THE REV. R. M. INSKIP,

H.M.S. BRITANNIA.



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PREFACE.

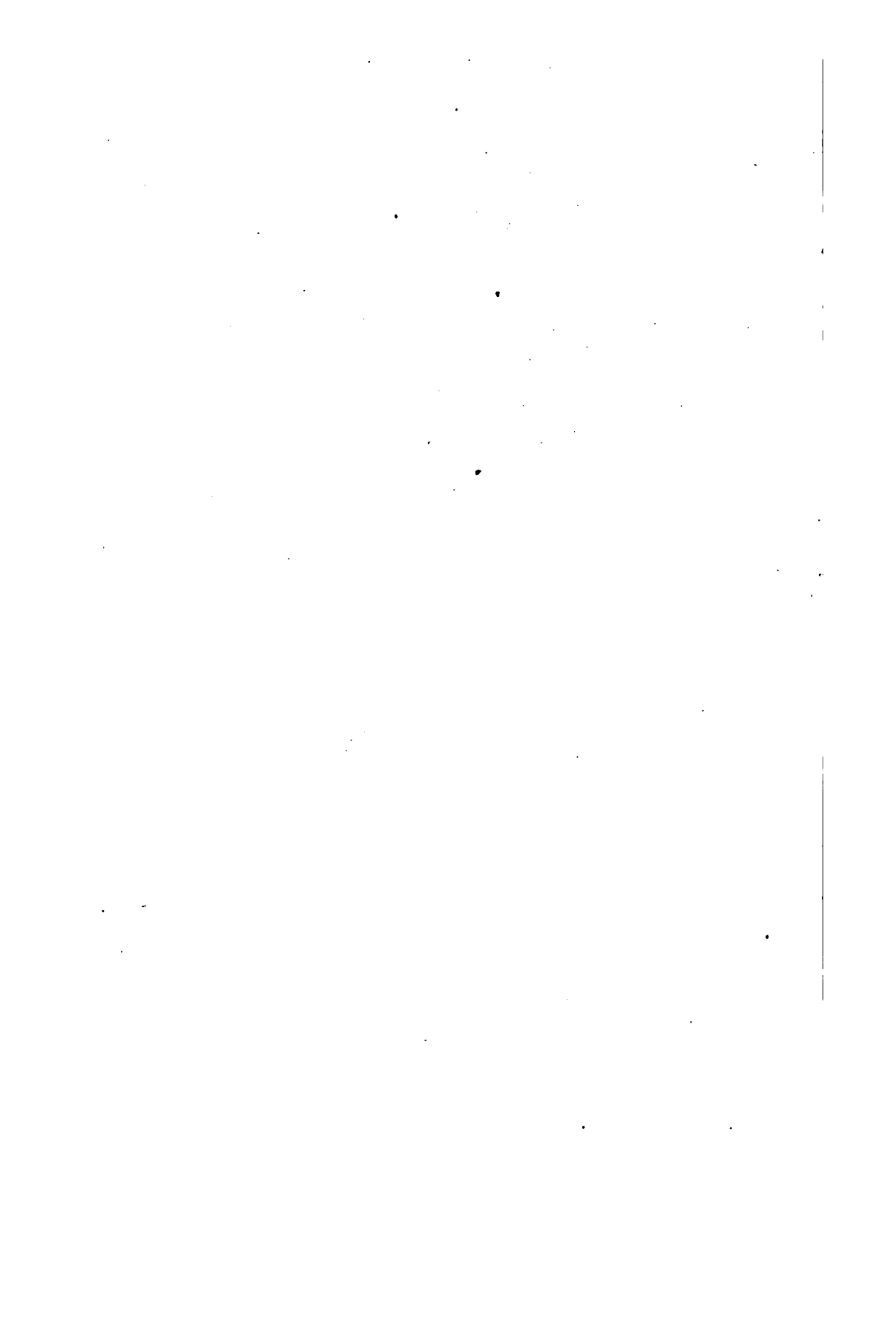
THE following pages include, in a practical form, all the most useful Rules of Navigation and Nautical Astronomy.

The Rules have been made as clear and easy to understand as the nature of the subject would admit.

Notes are added to the Rules to render them applicable to particular cases; and numerous Examples with their Answers are given.

An extract from the Nautical Almanac is annexed, containing the Elements used in working the Examples.

January, 1869.



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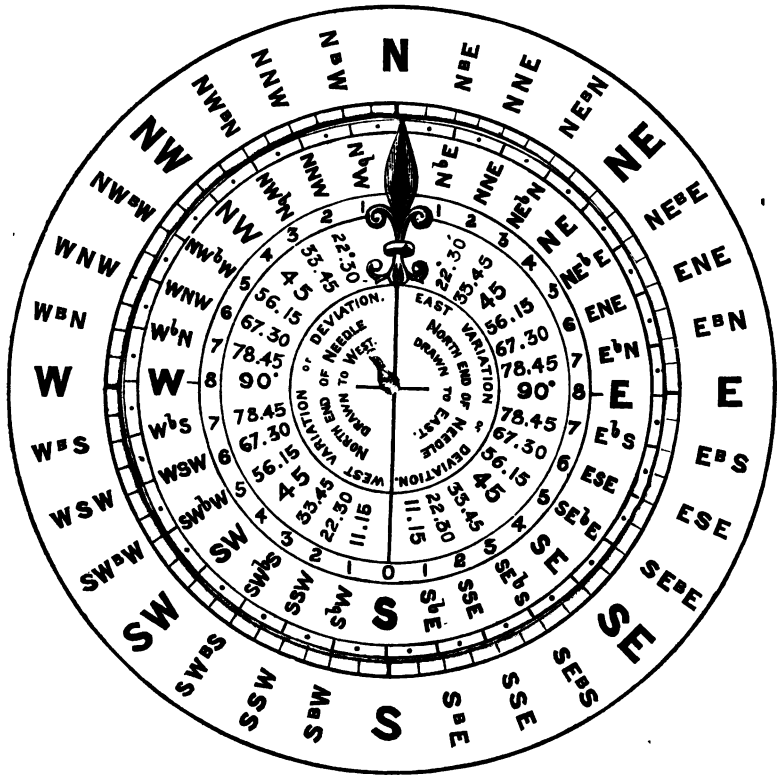
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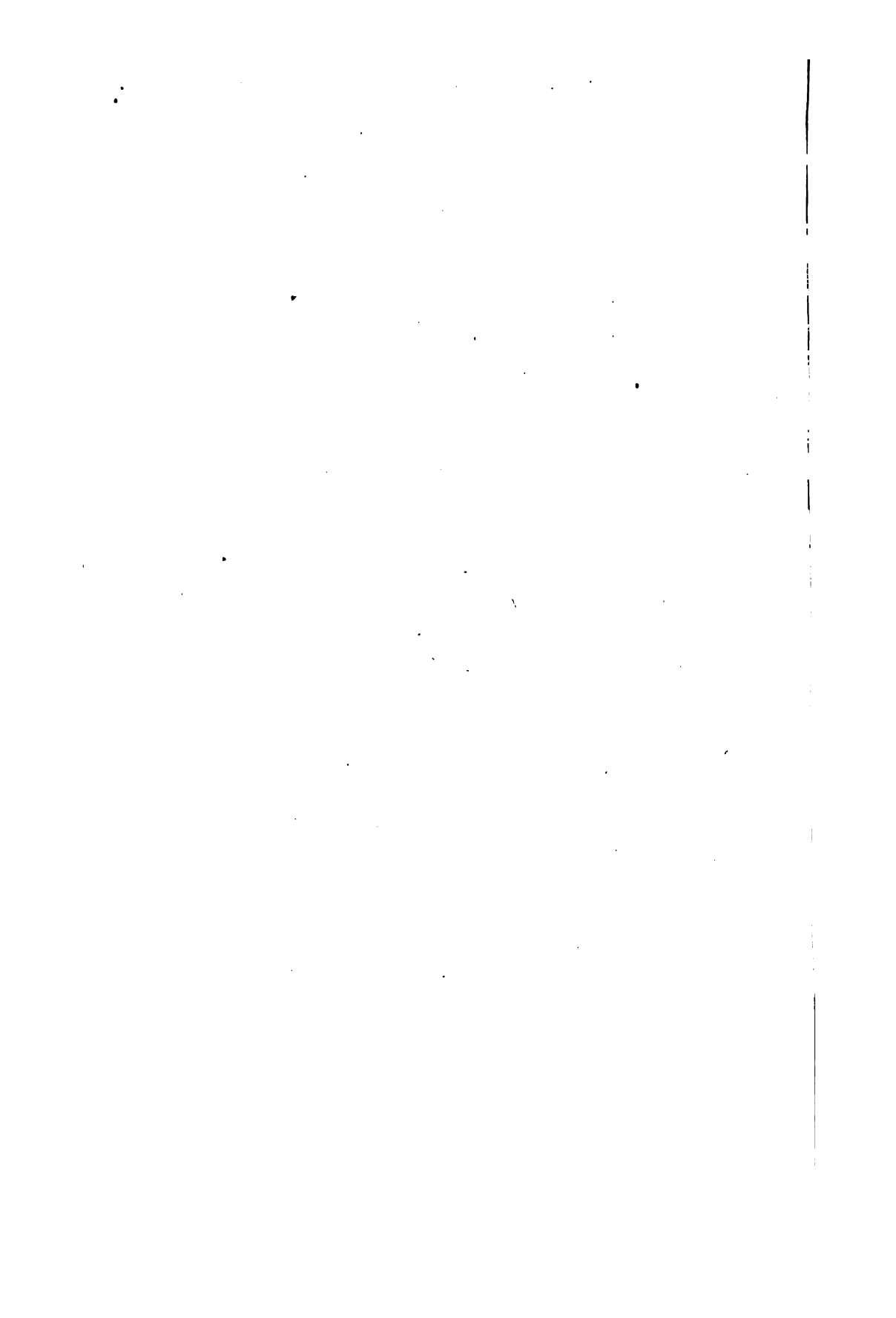
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NAVIGATION.

NAVIGATION refers more especially to that branch of the art of conducting a ship from one part of the world to another which is performed by means of the mariner's compass and log line ; and is usually divided into the following Rules, called the "Sailings," viz., Plane, Parallel, Traverse, Middle Latitude, and Mercator's Sailing, all of which are solved by the application of the rules of Plane Trigonometry.

PLANE SAILING.

IN this sailing, the distance being limited to so small a portion of the Globe as an ordinary day's run, no error of consequence can arise by considering

The meridians to be *straight* lines parallel to each other.

The parallels of latitude also to be *straight* lines cutting the meridians at right angles.

The rhumb lines* also *straight* lines.

The *figure* for plane sailing is a right-angled triangle, in which the vertical side (north and south line) represents the *difference of latitude*.

The horizontal side (east and west line) represents the *departure*.

The hypotenuse represents the *distance*.

The *angle* contained between the meridian and the distance represents the course.

Hence all plane sailing questions are merely solutions of *right-angled plane triangles*.

The two cases usually required in this sailing are :—

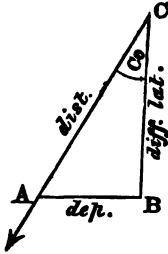
(1) When the course and distance are given, to find the difference of latitude and departure.

(2) When the difference of latitude and departure are given, to find the course and distance.

* The *rhumb* line is that line upon the surface of the terrestrial globe which makes the same angle with all the *meridians*, as it crosses them in succession.

PLANE SAILING.

Example 1.—A ship sails S. 30° W. until her distance is 120 nautical miles; find the difference of latitude and departure she has made.



Given { Co. S. 30° W. } to find { diff. lat.
dist. 120m. } and dep.

To find the diff. lat.

$$\frac{BC}{AC} \text{ or } \frac{\text{diff. lat.}}{\text{dist.}} = \cos. \text{ course.}$$

$$\therefore \text{diff. lat.} = \text{dist.} \cos. \text{ course. (1)}$$

$$\therefore \log. \text{diff. lat.} = \log. \text{dist.} + \text{Log. cos. co.} - 10.$$

	(1)	
log. dist. 120	2.079181	
Log. cos. co. 30°	9.937531	
	2.016712	
log. diff. lat.	2.016712	

$$\therefore \text{diff. lat.} = 103.9 \text{ miles.}$$

To find the dep.

$$\frac{AB}{AC} \text{ or } \frac{\text{dep.}}{\text{dist.}} = \sin. \text{ course.}$$

$$\therefore \text{dep.} = \text{dist.} \sin. \text{ course. (2)}$$

$$\therefore \log. \text{dep.} = \log. \text{dist.} + \text{Log. sin. co.} - 10.$$

	(2)	
log. dist. 120	2.079181	
Log. sin. co. 30°	9.698970	
	1.778151	
log. dep.	1.778151	

$$\therefore \text{dep.} = 60 \text{ miles.}$$

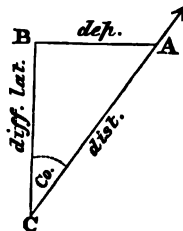
Note.—In every log. sin., log. cos., &c., the Tabular logarithm has 10 added to its *index*, and therefore 10 must be subtracted from the results above.

To find the diff. lat. and dep. by Traverse Table (by Inspection).

Enter the Traverse Table with the given distance 120, at the top (as dist.), and on the left hand side of the page look for 30°, as a course, and the corresponding numbers in the diff. lat. column will be 103.9, and in the dep. column will be 60, the diff. lat. and dep. required.

Example 2.—A ship makes 80 miles diff. lat. (northing), and 55 miles dep. (easting) ; find the course and distance.

Given $\left\{ \begin{array}{l} \text{diff. lat.} = 80 \text{ m.} \\ \text{dep.} = 55 \text{ m.} \end{array} \right\}$ to find $\left\{ \begin{array}{l} \text{Course} \\ \text{and} \\ \text{dist.} \end{array} \right\}$



To find the course.

To find the dist.

$$\tan. \text{ co.} = \frac{AB}{BC} \text{ or } \frac{\text{dep.}}{\text{diff. lat.}} \quad (1) \quad \frac{AC}{BC} \text{ or } \frac{\text{dist.}}{\text{diff. lat.}} = \text{Sec. course.}$$

$$\therefore \text{dist.} = \text{diff. lat. sec. co.} \quad (2)$$

$\therefore \text{Log. tan. co.} - 10 = \text{log. dep.} - \text{log. diff. lat.} \quad \therefore \text{log. dist.} = \text{log. diff. lat.} + \text{Log. sec. co.} - 10.$
 or, $\text{Log. tan. co.} = 10 + \text{log. dep.} - \text{log. diff. lat.}$

	(1)		(2)	
10 + log. dep.	55	11.740363	log. diff. lat. 80	1.903090
log. diff. lat. 80		1.903090	Log. sec. co. 34° 30' 30"	10.084050
		9.837273	log. dist.	1.987140
			\therefore course N. 34° 30' 30" E.*	\therefore dist. 97 miles.

To find the course and distance by Traverse Table (by Inspection).

Seek in the Traverse Table for diff. lat. 80, and dep. 55, and find where these numbers, or the nearest numbers to those come together. In this instance we cannot find them to agree exactly, but have 80.4 in the diff. lat. column, and 54.2 in the dep. column, which give the course 34° (on the left hand side of the page), and the dist. 97 (at the top of the page). Hence the course is N. 34° E. (nearly), and the dist. 97m.

* To determine the *direction* of the course sailed. By the question the ship has made *northing* and *easting*, she has therefore gone *between* N. and E., having made an angle of 34° 30' 30" with the *meridian*, which angle we must indicate by the letters N. and E. to denote the quarter of the compass sailed in : hence the *course* is N. 34° 30' 30" E.

PLANE SAILING.

EXAMPLES FOR EXERCISE.

1. A ship sails from Lisbon N.W. by W., 128m. ; required her diff. lat. and dep.—*See* Ex. 1, page 2.

Ans. Diff. lat. $71^{\circ}1'$ miles, dep. 106.4 miles.

2. A ship from lat. $22^{\circ}55'$ S. sails S.E. $\frac{1}{2}$ E. 80 miles ; find her lat. arrived at and dep. made.—*See* Ex. 1, page 2, and Rule (a) page 12.

Ans. Lat. arrived at $23^{\circ}45'$ S., dep. 61.8 miles.

3. A ship from lat. $42^{\circ}20'$ N. sails S. 80° E., for 20h. at the rate of 9 knots ; find her lat. in and dep. made.

Ans. Lat. arrived at $41^{\circ}48'$ N., dep. 177.3 miles.

4. A ship from lat. $1^{\circ}20'$ N. sails S. 20° E., 200m. ; required her lat. in and dep. made.

Ans. Lat. in $1^{\circ}48'$ S., dep. 68.4 miles.

5. A ship sails N. 110m. and then E. 60m. ; required her course and dist. made good.—*See* Ex. 2, page 3.

Ans. Course made good N. $28^{\circ}36'30''$ E., dist. 125.3 miles.

6. A ship sails from lat. $20^{\circ}10'$ N. until she arrives in lat. $23^{\circ}24'$ N., having made 140m. of easting ; required her course and distance made good.

Ans. Course made good N. $35^{\circ}49'0''$ E. dist. 239.2.

7. A ship in lat. $48^{\circ}30'$ S. sails on that parallel until she is 78m. to the westward of the place started from, and then steers S. until she arrives in lat. $50^{\circ}20'$ S. ; required her course and distance made good.

Ans. Course made good S. $35^{\circ}20'30''$ W., dist. 134.8 miles.

8. A ship sails S. 50m. and then E. 50m. ; required the course and distance made good.

Ans. Course S.E., or S. 45° E., dist. 70.7 miles.

9. A ship from San Francisco in lat. $37^{\circ}48'$ N. sails in the N.W. quarter 115m., and then finds that she is 80m. west of the

meridian of that place ; required the course steered, and latitude of the ship.

Ans. Course steered N. $44^{\circ} 4' 45''$ W., lat of ship $39^{\circ} 11'$ N.

10. A ship from lat. $50^{\circ} 35'$ N. sails N. 20° W., until she has elevated the N. Pole 2° ; required her lat. in and also the dist. and dep. made good.

Ans. Lat. in $52^{\circ} 35'$ N., dist. 127.7 miles, dep. 43.7 miles.

11. A ship from lat. $60^{\circ} 40'$ S. sails N. by E. until she has depressed the S. Pole 3° ; required her lat. in, and distance run.

Ans. Lat. in $57^{\circ} 40'$ S., dist. 183.5 miles.

12. A ship from lat. $30^{\circ} 20'$ S. sails 100m. in the S.W. quarter, and finds that she is, in lat. $31^{\circ} 40'$ S.; required the course and dep.

Ans. Course S. $36^{\circ} 52' 15''$ W., dep. 60 miles.

13. Two ports are on the same parallel of lat., and 58m. distant. A ship sails on the meridian of the easterly one until she is 64m. north of it, and then speaks a barque that had sailed from the westerly port ; required the course and distance of the barque.

Ans. Course N. $41^{\circ} 11'$ E., dist. 86.4 miles.

14. Two ships under the same meridian and 86 miles distant are to the eastward of a port, which is 50 miles from the northernmost ship, and on the same parallel with her ; required the course and dist. of each ship to reach the port.

Ans. 1st course W. dist. 50m., 2nd course N. $30^{\circ} 10' 30''$ W., dist. 99m.

15. A ship on the meridian of Fayal, in lat. $38^{\circ} 32'$ N., sails from lat. $35^{\circ} 53'$ N. on a direct course until she meets a brig which had run due W. 80m. from that place ; required the course steered by the ship, and her distance run.

Ans. Course steered N. $26^{\circ} 42' 30''$ W., dist. 178 miles.

PARALLEL SAILING.

When a ship sails on a parallel of latitude from a given meridian, *she changes her longitude only*; but the *distance between any two meridians* varies according to the parallel sailed on, and since longitude is measured on the equator, the spherical form of the Earth's surface must be taken into consideration before we can determine the relation between the departure, the difference of longitude, and the latitude of the parallel.

The method by which this relation is ascertained is called *parallel sailing*.

In this sailing a ship's track is supposed to be on a *parallel of latitude*, as from A to B. (See Fig. Ex. A.) The latitude of A or of B is represented by AD or BL.

The *diff. long.* between A and B, is the arc of the equator DL.

The arc AB, is the *departure* (or meridian distance) between A and B.

By Spherical Trigonometry.

It may be proved that,

$$\frac{\text{The arc AB of the parallel}}{\text{The arc DL of the equator}} = \cos. AD \text{ or } \cos. BL. \quad (\text{See Fig. Ex. A.})$$

that is, $\frac{\text{dep.}}{\text{diff. long.}} = \cos. \text{ lat.}$; whence the following formulæ.

- | | | |
|--|---|---|
| (A) $\frac{\text{dep.}}{\text{diff. long.}} = \cos. \text{ lat.}$
∴ $\text{dep.} = \text{diff. long.} \cos. \text{ lat.}$ | } | When the <i>lat.</i> of the parallel and <i>diff. long.</i> between the two places are given; to find the <i>dep.</i> (or mer. dist.) between the two places situated on that parallel. |
| (B) $\frac{\text{diff. long.}}{\text{dep.}} = \sec. \text{ lat.}$
∴ $\text{diff. long.} = \text{dep.} \sec. \text{ lat.}$ | } | When the <i>dep.</i> (or mer. dist.) between the two places, and the <i>lat.</i> of the parallel are given; to find the <i>diff. long.</i> between the two places. |
| (C) $\cos. \text{ lat.} = \frac{\text{dep.}}{\text{diff. long.}}$ | } | When the <i>dep.</i> (or mer. dist.) and <i>diff. long.</i> between the two places are given; to find the <i>lat.</i> of the parallel on which the two places are situated. |

EXAMPLE (A).

A ship on the parallel of 50° N. (or S.) latitude, makes 200 miles difference of longitude; find the corresponding departure (or meridian distance).

Given $AD = BL = \text{lat. } 50^\circ$.
 $DL = \text{diff. long.} = 200 \text{ m.}$

To find $AB = \text{required departure}$
 (or meridian distance).

Then $\frac{AB}{DL} = \cos. AD$.

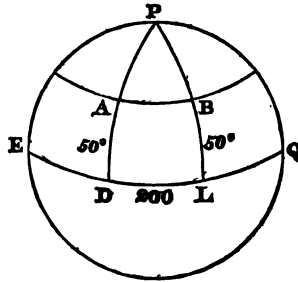
$\therefore \frac{\text{dep.}}{\text{diff. long.}} = \cos. \text{ lat.}$

$\therefore \text{dep.} = \text{diff. long.} \times \cos. \text{ lat.}$

$\therefore \log. \text{ dep.} = \log. \text{ diff. long.} + \text{Log. cos. lat.} - 10$.

log. diff. long. 200	2'301030
Log. cos. lat. 50°	9'808067
log. dep.	<u>2'109097</u>

$\therefore \text{dep. (or mer. dist.)} = 128.5 \text{ miles.}$



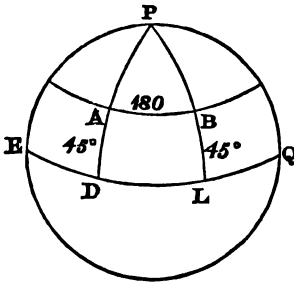
By Inspection.

With the latitude of the parallel as a *course*, and the diff. long. as a *distance*, enter the Traverse Table, and the corresponding *diff. lat.* will be the departure required.

Example.—Enter the Traverse Table with 50° as a *course*, and 200 in the *dist.* column: the corresponding *diff lat.* 128.6, will be the departure required.

EXAMPLE (B).

A ship on the parallel of 45° N. latitude, makes 180 miles of departure: find the corresponding difference of longitude.



Given $AB = \text{dep. (or mer. dist.)} = 180\text{m.}$

$AD = BL = \text{lat. } 45^\circ.$

To find $DL = \text{required diff. long.}$

Then $\frac{DL}{AB} = \text{sec. AD};$

or, $\frac{\text{diff. long.}}{\text{dep.}} = \text{sec. lat.}$

$\therefore \text{diff. long.} = \text{dep.} \times \text{sec. lat.}$

$\therefore \text{log. diff. long.} = \text{log. dep.} + \text{Log. sec. lat.} - 10.$

log. dep. 180 2.255273

Log. sec. lat. 45° 10.150515

log. diff. long. 2.405788

$\therefore \text{diff. long.} = 254.5 \text{ miles.}$

By Inspection.

With the latitude of the parallel as a *course*, and the dep. as *diff. lat.*, enter the Traverse Table, and the corresponding *dist.* will be the diff. long. required.

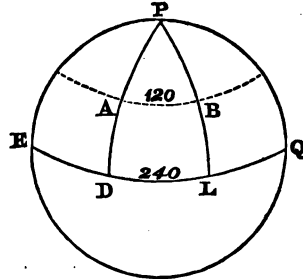
Example.—Enter the Traverse Table with 45° as *course*, and 180 (or the nearest number to it) as *diff. lat.*, the corresponding *distance*, 255, will be the diff. long. required.

EXAMPLE (C).

On what parallel of latitude must a ship sail 120 miles due E., or due W., to change her longitude 240 miles?

Given AB = dep. (or mer. dist.) = 120 m.
DL = diff. long. = 240 m.

To find AD = BL = latitude required.



$$\text{Then cos. lat.} = \frac{\text{dep.}}{\text{diff. long.}}$$

$$\therefore \text{Log. cos. lat.} - 10 = \text{log. dep.} - \text{log. diff. long.}$$

$$\text{or, Log. cos. lat.} = 10 + \text{log. dep.} - \text{log. diff. long.}$$

10 + log. dep.	120	12°079181
log. diff. long.	240	2°380211
		9°698970
Log. cos. lat.		9°698970

$$\therefore \text{Latitude of parallel} = 60^\circ.$$

By Inspection.

With the diff. long. as *distance*, and the dep. as *diff. lat.*, enter the Traverse Table, and the corresponding *course* will be the latitude required.

Example. Enter the Traverse Table with 240 as *dist.* and 120 (or the nearest number to it) as *diff. lat.*, and the corresponding *course*, 60°, will be the latitude required.

Note.—We can also solve these problems by constructing a right-angled plane triangle, in which the base represents the departure, the angle at the base the *latitude* of the parallel, and the hypotenuse represents the *difference of longitude*.

PARALLEL SAILING.

EXAMPLES FOR EXERCISE.

1. A ship on the parallel of 55° N. sails due E. from long. $35^{\circ} 30'$ W. to long. $25^{\circ} 30'$ W. ; find the meridian distance (dep.) she has made.—*See* Example (A) page 7.

Ans. Dep., 344.1m.

2. A ship is on the parallel of Cape Race, in lat. $46^{\circ} 40'$ N., and long. $47^{\circ} 8'$ W. ; how far must she sail in nautical miles to reach the Cape ?

Ans. Dist. to sail, 246.3m.

3. Two ships in 60° S. are 200 nautical miles apart ; what is their difference of longitude ?—*See* Example (B) page 8.

Ans. Diff. long., 400m., or $6^{\circ} 40'$.

4. How far must a ship sail to reach the meridian of Valparaiso, in long. $71^{\circ} 41'$ W., if she is 10 miles *due* south of Cape Horn, in lat. $55^{\circ} 59'$ S., and long. $67^{\circ} 16'$ W. ?

Ans. Dist. to sail, 147.9m.

5. The diff. long. between two ships is 500m., and their meridian distance is 250m. ; what parallel are they on ?—*See* Example (C) page 9.

Ans. The parallel of 60° .

6. A ship on the meridian of St. Paul's, in long. $77^{\circ} 35'$ E., and in lat. 54° S., sails due W. until she is on the meridian of Port Louis (Isle of France), in long. $57^{\circ} 32'$ E. ; how many nautical miles must she have run ?

Ans. Dist. run, 707.1m.

7. In what latitude will 300 miles of dep. equal 600 miles of diff. long. ?

Ans. In lat. 60° .

8. How far must a ship sail due W. from Ascension (Atlantic), in lat. $7^{\circ} 55'$ S., and long. $14^{\circ} 25'$ W., to change her long. 3° ?

Ans. Dist. 178.3m.

9. Being 150m. (in diff. long.) east of Halifax, in lat. $44^{\circ} 40'$ N., how far must I sail to reach the harbour?

Ans. Dist. to sail, 106.7m.

10. In what latitude must a ship be that each mile of dep. (per. dist.) shall be equal to two miles of diff. long.?

Ans. In lat. 60° north or south.

11. How far must a ship sail due W. from Cape Clear, in lat. $51^{\circ} 26'$ N., and long. $9^{\circ} 22'$ W., to reach the meridian of Funchal (Madeira), in long. $16^{\circ} 55'$ W.?

Ans. Dist. to sail, 280m.

12. What is the length of a degree (in nautical miles) on the parallel of 70° ?

Ans. Length of a degree, 20.52m.

13. How much longitude will a ship run down in a day, when in the lat. of Cape Horn, $55^{\circ} 59'$ S., if she sails due W. at the rate of 10 knots per hour?

Ans. $7^{\circ} 9'$.

14. Two vessels leave Fayal, in lat. $38^{\circ} 32'$ N., and long. $28^{\circ} 38'$ W., bound to Halifax, in lat. $44^{\circ} 40'$ N., and long. $63^{\circ} 35'$ W., one sails *due N.* until she is in the lat. of Halifax, and then steers *due W.* until she reaches it; the other sails *due W.* until she is in the long. of Halifax, and then steers *due N.* until she reaches it; which will arrive there first if they both sail at the same rate, and how many miles will she gain on the other?

Ans. The former vessel, and will gain 149 miles.

To find the latitude arrived at (lat. in) when the latitude left and diff. lat. are given.

RULES.*

(a) When the lat. left and diff. lat. are of the *same* name, their *sum* will be the lat. arrived at.

(b) When the lat. left and diff. lat. are of *contrary* names, take their difference, and the remainder will be the lat. arrived at, and of the name of the greater.

EXAMPLES.

1. A ship from a place A, in lat. $17^{\circ} 40'$ N. makes 130 miles diff. lat. (northing); find the latitude arrived at.

$$\begin{array}{r} \text{lat. left} \qquad \qquad \qquad 17^{\circ} 40' \text{ N.} \\ \text{diff. lat. } 130' = \qquad \qquad \qquad \underline{2 \quad 10 \text{ N.}} \\ \text{lat. arrived at, or lat. in} \quad 19 \quad 50 \text{ N.} \end{array}$$

2. A ship leaves a place A, on the equator, and makes 114 miles diff. lat. (southing); what is her latitude in?

$$\begin{array}{r} \text{lat. left} \qquad \qquad \qquad 0^{\circ} \quad 0' \\ \text{diff. lat. } 114' = \qquad \qquad \qquad \underline{1 \quad 54 \text{ S.}} \\ \text{lat. in (arrived at)} \quad 1 \quad 54 \text{ S.} \end{array}$$

3. A ship from a place A, in lat. $21^{\circ} 15'$ S., makes 160 miles diff. lat. (north); what is her lat. arrived at?

$$\begin{array}{r} \text{lat. from} \qquad \qquad \qquad 21^{\circ} 15' \text{ S.} \\ \text{diff. lat. } 160' = \qquad \qquad \qquad \underline{2 \quad 40 \text{ N.}} \\ \text{lat. arrived at} \qquad \qquad \qquad 18 \quad 35 \text{ S.} \end{array}$$

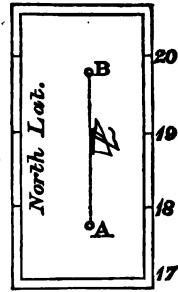
4. A ship from a place A, in lat. $1^{\circ} 44'$ N., makes 235 miles diff. lat. (south); what is her lat. arrived at?

$$\begin{array}{r} \text{lat. left} \qquad \qquad \qquad 1^{\circ} 44' \text{ N.} \\ \text{diff. lat. } 235' = \qquad \qquad \qquad \underline{3 \quad 55 \text{ S.}} \\ \text{lat. arrived at} \qquad \qquad \qquad 2 \quad 11 \text{ S.} \end{array}$$

* In order to remember these Rules, observe that, when a ship is receding from the equator, she is getting to a greater latitude; and when approaching it, to a less. In the former case, we *add* the diff. lat.; in the latter, *subtract* it.

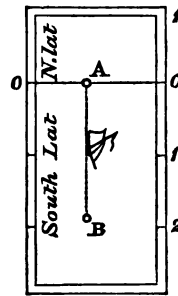
NOTES.

Example 1.—In this example the ship has gone $2^{\circ} 10'$ to the north, and consequently *increased* her distance from the equator.



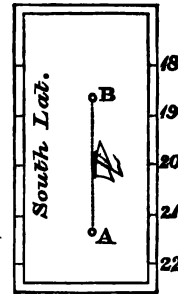
Ex. 1.

Example 2.—In this example the ship has made $1^{\circ} 54'$ diff. lat. to the south, the whole of which was from the equator itself; and therefore her diff. lat. is the same as the lat. arrived at.



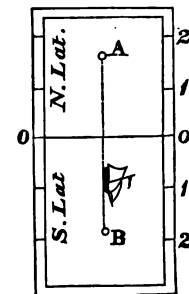
Ex. 2.

Example 3.—In this example the ship was in south lat., and has made $2^{\circ} 40'$ diff. lat. to the north, and consequently neared the equator by that number of miles, and *diminished* her south latitude.



Ex. 3.

Example 4.—In this example the diff. lat., $3^{\circ} 55'$ south, is *greater* than the lat. left, which was *north*; she has, therefore, *crossed* the equator and changed the name of her latitude to south.



Ex. 4.

To find the longitude arrived at (long. in) when the longitude left and diff. long. are given.

RULES.*

(a) When the long. left and diff. long. are of the *same* name, their sum will be the long. in arrived at.

(b) When the long. left and diff. long. are of *contrary* names, the difference between the long. left and diff. long. will be the long. arrived at, and of the name of the greater.

EXAMPLES.

1. A ship from a place A, in longitude $27^{\circ} 50' W.$ makes 50 miles diff. long. west ; required the longitude arrived at.

$$\begin{array}{r} \text{long. left} \qquad 27^{\circ} 50' W. \\ \text{diff. long. } 50' = \quad \underline{0 \quad 50} W. \\ \text{long. arrived at} \quad 28 \quad 40 W. \end{array}$$

2. A ship from a place A, on the meridian of Greenwich, makes 127 miles diff. long. E. ; required her long. in.

$$\begin{array}{r} \text{long. left} \qquad \qquad 0^{\circ} \quad 0' \\ \text{diff. long. } 127' = \quad \underline{2 \quad 7} E. \\ \text{long. in} \qquad \qquad \quad 2 \quad 7 E. \end{array}$$

3. A ship leaves a port A, in longitude $178^{\circ} 26' E.$, and makes 264 miles diff. long. E. ; what is her longitude arrived at ?

$$\begin{array}{r} \text{long. left} \qquad 178^{\circ} 26' E. \\ \text{diff. long. } 264' = \quad \underline{4 \quad 24} E. \\ \qquad \qquad \qquad 182 \quad 50 E. \\ \qquad \qquad \qquad \underline{360} \\ \text{long. in (arrived at)} \quad 177 \quad 10 W. \end{array} \left. \begin{array}{l} \\ \\ \\ \end{array} \right\} \begin{array}{l} \text{In this case the sum of the} \\ \text{long. left and diff. long.} \\ \text{exceeds } 180^{\circ}, \text{ subtract} \\ \text{from } 360^{\circ}, \text{ and change} \\ \text{the name from E. to W.} \end{array}$$

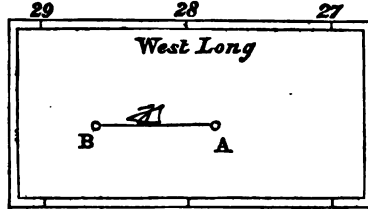
4. A ship from a place A, in longitude $2^{\circ} 30' E.$, makes 330 miles diff. long. W. ; find her longitude arrived at.

$$\begin{array}{r} \text{long. left} \qquad \qquad 2^{\circ} \quad 30' E. \\ \text{diff. long. } 330' = \quad \underline{5 \quad 30} W. \\ \text{long. in (arrived at)} \quad 3 \quad 0 W. \end{array}$$

* If the ship is receding from the first meridian, her longitude is getting *greater* ; if approaching it, *less*.

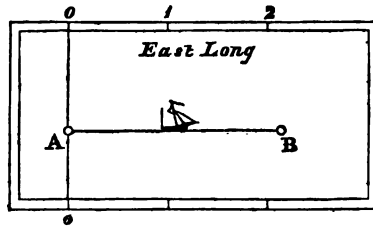
NOTES.

Example 1.—In this example the ship has made 50 m. diff. long. *west*, and consequently *increased* her *west longitude* by 50 m.



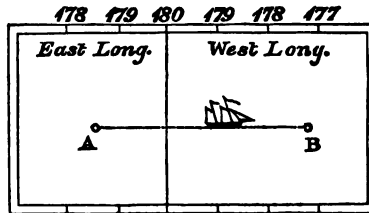
Ex. 1.

Example 2.—In this example the ship left the meridian of Greenwich, and has since made $2^{\circ} 7'$ diff. long. *away* from it to the east; hence her long. arrived at is equal to her diff. long.



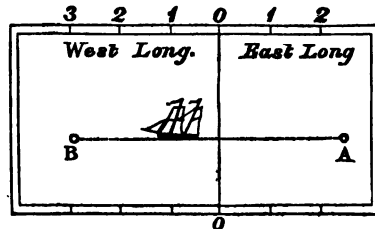
Ex. 2.

Example 3.—In this example the ship has gone from east into west longitude, and by crossing the meridian of 180° , has changed the *name* of her longitude from east to west.



Ex. 3.

Example 4.—In this example the ship has made $5^{\circ} 30'$ diff. long. west, and thereby crossed the meridian of Greenwich, having run from east into west longitude.



Ex. 4.

TRAVERSE SAILING.

IN the navigation of a ship it frequently happens that, from the prevalence of contrary winds and other causes, she cannot be kept on the *same course* for any number of hours, and thus, during a day, or any given period of time, she will sail on several courses. In order to reduce these to a *single course* and distance, as well as to determine the difference of latitude and departure she has made, we have the assistance of a Table so constructed as to give the *diff. lat.* and *dep.* for every *course* from $\frac{1}{4}$ to $7\frac{1}{4}$ points, as well as from 1° to 89° and every *distance* from 1 to 296 miles; (*see* Inman, Table X.)* the use of which is given in the following page.

Draw out a Form (called a Traverse Table), as under.

Corrected Courses.	Dist.	DIFF. LAT.		DEP.	
		N.	S.	E.	W.
(1)	(2)	(3)	(4)	(5)	(6)

In the first column (1) write down the several corrected courses.

In the second column (2) write down the distances made on the several courses.

In the third and fourth columns, (3) and (4), are to be placed the differences of latitude taken from Table X.

In the fifth and sixth columns, (5) and (6), are to be placed the departures, taken from Table X.

* The lower division of the Table gives the Course in Degrees.

Before commencing to take out the diff. lat. and dep. corresponding to the several courses and distances, draw a line ——— through each column that will not be required.

Note 1.—When the distance exceeds 296m. (the limits of the Table) divide it by 2, and take out the diff. lat. and dep. for this new distance, *multiply* them by 2, and place the results in their proper columns as in ordinary cases.

Note 2.—When there is a decimal place in the distance look for it as a whole number in the Traverse Table (x.), and point off one place to the left in the diff. lat. and dep. found, and in this form set them down in the Traverse Table.

To find the DIFF. LAT. and DEP. corresponding to a given course and distance.

When the Course is less than 4 Points.

Enter Table x. with the given course on the *left-hand* side of the page, and look for the distance at the top of the page; take out the corresponding *diff. lat.* and *dep.*, placing the diff. lat. in the N. or S. column, (3) or (4) of the foregoing form, according as the course contains *northing* or *southing*; and the *dep.* in the E. or W. column, (5) or (6), according as the course contains *easting* or *westing*.

When the Course is *more* than 4 Points.

Take the course from the *right-hand* side of the page, in which case the columns of diff. lat. and dep. become reversed, that is, diff. lat. at the *top* of the page, is dep. at the *bottom* of the page; and *dep.* at the top, is *diff. lat.* at the bottom of the page.

When a Course is *due* N. or *due* S.

The distance run will be the *diff. lat.*, and must be placed in the N. or S. column, according as the ship has gone due N. or due S.

When a Course is *due* E. or *due* W.

The distance will be the *dep.*, and must be placed in the E. or W. column, according as the ship has gone due E. or due W.

Course given in *Degrees.*

It must be looked out in the *lower* division of Table x. When not more than 45° , on the left-hand side of the page; and when it exceeds 45° , on the right-hand side of the page.

To find the DIFF. LAT. and DEP. made good.

After the N. S. E. and W. columns have been filled in with the differences of lat. and departure corresponding to the several courses and their distances, add up each column separately. Then take the less diff. lat. from the greater and mark the remainder N. or S., according to which is the greater.

Also take the less dep. from the greater, and mark the result E. or W., according as the easting or westing is the greater.

These remainders will be called the diff. lat. and dep. *made good* for the day, or any other given interval.

EXAMPLE, to illustrate the foregoing Rules.

A ship from latitude $50^\circ 9' N.$, and longitude $12^\circ 35' W.$, sails S.S.W. 25 miles, N.W. $\frac{1}{2}$ N. 30 miles, W. by S. 18 miles, E. 10 miles, and S. 15 miles; find the diff. lat. and dep. she has made.

POINTS from N. or S.	COURSES.	Dist.	DIFF. LAT.		DEP.	
			N.	S.	E.	W.
2	S.S.W.	25	—	23'1	—	09'6
3 $\frac{1}{2}$	N.W. $\frac{1}{2}$ N.	30	23'2	—	—	19'0
7	W. by S.	18	—	03'7	—	18'6
8	E.	10	—	—	10'0	—
0	S.	15	—	15'0	—	—
			23'2	41'8 23'2 — 18'6*	10'0	47'2 10'0 — 37'2†
				* Diff. Lat. made good S.		† Dep. made good W.

Working of the above Example.

Find the number of *points* (reckoned *from N. or S.*) corresponding to each course; then enter Table X. and take out the differences of lat. and departures, as follows:—

1st Course. Enter the Table with (S.S.W.) 2 points, on the *left-hand* side of the page, and 25m. at the top. This will give 23·1 *diff. lat.* to be placed in the S. column (because the ship has made southing), and 09·6 dep. to be placed in the W. column (because the ship has made westing).

2nd Course. Enter the Table with (N.W. $\frac{1}{2}$ N.) 3 $\frac{1}{2}$ points, and dist. 30m., which gives 23·2 for the N. column, and 19·0 for the W. column.

3rd Course. Enter the Table with (W. by S.) 7 points on the *right-hand* side of the page, and distance 18m. at the *top*. This gives 03·7 for the *diff. lat.*, to be placed in the S. column, and 18·6 dep. to be placed in the W. column. (*See* remark, when course is *greater* than 4 points.)

4th Course. The Table is not required in this case, the ship having made all easting, the dist. 10m. must be placed as a *dep.* in the E. column.

5th Course. The Table is not required in this case, the ship having made all southing, the dist. 15m. is the *diff. lat.*, and must be placed in the S. column.

Now add up the N. S. E. and W. columns separately; take the N. *diff. lat.* from the S. *diff. lat.*, and the remainder 18·6 will be the *diff. lat. made good*, and must be marked S. because the ship has made most *southing*.

Also, take the E. dep. from the W. dep., the remainder 37·2 will be the *dep. made good*, and is W. because it was greater than the *easting*.

Hence, the result of all the separate courses and distances gives 18·6 *diff. lat.* S., and 37·2 *dep.* W.

TRAVERSE SAILING.

EXAMPLES FOR EXERCISE.

1. A ship sails N. by E. 30 miles, S.S.W. 15 miles, S.E. by S. 20 miles, N.N.W. 16 miles, S. 14 miles, and N.N.E. 18 miles ; find the diff. lat. and dep. made good.—*See* Example, page 18.

Ans. Diff. lat. $16^{\circ}3$ N. Dep. $12^{\circ}1$ E.

2. A ship sails S.S.E. $\frac{1}{2}$ E. 14 miles, S.W. $\frac{1}{2}$ S. 19 miles, N.E. by N. $\frac{1}{2}$ N. 21 miles, N.W. 14 miles, S.S.E. $\frac{1}{2}$ S. 27 miles, and S. $\frac{1}{2}$ W. 24 miles ; find the diff. lat. and dep. made good.

Ans. Diff. lat. $48^{\circ}3$ S. Dep. $0^{\circ}0$.

3. A ship sails S.W. by S. 16 miles, S.W. by W. 20 miles, N.E. $\frac{1}{2}$ E. 17 miles, E. $\frac{1}{2}$ S. 10 miles, W.N.W. $\frac{1}{2}$ W. 13 miles, and S.S.W. $\frac{1}{4}$ W. 22 miles ; find her diff. lat. and dep., and also her course and distance made good.

Ans. Diff. lat. $30^{\circ}7$ S. Course S. $38^{\circ}14'45''$ W. Dep. $24^{\circ}2$ W.
Dist. 39m.

4. A ship from lat. $16^{\circ}10'$ N., and long. $16^{\circ}10'$ W., sails N.W. by W. 27 miles, N.N.E. $\frac{3}{4}$ E. 60 miles, E. $\frac{1}{2}$ S. 42 miles, and S.S.W. $\frac{1}{2}$ W. 57 miles ; find her course and distance made good, and also her lat. and long. arrived at.

Ans. Course N. $63^{\circ}32'30''$ E. Lat. in $16^{\circ}22'$ N. Dist. $26^{\circ}26$ m.
Long. in $15^{\circ}45'42''$ W.

5. A ship from Fernando de Noronha, in lat. $3^{\circ}52'$ S., sails N. 28° W., 39 miles, N. 45° W. 16 miles, S. 78° W. 27 miles, S. 62° W. 31 miles ; find her course and distance made good, and also the lat. and long. she has reached.

Ans. Course N. $73^{\circ}0'$ W. Lat. in $3^{\circ}26'30''$ S. Dist. $87^{\circ}2$ m.
Long. in $33^{\circ}51'$ W.

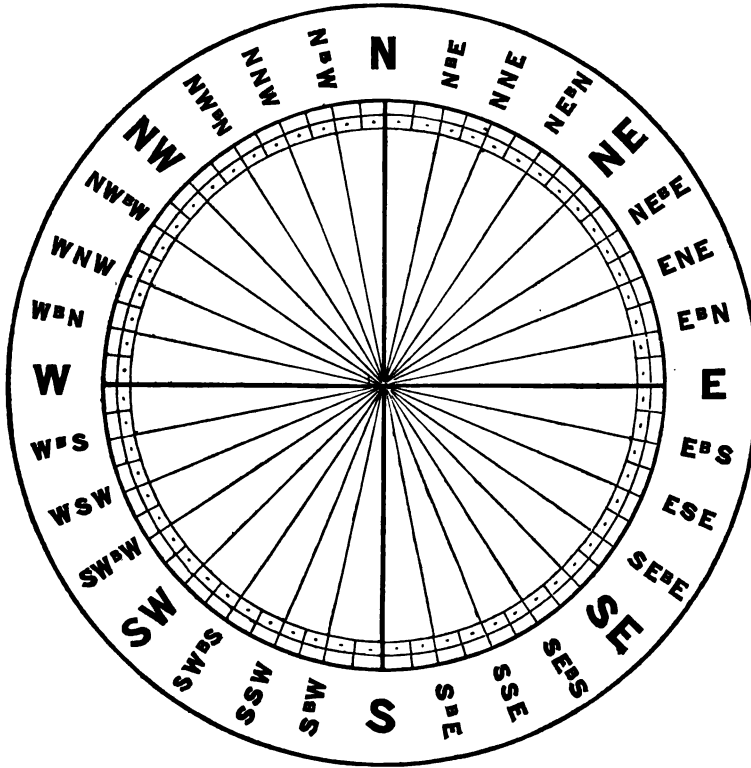
6. A ship on the Equator, and in long. $20^{\circ}16'$ W., sails N.N.E. $\frac{1}{2}$ E. 14 miles, W.S.W. 21 miles, E. $\frac{3}{4}$ S. 30 miles, S.S.E. $\frac{1}{4}$ E. 21 miles, and N.N.W. $\frac{1}{4}$ W. 15 miles ; find her course and distance made good, and also the lat. and long. in.

Ans. Course S. $74^{\circ}15'$ E. Lat. in $0^{\circ}5'30''$ S. Dist. $20^{\circ}26$ m.
Long. in $19^{\circ}56'30''$ E.

7. A ship from lat. $54^{\circ}15'$ S., and on the meridian of Greenwich, sails S. 54° W. 32 miles, S. 71° E. 40 miles, S. 10° W. 44 miles, N. 78° W. 15 miles, and N. 45° E. 27 miles ; find her lat. and long. in.

Ans. Lat. in $55^{\circ}8'$ S. Long. in $0^{\circ}15'$ E.

THE COMPASS.



Points.	From N.		From S.		Degrees.		
	N.	N.	S.	S.	°	'	"
$\frac{1}{4}$	N. $\frac{1}{4}$ E.	N. $\frac{1}{4}$ W.	S. $\frac{1}{4}$ E.	S. $\frac{1}{4}$ W.	2	48	45
$\frac{1}{2}$	N. $\frac{1}{2}$ E.	N. $\frac{1}{2}$ W.	S. $\frac{1}{2}$ E.	S. $\frac{1}{2}$ W.	5	37	30
$\frac{3}{4}$	N. $\frac{3}{4}$ E.	N. $\frac{3}{4}$ W.	S. $\frac{3}{4}$ E.	S. $\frac{3}{4}$ W.	8	26	15
1	N. b E.	N. b W.	S. b E.	S. b W.	11	15	0
2	N. N. E.	N. N. W.	S. S. E.	S. S. W.	22	30	0
3	N. E. b N.	N. W. b N.	S. E. b S.	S. W. b S.	33	45	0
4	N. E.	N. W.	S. E.	S. W.	45	0	0
5	N. E. b E.	N. W. b W.	S. E. b E.	S. W. b W.	56	15	0
6	E. N. E.	W. N. W.	E. S. E.	W. S. W.	67	30	0
7	E. b N.	W. b N.	E. b S.	W. b S.	78	45	0
8	East	West	East	West	90	0	0

TABLE *

For converting Points of the Compass and their fractional parts into Degrees.

Points.	No.	Degrees.	Points.	No.	Degrees.
		° ' "			° ' "
NORTH.	0	0 0 0	N.E.	4	45 0 0
	$\frac{1}{8}$	1 24 22		$\frac{1}{8}$	46 24 22
	$\frac{1}{4}$ —	2 48 45		$\frac{1}{4}$ —	47 48 45
	$\frac{3}{8}$	4 13 7		$\frac{3}{8}$	49 13 7
	$\frac{1}{2}$ —	5 37 30		$\frac{1}{2}$ —	50 37 30
	$\frac{5}{8}$	7 1 52		$\frac{5}{8}$	52 1 52
	$\frac{3}{4}$ —	8 26 15		$\frac{3}{4}$ —	53 26 15
	$\frac{7}{8}$	9 50 37		$\frac{7}{8}$	54 50 37
N. b E.	1	11 15 0	N.E. b E.	5	56 15 0
	$\frac{1}{8}$	12 39 22		$\frac{1}{8}$	57 39 22
	$\frac{1}{4}$ —	14 3 45		$\frac{1}{4}$ —	59 3 45
	$\frac{3}{8}$	15 28 7		$\frac{3}{8}$	60 28 7
	$\frac{1}{2}$ —	16 52 30		$\frac{1}{2}$ —	61 52 30
	$\frac{5}{8}$	18 16 52		$\frac{5}{8}$	63 16 52
	$\frac{3}{4}$ —	19 41 15		$\frac{3}{4}$ —	64 41 15
	$\frac{7}{8}$	21 5 37		$\frac{7}{8}$	66 5 37
N.N.E.	2	22 30 0	E.N.E.	6	67 30 0
	$\frac{1}{8}$	23 54 22		$\frac{1}{8}$	68 54 22
	$\frac{1}{4}$ —	25 18 45		$\frac{1}{4}$ —	70 18 45
	$\frac{3}{8}$	26 43 7		$\frac{3}{8}$	71 43 7
	$\frac{1}{2}$ —	28 7 30		$\frac{1}{2}$ —	73 7 30
	$\frac{5}{8}$	29 31 52		$\frac{5}{8}$	74 31 52
	$\frac{3}{4}$ —	30 56 15		$\frac{3}{4}$ —	75 56 15
	$\frac{7}{8}$	32 20 37		$\frac{7}{8}$	77 20 37
N.E. b N.	3	33 45 0	E. b N.	7	78 45 0
	$\frac{1}{8}$	35 9 22		$\frac{1}{8}$	80 9 22
	$\frac{1}{4}$ —	36 33 45		$\frac{1}{4}$ —	81 33 45
	$\frac{3}{8}$	37 58 7		$\frac{3}{8}$	82 58 7
	$\frac{1}{2}$ —	39 22 30		$\frac{1}{2}$ —	84 22 30
	$\frac{5}{8}$	40 46 52		$\frac{5}{8}$	85 46 52
	$\frac{3}{4}$ —	42 11 15		$\frac{3}{4}$ —	87 11 15
	$\frac{7}{8}$	43 35 37		$\frac{7}{8}$	88 35 37
N.E.	4	45 0 0	EAST.	8	90 0 0

* The remaining Quarters of the Compass may be converted into Points or Degrees by aid of this Table.

VARIATION OF THE COMPASS.

CORRECTING COURSES.

THE direction of the North point of the magnetic (or compass) needle, is not the same in all places, but is drawn aside more or less to the *right* or *left* of the true north point of the world.

The angle included between the direction of the Magnetic and true N. points, is called the VARIATION of the Compass.

[A] When the compass N. is drawn to the *right* of the true N. the variation is easterly.

[B] When the compass N. is drawn to the *left* of the true N. the variation is westerly.

In consequence of this divergence of the magnetic needle from the *true* meridian, the courses steered or bearings observed by compass will not correspond to the courses or bearings given by a chart, since charts are so constructed that the lines upon them shall represent the *true* points of the world, and therefore, whenever it is required to turn a course or bearing by compass into a course or bearing by chart, the variation of the compass must be applied.

The manner in which this is done may be shown as follows :—

Let two compass cards be placed one over the other, so as to have the same centre. The lower one to be *fixed*, and represent the true points of the world, whilst the upper one is *moveable*, and represents the magnetic compass; then, if the two N. points are made exactly to coincide, there would be *no* variation.

Again, suppose the *magnetic* N. to be moved aside 3 points to the *right* of the true N., then the N.W. by N. point of the compass would agree with the N. point of the world, and there would be 3 points easterly variation of the compass. In other words, when the ship is *steering* N.W. by N. by compass, she is in reality making a N. course by the world; and in order

to *correct* the compass course N.W. by N., so as to get N. (true), the variation must be allowed to the *right*.

[C] Hence, to correct a compass course for *easterly* variation, it must be allowed to the *right*, to obtain the *corresponding* true course.

Again, suppose the *magnetic* N. to be moved 4 points to the *left* of the true N., then the N.E. point of the compass would agree with the N. point of the world, and there would be 4 points westerly variation of the compass. In other words, when the ship is steering N.E. by compass, she is in reality making a N. course by the world ; and in order to *correct* the compass course N.E., so as to get N. (true), the variation must be allowed to the *left*.

[D] Hence, to correct a compass course for *westerly* variation, it must be allowed to the *left*, to obtain the *corresponding* true course.

The **true** course denotes the course as given by the *chart*,* unaffected either by variation or by deviation ; and is the course that would be steered by the compass if there were neither variation nor deviation.

The **correct magnetic** course denotes the course as given by the compass, affected by variation *only* ; and is the course that would be steered by the compass if there were no deviation.

The **compass** course is influenced both by variation and by deviation.

* A few of the Admiralty Charts are drawn for convenience on the *magnetic* meridian ; the navigator would of course pay attention to this occasional departure from the general practice of all charts and plans in this country being drawn on the *true* meridian.

CORRECTING MAGNETIC COURSES OR BEARINGS FOR VARIATION.

When the variation is easterly, allow it to the *right* of the correct magnetic course, to obtain the true course.

When the variation is westerly, allow it to the *left* of the correct magnetic course, to obtain the true course.

Note.—When allowing the variation, consider yourself at the CENTRE of the *compass*, and facing the point you wish to correct.

EXAMPLES.

Magnetic Courses.	Variation.	True Courses.	Magnetic Courses.	Variation.	True Courses.
N.N.W.	1 E.	N. b W.	N.N.E.	1 W.	N. b E.
N.W. b N.	1 "	N.N.W.	N.E. b N.	1 "	N.N.E.
N.W. b W.	2 "	N.W. b N.	N.E. b E.	2 "	N.E. b N.
W.S.W.	2 "	West	E.N.E.	2 "	N.E.
S.W.	4 "	West	N.E.	4 "	North
S. b W.	2 "	S.W. b S.	N. b E.	2 "	N. b W.
S. b E.	1 "	South	E. b S.	1 "	East

CORRECTING TRUE COURSES OR BEARINGS FOR VARIATION.

Correcting a *true* course for variation being exactly the *reverse operation* of correcting a correct magnetic course for variation, we have the following rules:—

When the variation is easterly, allow it to the *left* of the true course, to obtain the correct magnetic course.

When the variation is westerly, allow it to the *right* of the true course, to obtain the correct magnetic course.

EXAMPLES.

True Courses.	Variation.	Magnetic Courses.	True Courses.	Variation.	Magnetic Courses.
N.N.E.	2 E.	North	N.N.W.	2 W.	North
N.E. b N.	2 "	N. b E.	N. b W.	2 "	N. b E.
N.E.	1 "	N.E. b N.	N.W. b N.	1 "	N.N.W.
E.N.E.	2 "	N.E.	W.N.W.	1 "	N.W. b W.
East	2 "	E.N.E.	West	2 "	W.N.W.
E. b S.	3 "	E.N.E.	W.S.W.	2 "	West
S.E. b E.	3 "	East	S.W.	1 "	S.W. b W.
S.E. b S.	1 "	S.E.	S.S.W.	1 "	S.W. b S.
S.S.E.	2 "	S.E.	South	2 "	S.S.W.

CORRECTING STANDARD COMPASS COURSES OR BEARINGS FOR LOCAL DEVIATION.

The variation of the compass, before noticed, is caused by the magnetism of the *Earth* acting upon the needle, and deflecting it from the *true* meridian of the world; but when a compass is used on board a ship, it is subject also to another disturbance, generally known to seamen as *local deviation*.

This additional source of error in the *direction* of the needle arises from the magnetism of the various masses of iron surrounding it, such as armour-plating, guns, tanks, engines, &c.

In every ship the compass has a *local deviation*, but in some ships it is much greater than in others.

The amount of *local deviation* is contained in the angle between the *correct magnetic* direction of the ship's head and the direction of the ship's head as shown by the *affected* compass.

The deviation is named *easterly*, when the north end of the needle is drawn to the *right* of the magnetic meridian by the iron of the ship.

The deviation is named *westerly*, when the north end of the needle is drawn to the *left* of the magnetic meridian by the iron of the ship.

Since the deflection of the compass needle from the magnetic meridian (*local deviation*) is similar in effect to that of the deflection of the magnetic from the true meridian (*variation*), the local deviation must be applied to the ship's courses, or compass bearings, in the same manner as the variation, viz. :—

- (a) When the deviation is *easterly*, it must be allowed to the *right* of the compass course to obtain the *correct magnetic course*.
- (b) When the deviation is *westerly*, it must be allowed to the *left* of the compass course to obtain the *correct magnetic course*.

Note.—In iron ships an error is caused in the deviation when the ship “heels” to starboard or port. In some ships the error, when the ship's head is N. or S., is equal to, or even greater than the amount of heel. In other words, a “heel” of 10° may give rise to a deviation of 10° or more.

LEEWAY.

WHEN a ship is under sail, and the wind is blowing on her side, she is liable to be pressed by its force to *leeward* of the course she is endeavouring to steer. This lateral or side-way pressure is the cause of the *Leeway*, and its amount can be ascertained by observing the angle made by the direction of her keel and the line of wake she leaves on the surface of the water.

Since the effect of the wind is to drive the ship to the *left* of her intended course when she is on the *Starboard* tack; and to the *right* of her intended course when she is on the *Port* tack, we have the following rules for allowing the *Leeway* :—

(a) When the ship is on the Starboard tack, the leeway must be allowed to the *left* of the course steered.

(b) When the ship is on the Port tack, the leeway must be allowed to the *right* of the course steered.

CORRECTING COURSES FOR LEEWAY.

EXAMPLES.

Courses.	Winds.	<i>Leeway.</i>	Courses made good.	Courses.	Winds.	<i>Leeway.</i>	Courses made good.
	STAR. TACK.				PORT TACK.		
N.N.E.	East	2	North	W.S.W.	South	1	W. b S.
N.W. b N.	N.E. b N.	1	N.W.	N.W. b W.	S.W. b W.	3	N.N.W.
E.N.E.	S.E.	3	N.E. b N.	N.E.	N.N.W.	2	E.N.E.
S.W. b W.	N.W. b W.	2	S.W. b S.	E.N.E.	N.N.E.	1	E. b N.
S.S.E.	S.W.	3	S.E. b E.	E.S.E.	N.E.	2	S.E.
S. b W.	W. b S.	2	S. b E.	S.S.E.	East	3	S. b W.
S.E. b E.	S. b W.	1	E.S.E.	South	E.S.E.	3	S.W. b S.
East	S.S.E.	3	N.E. b E.	S.S.W.	S.E.	2	S.W.
N.E. b E.	S.E. b E.	2	N.E. b N.	W.S.W.	South	1	W. b S.
N.E.	E.S.E.	1	N.E. b N.	W. b N.	S.W. b S.	2	N.W. b W.

MIDDLE LATITUDE SAILING.

THIS sailing is founded on a supposition that the meridian distance *half-way* between the latitude *sailed from*, and that *bound to*, is equal to the departure which a ship makes in sailing from the one latitude to the other; and is a compound of *Plane* and *Parallel* sailing.

The two cases most commonly used in this sailing are—

- | | |
|--|---|
| (1) When the <i>latitudes</i> and <i>longitudes</i> of the two places are given | } To find the course and distance between them. |
| (2) When the <i>course</i> between the two places and their <i>latitudes</i> are given | } To find the difference of longitude. |

CASE (I).

Find the true course and distance from one place to another.

(a) On the left hand of the page, write down the latitude of the place from, underneath which put the latitude of the place bound to, or worked for.

(b) In the middle of the page, write down the latitude of the place from, underneath which put the latitude of the place bound to, or worked for.

(c) On the right hand of the page, write down the longitude of the place from, underneath which put down the longitude of the place bound to, or worked for.

(A) When the latitudes are both of the *same* name, take the less from the greater; but when they are of contrary names, take their *sum*; the result, in either case, multiplied by 60 will give the difference of latitude, *in miles*.

(B) Find the middle latitude as follows :—

(1) When the latitudes are both of the *same* name, half their SUM will give the *middle* latitude.

(2) When the latitudes are of *contrary* names, half the GREATER latitude will give the middle latitude.

(C) When the longitudes of the two places are both of the same name, take the less from the greater ; but when they are of contrary names, take their sum ; the result, in either case, multiplied by 60, will give the difference of longitude, in *miles*.

When the difference of longitude *exceeds* 180°, take it from 360°, the remainder, multiplied by 60, will be the diff. long. *in miles*.

EXAMPLE TO CASE (1).

Find the True Course and Distance from Brest to Halifax.

Working of Example.

(a)	(b)	(c)
Lat. from (Brest) 48° 23' N.	Lat. Brest 48° 23' N.	Long. from (Brest) 4° 29' W.
Lat. to (Halifax) 44 44 N.	Lat. Halifax 44 44 N.	Long. to (Halifax) 63 36 W.
$\begin{array}{r} 3\ 39 \\ \hline 60 \end{array}$	$\begin{array}{r} 2) 93\ 7 \\ \hline \end{array}$	$\begin{array}{r} 59\ 7 \\ \hline 60 \end{array}$
diff. lat. 219 <i>m</i> .	mid. lat. 46° 33'	diff. long. 3547 <i>m</i> .

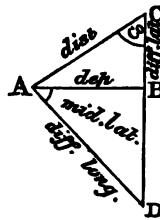
The figure used in this sailing is made up of *two* right-angled triangles, having one side (*dep.*) common to both.

One triangle is the figure used for *parallel* sailing, and the other is the figure used for *plane* sailing.

In the present case we have,

1st. In the parallel sailing figure, the *mid. lat.* and *diff. long.* are given, from which to find the *dep.*

2nd. In the plane sailing figure, the *diff. lat.* and the *dep.* (just found) are given, to determine the course and distance.



Note.—This figure is drawn out of proportion, in order to show the names of its parts with distinctness.

(1st) in $\triangle ABD$.

By Parallel Sailing.

$$\text{Given } \left\{ \begin{array}{l} \angle BAD = 46^\circ 33' \\ \text{the mid. lat.}^* \\ AD = 3547 \\ \text{the diff. long.} \end{array} \right\} \text{ to find } \left\{ \begin{array}{l} AB \\ \text{the} \\ \text{dep.} \end{array} \right.$$

$$\frac{AB}{AD} \text{ or } \frac{\text{dep.}}{\text{diff. long.}} = \cos. \text{ mid. lat.}^*$$

$$\therefore \text{dep.} = \text{diff. long.} \cos. \text{ mid. lat. (1)}$$

$$\text{log. diff. long. } 3547 \quad 3.549861$$

$$\text{log. cos. } 46^\circ 33' \quad 9.837412$$

$$\text{log. dep.} \quad \underline{3.387273}$$

$$\therefore \text{dep.} = 2439 \text{ miles.}$$

* Here the *middle* latitude between Brest and Halifax is taken as representing the *parallel half-way* between the two latitudes.

By Inspection.

The *dep.* is found from the Traverse Table, by the same method as in Ex. (A) of Parallel Sailing.

(2nd) In $\triangle ABD$.

By Plane Sailing.

$$\text{Given } \left\{ \begin{array}{l} AB = 2439 \text{ dep.} \\ BC = 219 \text{ diff. lat.} \end{array} \right\} \text{ to find } \left\{ \begin{array}{l} \text{co.} \\ \text{and} \\ \text{dist.} \end{array} \right.$$

$$\frac{AB}{BC} \text{ or } \frac{\text{dep.}}{\text{diff. lat.}} = \tan. \text{ co.} \quad (2)$$

$$\text{log. dep. } 2439 \quad 13.387273 (+10)$$

$$\text{log. diff. lat. } 219 \quad 2.340444$$

$$\text{log. tan. co.} \quad \underline{11.046829}$$

$$\therefore \text{Course} = 84^\circ 52' 15'' \\ \text{or S. } 84^\circ 52' 15'' \text{ W.}$$

$$\frac{AC}{BC} \text{ or } \frac{\text{dist.}}{\text{diff. lat.}} = \sec. \text{ co.}$$

$$\therefore \text{dist.} = \text{diff. lat.} \sec. \text{ co.} \quad (3)$$

$$\text{log. diff. lat. } 219 \quad 2.340444$$

$$\text{log. sec. co. } 84^\circ 52' 15'' \quad \underline{11.048656}$$

$$\text{log. dist.} \quad 3.389100$$

$$\therefore \text{dist.} = 2450 \text{ miles.}$$

By Inspection.

The *dist.* is found from Traverse Table, by the same method as in Ex. 2, of Plane Sailing.

Note.—The course can be found without first obtaining the *departure*, by the formula,

$$\tan. \text{ co.} = \frac{\text{diff. long.} \cos. \text{ mid. lat.}}{\text{diff. lat.}}$$

CASE (2).

To find the difference of longitude.

(a) Write down the lat. of the place from, on the left-hand side of the page, underneath which put the lat. arrived at, and find the difference of latitude.

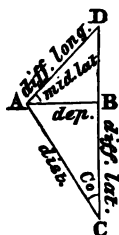
(b) Write down the lat. of the place from, on the right-hand side of the page, underneath which put the lat. arrived at, and find the middle latitude.

EXAMPLE TO CASE (2).

A ship sails N.W. by N., from lat. $44^{\circ} 15' N.$, until she arrives in lat. $46^{\circ} 15' N.$; find the difference of longitude she has made.

Working of Example.

(a)	(b)
Lat. from ... $44^{\circ} 15' N.$	Lat. from ... $44^{\circ} 15' N.$
Lat. in $46^{\circ} 15' N.$	Lat. in $46^{\circ} 15' N.$
<hr style="width: 50%; margin: 0;"/>	<hr style="width: 50%; margin: 0;"/>
$2^{\circ} 0'$	$2)90^{\circ} 30'$
60	<hr style="width: 50%; margin: 0;"/>
<hr style="width: 50%; margin: 0;"/>	mid. lat. $45^{\circ} 15'$
diff. lat. 120 miles.	



(1st) In $\triangle ABC$.
By Plane Sailing.

Given	{	co. N. $33^{\circ} 45' W.$ diff. lat. = 120m.	}	to find	}	dep.
<hr style="width: 80%; margin: 0;"/>						
		$\frac{AB}{BC}$ or $\frac{dep.}{diff. lat.}$			= tan. co.	
$\therefore dep. = diff. lat. \tan. co. \quad (1)$						
		log. diff. lat. 120m.	2.079181			
		log. tan. co. $33^{\circ} 45'$	9.824893			
		<hr style="width: 50%; margin: 0;"/>				
		log. dep.	1.904074			
		$\therefore dep. = 80.18m.$				

By Inspection.
Enter the Traverse Table with the given course, and given diff. lat. in their proper columns, and in the dep. column will be found the dep. required.

(2nd) In $\triangle ABD$.
(See Ex. (B.) Parallel Sailing.)

$\frac{AD}{AB}$ or $\frac{diff. long.}{dep.}$		= sec. mid. lat.	
$\therefore diff. long. = dep. \sec. mid. lat. \quad (2)$			
		log. dep.	1.904074
		log. sec. m. lat. $45^{\circ} 15'$	10.152418
		<hr style="width: 50%; margin: 0;"/>	
		log. diff. long.	2.056492
		$\therefore diff. long. = 113.9m. \text{ westing.}$	
		$= 1^{\circ} 53' 54'' W.$	

By Inspection.
The diff. long. is found from the Traverse Table by the method given in Ex. (B), Parallel Sailing.

Note.—The diff. long. can be found without first obtaining the *departure*, by the formula,

$$diff. long. = diff. lat. \tan. co. \sec. mid. lat.$$

MIDDLE LATITUDE SAILING.

EXAMPLES.

1. Find the *true* course and distance from Espichel Cape (Portugal), in lat. $38^{\circ} 25' N.$, and long. $9^{\circ} 13' W.$, to St. Mary's (Azores), in lat. $36^{\circ} 57' N.$, and long. $25^{\circ} 9' W.$

Ans. True course S. $83^{\circ} 20' 45'' W.$ Dist. 762m.

2. Find the *true* course and distance from Esquimaux Isles (Labrador), in lat. $58^{\circ} 35' N.$, and long. $56^{\circ} 21' W.$, to Bergen (Norway), in lat. $60^{\circ} 24' N.$, and long. $5^{\circ} 18' E.$

Ans. True course N. $80^{\circ} 2' 30'' E.$ Dist. 2018m.

3. Find the *true* course and distance from Goree (W.C. Africa), in lat. $14^{\circ} 40' N.$, and long. $17^{\circ} 24' W.$, to St. Salvador (Brazil), in lat. $13^{\circ} 5' S.$, and long. $38^{\circ} 28' W.$

Ans. True course S. $36^{\circ} 58' 0'' W.$ Dist. 2084m.

4. Find the standard *compass* course and distance from Guadaloupe I., in lat. $16^{\circ} 0' N.$, and long. $61^{\circ} 45' W.$, to Sal. I., in lat. $16^{\circ} 51' N.$, and long. $22^{\circ} 55' W.$ Var. $10^{\circ} W.$ Dev. $9^{\circ} E.$

Ans. True course N. $88^{\circ} 42' 15'' E.$ Comp. course N. $89^{\circ} 42' 15'' E.$ Dist. 2240m.

5. Find the standard *compass* course and distance from Gough's I., in lat. $40^{\circ} 19' S.$, and long. $9^{\circ} 44' W.$, to Grim Cape, in lat. $40^{\circ} 40' S.$, and long. $144^{\circ} 42' E.$ Var. $4^{\circ} W.$ Dev. $12^{\circ} E.$

Ans. True course S. $89^{\circ} 49' 45'' E.$ Comp. course N. $82^{\circ} 10' 15'' E.$ Dist. 7047m.

6. Find the standard *compass* course and distance from Woahoo, in lat. $21^{\circ} 18' N.$, and long. $157^{\circ} 55' W.$, to San Francisco, in lat. $37^{\circ} 48' N.$, and long. $122^{\circ} 24' W.$ Var. $10^{\circ} E.$ Dev. $10^{\circ} E.$

Ans. True course N. $61^{\circ} 53' E.$ Comp. course N. $41^{\circ} 53' E.$ Dist. 2101m.

7. A ship in lat. $9^{\circ} 15' N.$ sails E.N.E. until she is in lat. $11^{\circ} 25' N.$; find her difference of longitude.

Ans. Diff. long. $5^{\circ} 19' E.$

8. A ship in lat. $10^{\circ} 17' S.$ sails E.N.E. until she is in lat. $8^{\circ} 55' S.$; find her difference of longitude.

Ans. Diff. long. $3^{\circ} 21' E.$

9. A ship from Cape Frio (Brazil), in lat. $23^{\circ} 1' S.$, and long. $41^{\circ} 58' W.$, sails S.E. by E. until she is in lat. $25^{\circ} 25' S.$; find her difference of longitude, and long. arrived at.

Ans. Diff. long. $3^{\circ} 56' E.$ Long. in $38^{\circ} 2' W.$

MERCATOR'S SAILING.

THIS Sailing is derived from Mercator's projection of the Globe upon a plane, or *flat* surface, in which the degrees of longitude are *everywhere equal*; the degrees of latitude *increase* from the Equator towards the Poles; and the meridians, parallels, and rhumb lines (or courses made by the ship) are *all* represented by straight lines.

The two cases most commonly used in this sailing are,

- | | |
|---|--------------------------|
| (1) When the <i>latitudes</i> and <i>longi-</i> | } To find the course and |
| <i>tudes</i> of two places are given | |
| (2) When the <i>course</i> between two | } To find the difference |
| places and their <i>latitudes</i> are given | |

CASE (I).

Find the *True Course* and *Distance* from one place to another.

(a) On the left hand of the page, write down the latitude of the place *from*, underneath which put the latitude of the place *bound to* (or worked for).

(b) Write down the corresponding meridional parts of the two latitudes in the middle of the page.

(c) On the right hand of the page, write down the longitude of the place *from*, underneath which put the longitude of the place *bound to* (or worked for).

(A) When the latitudes are both of the *same* name, take the less from the greater; but when they are of *contrary* names, take their sum; the result, in either case, multiplied by 60, will give the *true diff. lat. in miles*.

(B) Having taken out the meridional parts of the two latitudes from Table y, Inman, take their *difference* when the latitudes are of the *same* name; but their *sum* when the latitudes are of *contrary* names. The result, in either case, will be the *meridional* diff. lat. (mer. diff. lat.) in *miles*.

(C) When the longitudes of the two places are both of the *same* name, take the less from the greater; but when they are of *contrary* names, take their sum; the result, in either case, multiplied by 60, will give the diff. long. in *miles*.

Note.—When the diff. long. *exceeds* 180 degrees, take it from 360 degrees, the remainder must be used as the diff. long. when turned into *miles*.

The difference of latitude (in geographical miles) between two places on the Globe, is called the *true* difference of latitude.

The difference of latitude (in geographical miles) between two places on the Chart, is called the *meridional* difference of latitude.

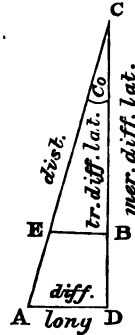
EXAMPLE TO CASE (1).

Find the True Course and Distance from Cape Clear to Porto Santo.

Working of Example.

(a)	(b)	(c)
Lat. C. Clear (from) 51° 25' N.	Mer. par. 3609	Long. C. Clear 9° 29' W.
Lat. P. Santo (to) 33 3 N.	Mer. par. 2103	Long. P. Santo 16 17 W.
18 22	1506	6 48 W.
60	mer. diff. lat.	60
True diff. lat. 1102 miles.		diff. long. 408 miles

Note. — 1st. Work with the triangle CDA, which is formed by side CD, the *mer. diff. lat.*, and side DA the *diff. long.*, finding $\angle C$ (the course), which is *common* to both triangles. 2nd. Having found the common $\angle C$, work with triangle EBC, which is formed by BC, the *true diff. lat.*, and CE the *distance* required.



[1st.]
 In $\triangle CDA$,
 Given $\left\{ \begin{array}{l} \text{mer. diff. lat.} \\ 1506\text{m.} \\ \text{diff. long. } 408\text{m.} \end{array} \right\}$ to find $\left\{ \begin{array}{l} \angle C, \text{ the} \\ \text{course} \end{array} \right\}$
 $\frac{DA}{DC}$ or $\frac{\text{diff. long.}}{\text{mer. diff. lat.}} = \tan. \text{co. } (1)$
 $\log. \text{diff. long. } 408 \quad 12.610660 + 10$
 $\log. \text{m. diff. lat. } 1506 \quad 3.177825$
 $\log. \tan. \text{co.} \quad 9.432835$
 $\therefore \text{course} = 15^\circ 9' 30''$
 or, true course = S. $15^\circ 9' 30''$ W.

[2nd.]
 In $\triangle CBE$,
 Given $\left\{ \begin{array}{l} \text{true diff. lat. } 1102\text{ m.} \\ \text{true co. } 15^\circ 9' 30'' \end{array} \right\}$ to find $\left\{ \begin{array}{l} \text{distance} \end{array} \right\}$
 $\frac{CE}{CB}$ or $\frac{\text{dist.}}{\text{tr. diff. lat.}} = \sec. \text{course } (2)$
 $\therefore \text{dist.} = \text{tr. diff. lat. sec. co. } (2)$
 $\log. \text{tr. diff. lat. } 1102 \quad 3.042182$
 $\log. \sec. \text{co. } 15^\circ 9' 30'' \quad 10.015380$
 $\log. \text{dist.} \quad 3.057562$
 $\therefore \text{distance} = 1141 \text{ miles.}$

Note.—The course must be marked S.....W., because Porto Santo (the place worked for) is to the *southward* and *westward* of Cape Clear—the place *from*.

By Inspection.

With the *mer. diff. lat.* as *diff. lat.*, and *diff. long.* as *dep.*, enter the Traverse Table, and find the corresponding *course*. This will be the course required.

Again, with this *course*, and the *true diff. lat.* as *diff. lat.*, enter the Traverse Table, and find the corresponding *distance*. This will be the distance required.

Example.—Enter the Traverse Table with *mer. diff. lat.* 150, 6 (or the nearest number to it) as *diff. lat.*, and the *diff. long.* 40, 8 (or the nearest number to it) as *dep.*, the corresponding *course* is 15° . This will be course required.

Again, enter the Traverse Table with this *course* 15° , and the *true diff. lat.* 110, 2 (or the nearest number to it) as *diff. lat.*, and find the corresponding *dist.* 114 (that is, 1140). This will be the distance required. (*See Note.*)

Note.—In the above example, since the numbers which represent the *mer. diff. lat.*, *diff. long.*, and *true diff. lat.* respectively, are greater than can be found in the Traverse Table, we cut off one figure to the *right* (that is, divide by 10) and seek for these new numbers in the Table; the division thus made, does not effect the *course*, but when finding the *distance* we multiply what the Table gives by 10, which will be the distance required.

CASE (2).

To find the Difference of Longitude.

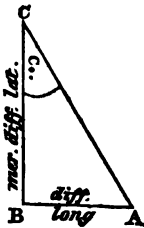
Write down the two latitudes, and take out their mer. parts; and find the *mer. diff. lat.* by the same rule as in Example (1).

EXAMPLE TO CASE (2).

A ship from latitude $50^{\circ} 10' N.$, sails S.E. by S. until she is in lat. $48^{\circ} 10' N.$; find the difference of longitude she has made.

Working of Example.

Lat. from $50^{\circ} 10' N.$	Mer. part. 3490
Lat. in $48^{\circ} 10' N.$	Mer. part. 3306
	Mer. diff. lat. 184 miles.



Given { mer. diff. lat. 184m. } to find { diff. long. }
 { course S. $33^{\circ} 45' E.$ }

$$\frac{BA}{BC} \text{ or } \frac{\text{diff. long.}}{\text{mer. diff. lat.}} = \tan. \text{ co.}$$

$$\therefore \text{diff. long.} = \text{mer. diff. lat.} \tan. \text{ co.}$$

log. mer. diff. lat. 184m.	2.264818
log. tan. co. $33^{\circ} 45'$	9.824893
	2.089711
log. diff. long.	2.089711
\therefore diff. long. = 123 miles, Easting.	

$$60 \) \ 123'$$

$$\text{or, diff. long. } \underline{2^{\circ} 3'} \text{ East.}$$

By Inspection.

With the given *course*, and the mer. diff. lat., as *diff. lat.*, enter the Traverse Table, and the corresponding *dep.* will be the diff. long. required.

Example.—Enter the Traverse Table with course 3 points and the mer. diff. lat. 184 (or the nearest number to it) as *diff. lat.*, and the corresponding *dep.* 123.3 will be the diff. long. required (nearly).

**To turn a TRUE Course into a COMPASS Course
(or, Course to Steer).**

1st. Correct for VARIATION.

Apply the variation to the true course as follows :

Easterly variation to the left.

Westerly variation to the right.

This will give the correct magnetic course.

2nd. Correct for DEVIATION.

Apply the deviation to the correct magnetic course as follows :

Easterly deviation to the left.

Westerly deviation to the right.

The result will give the standard compass course
(or, course to steer).

EXAMPLES.

1. The true course from Corunna to Cape Clear is N. 6° W.; what is the standard compass course, or course to steer, when the variation is 20° W., and the deviation 3° E. ?

True course	N. 6° W.	
Variation, 20° W.	20	allow to <i>right</i> of true course.
	14	
Corr. mag. course	N. 14 E.	
Deviation, 3° E.	3	allow to <i>left</i> of corr. mag. course.
	11	
Stand. comp. course	N. 11 E.	or, course to steer from Corunna to C. Clear.

2. The true course from Mobile to Bermuda is N. 84° E.; what is the standard compass course, or course to steer, when the variation is 5° W., and deviation 10° E. ?

True course	N. 84° E.	
Variation, 5° W.	5	allow to <i>right</i> of true course.
	89	
Corr. mag. course	N. 89 E.	
Deviation, 10° E.	10	allow to <i>left</i> of corr. mag. course.
	79	
Stand. comp. course	N. 79 E.	or, course to steer from Mobile to Bermuda.

MERCATOR'S SAILING.

EXAMPLES FOR EXERCISE.

1. Find the *true* course and distance from Mobile, in lat. $30^{\circ} 14' N.$ and long. $88^{\circ} 1' W.$, to Bermuda, in lat. $32^{\circ} 19' N.$ and long. $64^{\circ} 52' W.$ —See Example, p. 35.

Ans. Course N. $83^{\circ} 59' 0'' E.$ Dist. 1197 miles.

2. Find the *true* course and distance from Bon Cape (Barbary), in lat. $37^{\circ} 5' N.$ and long. $11^{\circ} 3' E.$, to Caprera I. (Mediterranean), in lat. $41^{\circ} 13' N.$ and long. $9^{\circ} 29' E.$

Ans. Course N. $16^{\circ} 27' 0'' W.$ Dist. 259 miles.

3. Find the *true* course and distance from Broad Haven (Ireland), in lat. $54^{\circ} 26' N.$ and long. $10^{\circ} 12' W.$, to Charlestown (United States), in lat. $32^{\circ} 46' N.$ and long. $79^{\circ} 55' W.$

Ans. Course S. $66^{\circ} 25' W.$ Dist. 3248 miles.

4. Find the *true* course and distance from Columbia R. (Ent.) in lat. $46^{\circ} 19' N.$ and long. $123^{\circ} 44' W.$, to Combay I. (Indian Ocean), in lat. $7^{\circ} 49' S.$ and long. $123^{\circ} 51' E.$

Note.—In crossing the meridian of 180° from W. long. to E. long. the course will be westerly.

Ans. Course S. $61^{\circ} 49' 30'' W.$ Dist. 6878 miles.

5. Find the standard *compass* course and distance from Congo Riv., in lat. $6^{\circ} 5' S.$ and long. $12^{\circ} 15' E.$, to St. Salvador, in lat. $13^{\circ} 5' S.$ and long. $38^{\circ} 28' W.$ Var. $5^{\circ} E.$ Dev. $7^{\circ} W.$

Ans. True course S. $82^{\circ} 1' W.$ Comp. course S. $84^{\circ} 1' W.$ Dist. 3030 miles.

6. Find the standard *compass* course and distance from Jervis' Island (Torres' Strait), in lat. $9^{\circ} 55' S.$ and long. $142^{\circ} 10' E.$, to Valparaiso (Chile), in lat. $33^{\circ} 2' S.$ and long. $71^{\circ} 41' W.$ Var. $15^{\circ} E.$ Dev. $9^{\circ} E.$

Note.—In crossing the meridian of 180° from E. long. to W. long. the course will be easterly.

Ans. True course S. $80^{\circ} 16' E.$ Comp. course N. $75^{\circ} 44' E.$ Dist. 8204 miles.

7. A ship from C. Clear, in lat. $51^{\circ} 26' N.$ and long. $9^{\circ} 22' W.$, sails S.S.W. until she is in lat. $40^{\circ} 0' N.$; find her difference of longitude.—See Example, p. 36.

Ans. Diff. long. $6^{\circ} 49' W.$

8. A ship from L'Orient (France), in lat. $47^{\circ} 43' N.$ and long. $3^{\circ} 21' W.$, sails N.W. by W. until she is in lat. $49^{\circ} 15' N.$; find her difference of longitude.

Ans. Diff. long. $3^{\circ} 23' W.$

MISCELLANEOUS EXAMPLES.

1. A point of land in lat. $0^{\circ} 9' 40''$ N. and long. $50^{\circ} 10' W.$ bore from a ship N. $47^{\circ} W.$, distant 22 miles; find the lat. and long. of the ship.

Ans. Lat. $0^{\circ} 5' 20'' S.$ Long. $49^{\circ} 54' W.$

2. A headland in lat. $0^{\circ} 2' 20'' S.$ and long. $98^{\circ} 15' 30'' E.$ bore from a ship (by standard compass) S.W. by S., distant 21 miles; find the lat. and long. of the ship. Var. $2^{\circ} E.$ Dev. $2^{\circ} W.$

Ans. Lat. $0^{\circ} 15' 10'' N.$ Long. $98^{\circ} 27' 5'' E.$

3. The centre of an island in lat. $0^{\circ} 10' S.$ and long. $66^{\circ} 15' W.$ is observed to bear from a ship S. $20^{\circ} W.$ (by standard compass), distant 18 miles; find her lat. and long. Var. $3^{\circ} E.$ Dev. $1^{\circ} W.$

Ans. Lat. $6' 42'' N.$ Long. $66^{\circ} 8' W.$

4. From a ship in lat. $50^{\circ} 5' N.$ and long. $2^{\circ} 7' W.$ a point of land was observed to bear N.N.W. $\frac{1}{4}W.$, and after sailing W.S.W. 10 miles, it bore N. $\frac{1}{2}W.$; required the lat. and long. of the point.

Ans. Lat. $50^{\circ} 31' N.$ Long. $2^{\circ} 36' W.$

5. A point of land in lat. $0^{\circ} 10' N.$ and long. $6^{\circ} 44' E.$ is observed to bear north (by compass) 30 miles from a ship bound to St. Helena, in lat. $15^{\circ} 55' S.$ and long. $5^{\circ} 44' W.$ How far is she from her destination, and what course must she steer to reach it? Var. $2^{\circ} W.$

Ans. Course S. $39^{\circ} 11' W.$ Dist. 1173 miles.

6. A rock, not marked on the chart, is discovered bearing N. $20^{\circ} W.$ (true) from a ship, and distant 2 miles. Being unable to obtain observations at the time, she sails N. $40^{\circ} E.$ 60 miles, and then finds by sights that she is in lat. $10^{\circ} 30' S.$ and long. $108^{\circ} 20' W.$; required the lat. and long. of the rock.

Ans. Lat. $11^{\circ} 14' S.$ Long. $109^{\circ} 0' W.$

7. A ship in lat. $0^{\circ} 7' 30'' S.$ and long. $149^{\circ} 15' 45'' E.$ sails N. $17^{\circ} E.$ 12 miles, when she grounds on a shoal. From her masthead the sea is observed to break over a rock, bearing

S. 20° E., distant 9 miles ; required the lat. and long. of the shoal on which she struck, and also of the rock.

Ans. Lat. of shoal $0^{\circ} 4' N.$ Long. $149^{\circ} 19' 15'' E.$ Lat. of rock $0^{\circ} 4' 30'' S.$ Long. $149^{\circ} 22' 20'' E.$

8. A ship sailing N.N.E. at the rate of 6 knots, passes a coral reef, the southern extremity of which bears N. 70° W., distant 3 miles. At the end of an hour it is found by observation that the ship is in lat. $11^{\circ} 30' S.$ and long. $134^{\circ} 25' W.$, when the northern extremity bears west, distant 4 miles ; required the lat. and long. of the southern extremity of the reef, its length, and the direction in which it lies.

Ans. Lat. of southern extremity $11^{\circ} 34' 30'' S.$ Long. $134^{\circ} 30' 15'' W.$
Length of reef 4.654 miles. Direction in which it lies, N. $13^{\circ} 52' E.$

9. Wanting to ascertain the breadth of a river, in the entrance to which are several islands ; from a cliff on the eastern side I sailed N.N.W. 50 miles, and then running S.S.W. I reached the western shore : required the breadth between these two positions.

Ans. 35.35 miles.

10. Two ships take their departure from the Lizard, in lat. $49^{\circ} 58' N.$, one bound to St. Michael's, which lies 715m. to the south and 745m. to the west ; the other bound to Lisbon, which lies 661m. to the south and 215m. to the west, reckoning from the Lizard. They sail in company S.W. $\frac{1}{2}$ W. 610 miles, and then part. What is the direct course and distance of each ship to her port ?

Ans. First ship, S. $43^{\circ} 7' E.$, 375.3m.
Second do., S. $39^{\circ} 49' W.$, 427.1m.

11. A ship, after sailing S.E. by S., 62 leagues from her port, in lat. $29^{\circ} 16' N.$, meets another ship that had run 124 leagues from a place bearing W. by N. of the port the first ship departed from ; what was the ship's course, and the distance of the two ports ?

Ans. Ship's course, S. $58^{\circ} 3' E.$ Dist. 77.18 leagues.

THE DEAD RECKONING.

IN the Navy, the time in the ship's log-book is reckoned from *midnight*, as civil or common time ; the first hour is therefore 1 o'clock in the morning, and the hours are carried on to 12, or *noon*, and then to 12, or *midnight*.

The log, or account of the ship's proceedings, begins and ends at *noon*.

The place of the ship is ascertained every day at NOON, from the *log-account* ; besides this, the course and distance made good during the last 24 hours, and the bearing and distance of the port bound to, or an intermediate point of land, with reference to which the course has to be *shaped*, so as to make it, or to avoid it, are to be found.

The work of computation necessary to obtain the above is called a *Day's Work*.

To work a Day's Work.

When a departure from a known point of land has been taken, consider it as a *course* in the *opposite* direction to the bearing.

1. Correct the departure course for deviation and variation, and place it in the Traverse Table, with its distance.

2. Then correct *each* course on the log-board for deviation and variation, also for leeway (if any), and place it with its corresponding distance run, in the Traverse Table.

3. When the ship is known to have been "set" in any given direction by a current (or other cause), consider this "set" as a course, and correct it for variation *only*, and place it with its corresponding distance in the Traverse Table.

4. If the ship is lying-to, take the *middle* point between the direction of her head when she comes up to the wind and when she falls off from it, for the *course*. Allow to this course the *mean* of the deviations upon the two directions of her head, as the deviation ; then correct for variation and leeway, as usual.

5. Having thus corrected all the courses, take out the diff. lat. and dep. for each from the Traverse Table (x. Inman).

6. Then add up each column separately, and take the less diff. lat. from the greater, marking the remainder with the name of the greater, which will give the diff. lat. made good.

7. Also, take the less dep. from the greater, marking the remainder with the name of the greater, which will give the dep. made good.

To find the Latitude in, by Dead Reckoning (D.R.).

(a) To the latitude by *observation* of the preceding noon, apply the difference of latitude made good, which will give the latitude in by D.R. as follows:—

(1) When the lat. left and diff. lat. are of the *same* name, their *sum* will be the lat. in.*

(2) When the lat. left and diff. lat. are of *contrary* names, take their *difference*, and the remainder will be the lat. in of the name of the greater.*

(b) When the latitude was not observed at *noon*, but at some other time, and from it the lat. D.R. has been deduced, the lat. thus obtained should be noted as “brought up,” because the lat. by D.R. when employed for comparison with the lat. by *observation* is considered as referred to the beginning of the day, unless the contrary is expressed.

(c) When there was no lat. by *observation* at the preceding noon, the lat. D.R. at that time must be used.

To find the Longitude in by Dead Reckoning (D.R.).

To the longitude D.R. of the preceding noon, apply the difference of longitude made good, which will give the long. in by D.R.

If the longitude at the preceding noon has been obtained from observations upon which dependence can be placed, it should be used instead of the long. D.R. to find the long. in.

(1) When the long. left and diff. long. are of the *same* name, their *sum* will be the long. in.*

* See Notes, pages 12 and 14.

(2) When the long. left and diff. long. are of *contrary* names, take their *difference*, which will be the long. in of the name of the greater.*

Note.—The diff. long. is obtained from the *dep.* made good, either by Ex. (B.) Parallel Sailing, or by Case (2) Middle Latitude Sailing, or by Case (2) Mercator's Sailing.

To find the True Course and Distance made good.

These are found by Mercator's Sailing. (*See* Rule, p. 33.)

To find the True Bearing and Distance.

These are found by Case 1, Middle Latitude Sailing; or by Case 1, Mercator's Sailing.

When the ship is on the same parallel of latitude as the place worked for, the true course will be due E. or due W.; and the distance must be found by Ex. A, Parallel Sailing.

To Shape the Course to Steer.

Having found the true bearing of the place bound to, the course to steer is obtained by applying the variation and deviation. (*See* Rules, p. 37.)

To find the Current (if any).

The course and distance between the place of the ship by D.R. and her place by obs. worked out as in the case of the true bearing and distance (above) will give the set and rate of the current. The *set* being marked as if going from D.R. to observation.

* See Notes, pages 12 and 14

June 14th, at noon, a point of land, in lat. $42^{\circ} 56' N.$, and long. $9^{\circ} 16' W.$, bore by standard compass E. by N., distant 18 miles (deviation $6^{\circ} W.$), afterwards sailed as by the following log account: required the latitude and longitude in at noon of June 15th.

Hours.	Knots.	Tenths.	Standard Compass Courses.	Lee-way Points.	Winds.	Deviation.	Remarks.
1	3	6	S.W. b W.	$\frac{1}{4}$	N.W. b W.	$8^{\circ} W.$	P.M.
2	3	5					
3	4	2					
4	4	4					
5	4	2					
6	4	0	W. b N.	$\frac{1}{4}$	N. b W.	$12^{\circ} W.$	Variation allowed, $2\frac{1}{4}$ points westerly.
7	4	5					
8	5	0					
9	5	2					
10	5	4					
11	6	5					
12	6	3					
1	6	2	E.N.E.	0	North	$13^{\circ} E.$	A.M.
2	5	5					
3	5	4					
4	5	0					
5	5	0					
6	4	6					
7	4	0	S.S.E.	$\frac{1}{4}$	East	$6^{\circ} E.$	A current set the ship the last 8 hours N. b W. (corr. mag.) $1\frac{1}{2}$ mile an hour.
8	3	3					
9	3	5					
10	3	0					
11	3	2					
12	3	0					

True Courses.—S. $47^{\circ} W.$ (dep. co.) $18'$. S. $17^{\circ} W.$ $19'9''$. S. $61^{\circ} W.$ $36'9''$.

N. $55^{\circ} E.$ $31'7''$. S. $39^{\circ} E.$ $20'$. N. $37^{\circ} W.$ (current co.) $12'$.

Diff. lat. $36'8'' S.$ Dep. $19'39'' W.$

Latitude in, $42^{\circ} 19' N.$ Longitude in, $9^{\circ} 43' W.$

August 20th, at noon, Ushant, in lat. 48° 28' N., long. 5° 3' W., bore S.E. by standard compass, distant 10 miles (deviation 11° W.), afterwards sailed as by the following log account: required the latitude and longitude of the ship on August 21st, at noon.

Hours.	Knots.	Tenths.	Standard Compass Courses.	Lee-way Points.	Winds.	Deviation.	Remarks.	
1	5	5	West	—	N.N.W.	14° W.	P.M.	
2	5	7						
3	6	0						
4	6	2						
5	7	0						
6	7	5						
7	8	3					Variation, 25° westerly	
8	7	2	South	—	W. b S.	3° E.		
9	7	0						
10	$\frac{3}{3}$	$\frac{2}{0}$	S.S.W.* N.N.W.	— $\frac{1}{2}$	West Ditto	4° W. 3° W.	9'30. Tacked.	
11	5	5						
12	4	2						
1	3	0					A.M.	
2	3	2						
3	—	}	Up S. b E.†	} 5	S.-westerly	{	4° E. 12° E.	2. Hove-to.
4	—		Off S.E. b E.					
5	—	}	Up W. b S.†	} 4	Ditto	{	12° W. 8° W.	5. Wore ship.
7	—		Off N.W. b W.					
8	—							
9	2	0	W.S.W.	2	South	10° W.	Allow 1 mile an hour for drift, whilst lying-to.	
10	3	0						
11	3	6					(See Art. 4, page 41).	
12	3	4						

True Courses.—N. 81° W. (dep. co.) 10'. S. 51° W. 46'2". S. 22° E. 14'2". S. 6° E. 3'2".
N. 45° W. 18'9". N. 73° E. 3'. N. 69° W. 3'. S. 55° W. 12'.

Diff. lat. 34'71" S. Dep. 62'22' W.

Latitude in, 47° 53' N. Longitude in, 6° 36' W.

* When the ship has altered course during the hour, a line is drawn between the two courses made in that interval.
† See (4), page 41.

Ship's Head by Standard Compass.	Deviation.	Ship's Head by Standard Compass.	Deviation.
NORTH.	0° E.	SOUTH.	0° E.
N. b E.	2° 0	S. b W.	2° 0 W.
N.N.E.	3° 0	S.S.W.	4° 0
N.E. b N.	5° 0	S.W. b S.	6° 0
N.E.	9° 0	S.W.	6° 0
N.E. b E.	11° 0	S.W. b W.	8° 0
E.N.E.	13° 0	W.S.W.	10° 0
E. b N.	14° 0	W. b S.	12° 0
EAST.	14° 0	WEST.	14° 0
E. b S.	15° 0	W. b N.	12° 0
E.S.E.	13° 0	W.N.W.	10° 0
S.E. b E.	11° 0	N.W. b W.	8° 0
S.E.	9° 0	N.W.	5° 0
S.E. b S.	7° 0	N.W. b N.	3° 0
S.S.E.	6° 0	N.N.W.	2° 0
S. b E.	4° 0 E.	N. b W.	1° 0 W.

Figures to denote the Force of the Wind.

- | | | |
|---------------------|---|---|
| 0. Calm. | | |
| 1. Light air, | } Or just sufficient to give steerage way.
Or that in which a ship, with all sail set, and clean full, would go in smooth water from | } 1 to 2 knots.
3 to 4 knots.
5 to 6 knots.
Royals, &c. |
| 2. Light breeze, | | |
| 3. Gentle breeze, | | |
| 4. Moderate breeze, | | |
| 5. Fresh breeze, | } Or, that to which she could just carry in chase, full and by | } Double-reefed topsails, jib, &c.
Treble-reefed topsails, &c.
Close-reefed topsails and courses. |
| 6. Strong breeze, | | |
| 7. Moderate gale, | | |
| 8. Fresh gale, | | |
| 9. Strong gale, | } Or that with which she could scarcely bear close-reefed main-topsail, and reefed foresail. | } Close-reefed topsails and courses. |
| 10. Whole gale, | | |
| 11. Storm, | Or that which would reduce her to storm staysails. | |
| 12. Hurricane, | Or that which no canvas could withstand. | |

Letters to denote the State of the Weather.

- | | |
|---|---|
| <i>b.</i> Blue sky, whether with clear or hazy atmosphere. | <i>q.</i> Squally. |
| <i>c.</i> Cloudy, but detached opening clouds. | <i>r.</i> Rain, continued rain. |
| <i>d.</i> Drizzling rain. | <i>s.</i> Snow. |
| <i>f.</i> Foggy. — <i>f.</i> Thick fog. | <i>t.</i> Thunder. |
| <i>g.</i> Gloomy, dark weather. | <i>u.</i> Ugly threatening appearance of the weather. |
| <i>h.</i> Hail. | <i>v.</i> Visibility of distant objects whether the sky is cloudy or not. |
| <i>l.</i> Lightning. | <i>w.</i> Wet, dew. |
| <i>m.</i> Misty, hazy atmosphere. | <i>•</i> under any letter indicates an extraordinary degree. |
| <i>o.</i> Overcast, the whole sky being covered with an impervious cloud. | |
| <i>p.</i> Passing temporary showers. | |

Thus by combining certain letters we can record the ordinary phenomena of the weather: for example, *bp* would denote blue sky, with passing temporary showers; *qrt* would denote squally weather, with rain and thunder; *gr* would denote gloomy, dark weather, with *very heavy* rain.

NAUTICAL ASTRONOMY.

ON TIME.

As the Earth revolves on its axis from west to east, the Sun *appears* to move in the contrary direction from east to west, making it 12 o'clock, or NOON, at each meridian it successively passes over in its daily course.

Thus, when the Sun is *on* the meridian of Greenwich it is 12 o'clock, *apparent time*, at that place, and the inhabitants of places to the *westward* of Greenwich, having to *wait for the Sun* to make it noon by their clocks, will find their time behind that of Greenwich; whilst those living to the *eastward* of Greenwich, having had the Sun on their meridians *before* the people at Greenwich, will find their time in advance of that of Greenwich.

Places 15° to the westward of Greenwich wait 1 hour, places 30° to the westward of Greenwich wait 2 hours, and so on, for the Sun to cross their meridians after it has passed that of Greenwich.

In a similar manner, places 15° to the eastward of Greenwich have the Sun on their meridian 1 hour before the people at Greenwich; places 30° to the eastward 2 hours, and so on; and their clocks are therefore in advance of the clocks at Greenwich, in proportion to their longitude in time.

Hence, all places to the westward of Greenwich are to be considered *slow* on Greenwich time, whilst all places to the eastward of Greenwich are to be considered *fast* on Greenwich time. Consequently the clocks to the westward will show *less* time, and the clocks to the eastward will show *more* time than the clocks at Greenwich, or than a chronometer set to *Greenwich time*.

EXAMPLE I.

Supposing it is required to find the time at Plymouth, in long. $4^{\circ} 24' W.$, at the instant the clocks at Greenwich are striking 12 at Noon.

Time at Greenwich	12h. om. os. (Noon)
Longitude of Plymouth in time (slow)	0 17 36
	<hr/>
Time at Plymouth	11 42 24 A.M.

Since the Sun does not come to the meridian of Plymouth until 17m. 36s. after it has passed the meridian of Greenwich.

EXAMPLE II.

Supposing it is required to find the time at Madras, in long. $80^{\circ} 14' 15'' E.$, when the clocks at Greenwich are striking 12 at Noon.

Time at Greenwich	oh. om. os. (Noon)
Longitude of Madras in time (fast)	5 20 57
	<hr/>
Time at Madras	5 20 57 P.M.

Since the Sun passes the meridian of Madras 5h. 20m. 57s. before it comes to the meridian of Greenwich.

It is often necessary to determine the Greenwich time, when the time at another place and the longitude are given.

In this case, by applying the longitude in *time* to the time at the place, the corresponding time at Greenwich can be obtained; remembering that places to the westward are behind, or *slow* of Greenwich, and places to the eastward are in advance, or *fast* of Greenwich, hence the following rules to find the Greenwich date.

THE GREENWICH DATE.

The term *Greenwich date* is here used, being preferable to Greenwich time, because in Nautical Astronomy it is essential to note the day as well as the hour.

The *Civil Day* is that in general use ; it begins at midnight and ends at the midnight following. It is divided into two parts of 12 hours each ; the first is marked A.M. (before noon), and the latter is marked P.M. (afternoon).

The *Astronomical Day* commences 12 hours after the Civil Day, that is, at NOON, and ends on the following Noon ; it is generally reckoned through the 24 hours, from Noon to Noon.

The Noon of the Civil Day, and the beginning of the Astronomical Day take place at the same moment.

The computations in the Nautical Almanac are made for *astronomical* time at the meridian of Greenwich.

TO FIND THE GREENWICH DATE.

RULES.

REDUCE the time at ship to *astronomical* time, by one of the following rules :—

(a) When the time at ship is A.M., add 12 to the *hours*, and put the date one day *back* ; the result will be the *astronomical time* at ship.

(b) When the time at ship is P.M., *omit* the letters P.M., and the ship time will be expressed *astronomically*.

LONGITUDE, **West**.—Turn the longitude into time (by the Table of Haversines), and *add* it to the astronomical time at ship ; the sum will be the *Greenwich date*, of the *same* day as the ship date.

Note 1.—If the sum be *greater* than 24 hours, reject 24h. ; the remainder will be the Greenwich date, of the day *after* the astronomical *ship* date.

LONGITUDE, **East.**—Turn the longitude into time ; *subtract* it from the astronomical time at ship, and the remainder will be the *Greenwich date*, of the *same* day as the astronomical ship date.

Note 2.—If the longitude in time be *greater* than the *astronomical* time, add 24 hours to the ship time, and put the day one back ; then *subtract* the longitude in time ; the remainder will be the *Greenwich date*, of the day *before* the *ship* date.

EXAMPLE I.

March 10th, at 2h. 50m. P.M. (mean time) at ship, in longitude 60° W. ; required the Greenwich date.

Time at ship	2h. 50m. March 10th (b)
Longitude 60°, in time	4 0 W. (add)
	<hr style="width: 50px; margin-left: auto; margin-right: 0;"/>
Greenwich date	6 50 March 10th.

EXAMPLE II.

September 14th, at 6h. 10m. A.M. apparent time at ship, in longitude 35° E. ; required the Greenwich date.

Civil time at ship	6h. 10m. A.M. Sept. 14th
Add	12 0
	<hr style="width: 50px; margin-left: auto; margin-right: 0;"/>
<i>Astronomical</i> time at ship	18 10 Sept. 13th (a)
Longitude 35°, in time	2 20 E. (subtract)
	<hr style="width: 50px; margin-left: auto; margin-right: 0;"/>
Greenwich date	15 50 Sept. 13th.

EXAMPLE to Note 1.

Find the Greenwich date corresponding to 9h. 35m. A.M. (civil time) at a ship, on May 1st, in longitude 65° W.

Time at ship	9h. 35m. A.M. May 1st	
Add	<u>12 0</u>	
<i>Astronomical</i> time at ship	21 35	April 30th (a)
Longitude 65° W., in time	<u>4 20</u>	(add)
Greenwich date	25 55	April 30th
Subtract	<u>24 0</u>	
Corresponding Greenwich date of the day <i>after</i> as- tronomical ship date (see <i>Note 1</i>)	1 55	May 1st

EXAMPLE to Note 2.

Find the Greenwich date corresponding to 5h. 25m. P.M. at a ship, on August 1st, in longitude 100° E.

Astronomical time at ship	5h. 25m. Aug. 1st (b)	
Add	<u>24 0</u>	
Time at ship	29 25	July 31st (<i>Note 2</i>)
Longitude 100° E., in time	<u>6 40</u>	(sub.)
Corresponding Greenwich date of day <i>before</i> the astronomical ship date	22 45	July 31st

EXAMPLES.

Find the Greenwich dates corresponding to the following ship dates and longitudes :—

1. Ship, Jan. 4th, at 3h. 30m. P.M., longitude 30° W. and longitude 30° E.
2. Ship, Aug. 9th, at 7h. 20m. P.M., longitude 110° W. and longitude 110° E.
3. Ship, Sept. 14th, at 11h. 30m. P.M., longitude $10^{\circ} 30'$ W. and longitude $10^{\circ} 30'$ E.
4. Ship, Oct. 21st, at 4h. 55m. P.M., longitude $35^{\circ} 10'$ W. and longitude $35^{\circ} 10'$ E.
5. Ship, Nov. 5th, at *midnight*, longitude $16^{\circ} 55'$ W. and longitude $16^{\circ} 55'$ E.
6. Ship, Dec. 2nd, at *midnight*, longitude 115° W. and longitude 115° E.
7. Ship, Feb. 14th, at 7h. 25m. A.M., longitude 100° W. and longitude 100° E.
8. Ship, Aug. 31st, at 9h. 50m. A.M., longitude $60^{\circ} 15'$ W. and longitude $60^{\circ} 15'$ E.

Note.—*Midnight* denotes 12 hours *after* noon, or 12h. P.M. of the given day.

Answers.

1. G.D. 5h. 30m. Jan. 4th, and 1h. 30m. Jan. 4th.
 2. G.D. 14h. 40m. Aug. 9th, and oh. om. (Noon) Aug. 9th.
 3. G.D. 12h. 12m. Sept. 14th, and 10h. 48m. Sept. 14th.
 4. G.D. 7h. 15m. 40s. Oct. 21st, and 2h. 34m. 20s. Oct. 21st.
 5. G.D. 13h. 7m. 40s. Nov. 5th, and 10h. 52m. 20s. Nov. 5th.
 6. G.D. 19h. 40m. Dec. 2nd, and 4h. 20m. Dec. 2nd.
 7. G.D. 26h. 5m. Feb. 13th (*a*), } and 12h. 45m. Feb. 13th (*a*).
or, 2h. 5m. Feb. 14th (*Note* 1) }
 8. G.D. 25h. 51m. Aug. 30th (*a*), } and 17h. 49m. Aug. 30th (*a*).
or, 1h. 51m. Aug. 31st (*Note* 1) }
-

EXAMPLES.

1. A chronometer was 4m. 24s. *fast* on mean time at Portsmouth ; find its error on M.T. Greenwich.

The chronometer was	} 4m. 24s. fast on M.T. Portsmouth
The long. of Portsmouth	} 4 24 slow on M.T. Greenwich
is 1° 6' west, or	}
Error of chronometer is	0 0

Here it is evident that since the time at Portsmouth is behind that of Greenwich, the chronometer, which was as much *fast* for Portsmouth as Portsmouth is *slow* on Greenwich, can have no error, and shows correct M.T. Greenwich.

2. A chronometer was 22m. 15s. *fast* on M.T. at Cape Town; find its error on Greenwich.

The chronometer was	} oh. 22m. 15s. fast on M.T. Cape Town
The long. of Cape Town	} 1 13 52 fast on M.T. Greenwich
is 18° 23' east, or	}
Error of chronometer is	1 36 7 fast on M.T. Greenwich

Here the time at Cape Town, being in *advance* of that at Greenwich, and the chronometer also in advance of Cape Town, the *sum* of these two quantities must be the error of the chronometer *fast* on M.T. Greenwich.

MISCELLANEOUS EXAMPLES.

1. On Jan. 1st, a chronometer was *slow* 3m. 15s. on mean time at Plymouth, in long. $4^{\circ} 9' 30''$ W. ; what was its error on Greenwich mean time ?

Ans. Error, oh. 19m. 53s. slow.

2. On May 25th, a chronometer was *fast* 32m. 30s. on mean time at Cork, in long. $8^{\circ} 27' 42''$ W. ; what was its error on mean time at Greenwich ?

Ans. Error, oh. 1m. 21s. slow.

3. On Nov. 20th, a chronometer was 5h. 18m. 10s. *slow* on mean time at Calcutta, in long. $88^{\circ} 19' 12''$ E. ; what was its error on mean time at Greenwich ?

Ans. Error, oh. 35m. 7s. fast.

4. On Aug. 4th, a chronometer was 4h. 22m. 15s. *fast* on Greenwich mean time ; what was its error on the clocks at New York, in long. $74^{\circ} 1'$ W. ?

Ans. Error, 9h. 18m. 19s. fast on New York.

5. A chronometer was 2m. 15s. *slow* for mean time at Funchal, Madeira, in long. $16^{\circ} 54' 42''$ W. ; and on arrival at Sta. Cruz, Teneriffe, in long. $16^{\circ} 14' 42''$ W., it was found to be 1m. 16s. *fast* on the mean time at that place ; what was its error on Greenwich mean time ?

Ans. Error, 1h. 3m. 43s. slow.

6. A telegram is made from Greenwich to Lisbon, in long. $9^{\circ} 8' 15''$ W., at 8h. 15m. 30s. A.M., and which takes 19 seconds in transmission ; what time should the clocks at Lisbon show when the telegram is received ?

Ans. Should show 7h. 39m. 16s. Lisbon time.

7. On Sept. 16th, Cape Hatteras, in long. $75^{\circ} 30'$ W., bore due N. (true) exactly at mean noon, when a chronometer showed 5h. 2m. os. ; what was its error on mean time at Greenwich ?

Ans. No error.

8. May 21st, at mean noon, the east point of Barbadoes (W. Indies), in long. $59^{\circ} 25'$ W., bore S. (true), when a chronometer showed 3h. 56m. 40s.; what was its error on mean time at Greenwich?

Ans. Error, oh. 1m. 0s. slow.

9. June 20th, at mean noon, the centre of Amsterdam I., in long. $77^{\circ} 35'$ E., bore due S. (true); what time ought a chronometer to show, if it was correct to mean time Greenwich, and of what date?

Ans. 18h. 49m. 40s., June 19th.

10. On Nov. 21st, at about 8h. A.M. mean time at ship, in long. $60^{\circ} 45'$ W., a chronometer showed oh. 4m. 10s., and its error on mean time Greenwich was 14m. 20s. fast; what is the correct mean time by chronometer, and of what date?

Ans. M.T. Chron. 23h. 49m. 50s., Nov. 20th.

11. On Oct. 6th, a chronometer showed 4h. 17m. 54s. the time at ship being about 3h. P.M. and long. $47^{\circ} 14' 30''$ E., error of chronometer *fast* on Greenwich mean time 5h. 9m. 20s.; what was the correct time by the chronometer, and of what date?

Ans. M.T. Chron. 23h. 8m. 34s., Oct. 5th.

12. Aug. 31st, a chronometer showed oh. 5m. 30s., the time at ship being about 8h. A.M. and the long. $61^{\circ} 10' 30''$ W., the error of the chronometer 1m. 55s. *slow* on Greenwich mean time; what was the correct time by the chronometer, and of what date?

Ans. M.T. Chron. oh. 7m. 25s., Aug. 31st.

13. At *mean* noon on June 1st, the Lizard, in lat. $49^{\circ} 58'$ N. and long. $5^{\circ} 12'$ W., bore N.N.E. by *compass*, distant 20 miles, var. 23° W., dev. 7° W., when a chronometer showed oh. 24m. 15s. mean time Greenwich; what was its error?

Ans. Error, oh. 3m. 15s. fast.

14. At *app.* noon, on June 16th, Cape Clear, in lat. $51^{\circ} 26'$ N. and long. $9^{\circ} 22'$ W., bore E. by *compass*, distant 16 miles, var. 24° W., dev. 2° E., when a chronometer showed oh. 44m. 20s.; what was its error on mean time at Greenwich?

Ans. Error, oh. 5m. 1s. fast.

15. A set of morning sights was taken for the longitude, the watch used showed 8h. 30m. 10s., and the corresponding time by a chronometer (A) was 8h. 38m. 10s. At 3h. 50m. 14s. P.M., by the same watch, another set of sights was taken; what corresponding time ought the chronometer to show, supposing neither the rate of the watch nor of the chronometer had altered in the interval?

Ans. 3h. 58m. 14s. P.M.

16. At 9h. 6m. 22s. A.M., a watch was *fast* 1h. 2m. 10s. on a chronometer (C). At 3h. 6m. 22s. P.M. a set of sights was taken, by the same watch, the rate of which was 24s. *losing*; what was the corresponding time by the chronometer, if it had no appreciable rate?

Ans. Time, 2h. 4m. 6s. P.M.

17. At 8h. P.M. a watch is exactly right for ship mean time; by 8h. A.M. on the following morning we have changed our longitude 72 miles to the *eastward* of our former position; how much must I alter the watch to set it to ship mean time?

Ans. To be put forward 4m. 48s.

18. At noon, in lat. 30° S., my watch is set right to mean time at ship, and by the following *noon* I have made a S.W. (true) course, distance 120 miles; what is the error of my watch on mean time at ship?

Ans. Error, 6m. 28s. fast.

19. On June 1st, a chronometer (A) is 15m. fast on M. T. G. and daily rate 3s. losing; when a chronometer (B) is found to be 1m. fast of (A). On June 21st they are again compared, and (B) is found to be 1m. slow on (A); what is B's daily rate, and error on M. T. G.

Ans. Daily rate of (B), 9s. losing.
Error on M. T. G., 13m. fast.

TO CORRECT THE RIGHT ASCENSION OF
MEAN SUN.

(*Sidereal Time.*)

RULE.

HAVING found the Greenwich date from the ship mean time and longitude, take from the Nautical Almanac (page II. of the month) the R.A. mean sun (given under the heading "Sidereal Time"), for noon of the Greenwich date.

Then from Table (o.) I. of Inman, page 10,* or from the Table of "Time Equivalents" in the Nautical Almanac take the increase of sidereal time (R.A. mean sun) corresponding to the hours and minutes of the *Greenwich date*, which *add* to the sidereal time taken out; the sum will be the sidereal time (R.A. mean sun) required.

EXAMPLE.

On Sept. 7th, at 9h. 50m. P.M. mean time at ship, in long. $34^{\circ} 28' E.$; find the right ascension of mean sun.

Green. Date.	Right Ascension Mean Sun (or Sidereal Time).								
Time at ship	h.	m.	s.	Sept.	7th	Sid. time, noon, Sept. 7th	h.	m.	s.
Long. in time	9	50	0	Sept.	7th	Correction, Tab. (o.) I.:	11	5	20.34
Green. date	2	17	52	E.	E.	For 7h.	+	1	8.99
	7	32	8	Sept.	7th	For 32m.	+	0	5.26
						For 8s.	+	0	0.02
						R.A. mean sun	11	6	34.61
						(Or sidereal time)			

EXAMPLES.

1. Find the right ascension of mean sun, on Sept. 10th, at 7h. 30m. M. T. Green.
2. On Sept. 14th, at 19h. 44m. 35s. M. T. G.
3. On Sept. 16th, at 2h. 14m. M. T. G.
4. On Sept. 20th, at midnight, in long. $60^{\circ} W.$, for M. T. G.
5. On Sept. 22nd, at 7h. 50m. A.M., in long. $42^{\circ} E.$

ANSWERS.

	h.	m.	s.
1.	11	18	24
2.	11	36	11
3.	11	41	11
4.	11	59	13
5.	12	3	20

TO FIND WHAT
STARS WILL PASS THE MERIDIAN
 BETWEEN ANY TWO GIVEN SHIP MEAN TIMES.*

RULE.

FOR each given ship mean time get a Greenwich date.

Then take from the Nautical Almanac the right ascension of mean sun and correct it for the hours and minutes of each Greenwich date, by Table (o.) I. Inman.

Add together each ship mean time expressed astronomically, and the corresponding R.A. mean sun (rejecting 24 hours from either sum if greater than 24 hours), which will give for results two separate right ascensions of meridian.

Then look in the Catalogue of Fixed Stars in the Nautical Almanac, and all those whose *right ascensions* are found to be between the two results (above), taken in order as limits, will pass the meridian between the given ship times.

Note.—The principal stars used for finding the latitude are given in the Nautical Almanac, pp. 336 to 385.

EXAMPLE.

What principal bright stars will pass the meridian of a ship in longitude 54° E. between 8h. and 11h. P.M. on Sept. 17th?

1st time at ship	8h. om. Sept. 17th	2nd time at ship	11h. om. Sept 17th
Long. in time	<u>3 36</u> E.	Long. in time	<u>3 36</u> E.
1st Green. date	4 24 Sept. 17th	2nd Green. date	7 24 Sept. 17th

* This rule will enable an observer to ascertain whether a star of sufficient magnitude will pass the meridian about the time at which he may require the latitude.

	h.	m.	s.		h.	m.	s.
R.A.M. sun 17th Sept.	11	44	45·87		11	44	45·87
Corr. Tab. (o.) 1, Inman:							
For 4h.	+	0	39·42	For 7h.	+	1	8·99
For 24m.	+	0	3·94	For 24m.	+	0	3·94
1st time at ship ...	11	45	29·23	2d time at ship ...	11	45	58·80
	8	0	0		11	0	0
1st R.A. meridian ...	19	45	29	2d R.A. meridian	22	45	59

Then the stars will be those in the Catalogue* whose right ascensions are *greater* than the 1st R.A. meridian, and *less* than the 2nd R.A. meridian, or, from α Pavonis to α Gruis (inclusive).

EXAMPLES.

1. What bright stars will pass the meridian of a ship in longitude 75° W. between 11h. P.M., Sept. 15th, and 1h. A.M. on Sept. 16th?

Ans. From Fomalhaut to β Ceti.

2. What bright stars will pass the meridian of Portsmouth, in long. $1^\circ 6'$ W., on Sept. 7th, between 8h. and 11h. P.M.?

Ans. From α Aquilæ to α Gruis.

3. What bright stars will pass the meridian of Greenwich between 7h. and 9h. P.M., Sept. 20th?

Ans. From α Aquilæ to α Cygni.

4. What bright stars will pass the meridian of Paris, in long. $2^\circ 20'$ E., between 2h. and 5h. A.M., Sept. 20th?

Ans. From α Arietis to α Tauri.

5. What bright stars will pass the meridian of Cape Blanco (Patagonia), in long. $65^\circ 44'$ W., between 8h. and 10h. P.M., Sept. 21st?

Ans. From α Pavonis to α Gruis.

* Stars of less than the 3rd magnitude are omitted in the Catalogue appended to this book.

TO FIND THE SHIP MEAN TIME WHEN A STAR IS ON THE MERIDIAN.

RULE.

Take from the Nautical Almanac the right ascension of the star, and also the right ascension of the mean sun.

Then, from the right ascension of the star (increased if necessary by 24 hours), subtract the right ascension of the mean sun; the remainder will be the ship mean time when the star is on the meridian.

Note.—Where the mean time of passage is required very accurately, the right ascension of the mean sun must be corrected for the *hours* and *minutes* of the Greenwich date (by Table (o.) I, Inman).

EXAMPLE.

At what mean time (nearly) will the star *α Aquilæ* pass the meridian of a ship in longitude $35^{\circ} 30' E.$ on Sept. 7th?

R.A. <i>α Aquilæ</i> , Sept. 7th	h.	m.	s.
			19	44	15
R.A. mean sun, Sept. 7th	11	5	20
Mean time of passage at ship (nearly) ...			8	38	55 Sept. 7th

The same (accurately, see *Note*).

	h.	m.							
Time at ship	8	39	Sept. 7th	R.A. mean sun, Sept. 7th	11	5	20	34	
Long. in time	2	22	E.	Corr. Tab. (o.) I, Inman :					
				For 6h.	+	0	59	14
Greenwich date	6	17	Sept. 7th	For 17m.	+	0	2	79
				R.A. mean sun	11	6	22
				h.	m.	s.			
R.A. <i>α Aquilæ</i>		19	44	15			
R.A. mean sun		11	6	22			
Mean time of passage*	...	8	37	53					

* Having found by the last Rule what stars will pass the meridian between any two given ship times, we can ascertain as above at what moment one of those stars will be on the meridian, and therefore suitable for determining the latitude by meridian altitude.

EXAMPLES.

1. At a ship in longitude 45° W., it is required to find at what time the star *a Aquilæ* will pass the meridian on Sept. 10th.

Ans. 8h. 25m. 2s.

2. At what time will *a Cephei* be on the meridian of a ship in longitude 70° E. on Sept. 8th?

Ans. 10h. 5m. 13s.

3. At what time will *a Scorpii* pass the meridian of a ship in longitude 150° E. on Sept. 7th?

Ans. 5h. 12m. 32s.

4. At what time will *Spica* pass the meridian of a ship in longitude $179^{\circ} 59'$ W. on Sept. 16th?

Ans. 1h. 35m. 4s.

Note.—If the zenith distance be required, take the difference between the declin. of the body and the latitude when the names are alike; marking it with the common name if the declin. be the greater, and with a contrary name if the declin. be the less. When the lat. and declin. are unlike, take their sum, and put to it the name of the declin. The result will be the meridian zenith distance, N. or S. of the zenith.

TO FIND THE LATITUDE BY THE MERIDIAN ALTITUDE OF A FIXED STAR.

RULE.

CORRECT the observed altitude for $\left\{ \begin{array}{l} \textit{Index error,} \\ \textit{Dip,} \\ \textit{Refraction,} \end{array} \right.$

which will give the true altitude.

Subtract the true altitude from 90° , and the remainder will be the *meridian zenith distance*, which mark N. or S. according as the observer is N. or S. of the star.

Then, from the Naut. Almanac (Table of Fixed Stars) take out the declin. for the day *nearest* to that at ship; mark it with its proper name N. or S., and put it underneath the zen. dist:

If the zen. dist. and declin. are of *like* names, their *sum* will be the latitude, of the name of either.

If the zen. dist. and declin. are of *contrary* names, their *difference* will be the latitude, of the name of the *greater*.

EXAMPLE.

January 15th, the obs. mer. alt. Markab. was $51^\circ 10' 15''$ (Z.N.), I.E. $-1' 10''$, dip 21 ft.; required the latitude.

Obs. alt.	51° 10' 15" (Zen. N.)
Index error	— 1 10
	<hr style="width: 100%;"/>
Dip	51 9 5
	— 4 31 (Tab. e.)
	<hr style="width: 100%;"/>
Ref.	51 4 34
	— 0 47 (Tab. n.)
	<hr style="width: 100%;"/>
True alt.	51 3 47
	90 0 0
	<hr style="width: 100%;"/>
Zen. dist.	38 56 13 N.
Declin.	14 29 6 N.
	<hr style="width: 100%;"/>
Latitude	53 25 19 N.

**TO FIND THE LATITUDE BY THE MERIDIAN
ALTITUDE OF A FIXED STAR UNDER THE POLE.**

RULES.

(A) CORRECT the observed altitude for $\left\{ \begin{array}{l} \text{Index error,} \\ \text{Dip,} \\ \text{Refraction,} \end{array} \right.$
which will give the true altitude.

Then *add* 90° to the true altitude, and from the sum *subtract* the star's declin. ; the remainder will be the latitude, *always* of the same *name* as the Pole.

Or,

(B) Correct the observed altitude for $\left\{ \begin{array}{l} \text{Index error,} \\ \text{Dip,} \\ \text{Refraction,} \end{array} \right.$
which will give the true altitude.

Then *add* the star's polar distance ($90^\circ - \text{declin.}$) to the true altitude, the sum will be the latitude, of the same *name* as the Pole.

EXAMPLE (A).

June 20th, the obs. mer. alt. α Cephei (under N. Pole), was $30^\circ 10' 30''$, I.E. $+2' 15''$, height of eye 15 ft. ; required latitude.

Obs. alt.	30° 10' 30"	
Index error	+ 2 15	
	<hr style="width: 100%;"/>	
	30 12 45	
Dip	- 3 49 (Tab. e.)	
	<hr style="width: 100%;"/>	
	30 8 56	
Ref.	- 1 40 (Tab. n.)	
	<hr style="width: 100%;"/>	
True alt.	30 7 16	
Add	90 0 0	
	<hr style="width: 100%;"/>	
	120 7 16	
Declin.	62 1 6 N. (p. 379, N.A.)	
	<hr style="width: 100%;"/>	
Latitude	58 6 10 N.	

EXAMPLE (B).

True alt.	30° 7' 16"	
Pol. dist.	27 58 54	
	<hr style="width: 100%;"/>	
Latitude	58 6 10 N.	

TO CORRECT THE SUN'S DECLINATION.

RULE.

BY HOURLY DIFFERENCES.

HAVING found the Greenwich date from the ship time and longitude, take out of the Nautical Almanac the sun's declin. for apparent or mean noon of the Greenwich date (as required by the Question) ; also take from page I. the difference for 1h.

Multiply this *diff.* by the hours and decimal parts of an hour (*Note 2*) of the *Green. date*: reduce the product to minutes and seconds and *add* it to the declin. (taken out) when the declin. is *increasing*, or *subtract* it from the declin. (taken out) when the declin. is *decreasing*: the result will be the sun's declin. at the given *Greenwich time*.

Note 1.—The correction for hourly *diff.* is added to or subtracted from the declin. when it (*the declin.*) is increasing or decreasing, and not because the *hourly diff.* increases or decreases.

Note 2.—To find the decimal parts of an hour, divide the given minutes by 6. Thus, 7h. 42m. = $7^{\text{h}}7$, and 19h. 20m. = $19^{\text{h}}3$ (nearly).

Note 3.—In the case where the *correction* (when subtractive) *exceeds* the declin. for the day taken out, the *difference* between the declin. and the correction will give the declin. required, but which must then be marked with a *contrary* name to that given in the Nautical Almanac.

TO FIND THE LATITUDE BY THE MERIDIAN ALTITUDE OF THE SUN.

RULES.

(A) *In West Longitude.*

1. Turn the longitude of the ship into time, and *add* it to oh. om. (see *Note*). The sum, putting before it the day of the month at *ship*, will be the Greenwich date.

(B) *In East Longitude.*

2. Turn the longitude of the ship into time, and *subtract* it from 24h. om. (see *Note*). The remainder, putting before it the day of the month *preceding* that at *ship*, will be the Greenwich date.

3. Take from the Nautical Almanac (page I. of the month) the declin. at apparent noon, and also the hourly diff.; multiply the hourly diff. by the hours and decimal parts of an hour of the Greenwich date, and apply the correction thus obtained to the declin., which mark with its proper name, N. or S., as in the preceding Rule.

4. Also take out the sun's semidiameter from the Nautical Almanac (page II. of the month), for noon of the Greenwich date.

5. Correct the observed meridian altitude,

For	{	Index error,	+ <i>off the arc,</i>	or	- <i>on the arc.</i>	
		Dip (Tab. e. Inman)	<i>always -</i>			
		Ref. - Par. (Tab. m. Inman)	<i>always -</i>			
		Semidiameter	+ L.L.	<i>- U.L.</i>		

which will give the true altitude of sun's centre.

Note.—When finding the *Greenwich date* in this question, it is usual to consider NOON (the time of observation), as oh. om. of the *day at ship*; or as 24h. om. of the *day preceding* that at *ship*, both being the same moment of time.

6. Subtract the true altitude from 90° , which will give the zenith distance ; mark this N. or S., according as the *observer* is to the N. or S. of the sun.

7. Under the zenith distance place the declin. (both being marked with their proper names, N. or S.).

When the names are *alike* take the *sum*, which will give the latitude, of the name of either zenith distance or declin.

8. When the zenith distance and declin. are of *contrary* names, take the less from the greater for the latitude, which will be of the name of the greater.

EXAMPLE.

Sept. 28th, in long. 45° W., the obs. mer. alt. of the sun's L.L. was $47^\circ 10' 30''$ (zen. N.). Index error $-2' 40''$, height of the eye 18 feet ; required the latitude.

(1) Green. Date.	(3) Declin., page I, N. A.	(3) Change, p. I, N. A.
Time at Ship (NOON)* or $0^h 0^m$ Sp. 28 Long. in time $3^h 0^m$ W.	Dec. 28 Sept. $2^\circ 1' 54''$ S. Corr. for 3h $+ 2 55$	Hourly diff. $58'' 46$ × by G.D. 3
Green. date $3^h 0^m$ Sp. 28	Reduced dec. $2 4 49$ S.	$6,017,5'' 38$
* (See <i>Nota</i> .)		Corr. for 3h. G.D. $2' 55''$

To find the Latitude.

(5) Obs. alt. L.L.	$47^\circ 10' 30''$ (Zen. N.)
In. error	$- 2 40$
	<hr/>
Dip 18 ft.	$47 7 50$ $- 4 11$ (Tab. e.)
	<hr/>
App. alt. L.L.	$47 3 39$
Ref.	$- 0 48$ (Tab. m.)
	<hr/>
True alt. L.L.	$47 2 51$
(4) Semid.	$+ 16 1$ N.A. p. II. month.
	<hr/>
True alt. centre	$47 18 52$ 90
	<hr/>
(6) Zen. dist.	$42 41 8$ N.
Declin.	$2 4 49$ S.
	<hr/>
(8) Latitude	$40 36 19$ N.

LATITUDE BY POLE STAR (POLARIS).

RULE.

Nautical Almanac Method.

1. WITH the mean time at ship and longitude, get the Greenwich date.

2. For this date take out the *sidereal time* (R.A. mean sun) (page II. of the month in the Nautical Almanac).

Then enter Table (o.) I, Inman, page 10,* and take out the increase for the hours and minutes of the Greenwich date, which correction must be *added* to the sidereal time.

3. To the sidereal time (R.A. mean sun) thus corrected, *add* the *astronomical time at ship*. Call the result the right ascension of meridian (or sidereal time *of observation*).

Note.—When the *sum* of the sidereal time and *astronomical time at ship* exceeds 24h., reject 24h. and use the remainder in looking out the 1st, 2nd, and 3rd corrections from the Nautical Almanac.

4. Correct the observed altitude for index error, dip, and refraction, and get the true altitude.

5. From the true altitude subtract 1'. This will give the reduced altitude.

6. Enter Table I. (page 499 of Nautical Almanac) with the *sidereal time of observation*, on its proper side of the column of *Correction*. Take out the *first* correction, which must be marked with the algebraic sign *next* to the column "sidereal time."

7. Apply this correction to the reduced altitude, according to the sign + or -. This will be the *approximate* latitude.

8. Then enter Table II. (pages 500, 501 of Nautical Almanac) with the *nearest* degree to the true altitude at the *top*, and sidereal time of observation at the *side*. Take out the *second* correction, which is *always to be added*.

9. Then enter Table III. (pages 500, 501 of Nautical Almanac) with the month at the *top*, and hour of sidereal time of observation at the *side*, take out the *third* correction, which is *always to be added*.

10. Add these two corrections to the *approximate* latitude, the sum will give the latitude (*always north*).

* By *exact* interpolation of the Tables, a difference of a few seconds in the latitude would have resulted : but at sea, for all practical purposes, *seconds* may be omitted.

EXAMPLE—A.M. TIME AT SHIP.

Sept. 4th, at 4h. 52m. A.M. (mean time) in long. 60° E., the obs. alt. of the Pole Star was 20° 10' 40'', index error -2' 15'', height of the eye 15 feet; required the latitude.

(1) Green. Date.	(2) Sidereal Time (or R.A. Mean Sun).	(3) R.A. Meridian (or Sid. Time of Obs.).
Time at ship $\begin{array}{r} \text{h. m.} \\ 4 \ 52 \text{ A.M., Sept. 4th} \\ 12 \end{array}$	R.A.M. sun $\begin{array}{r} \text{h. m. s.} \\ 10 \ 49 \ 34 \cdot 14 \end{array}$	R.A.M. sun $\begin{array}{r} \text{h. m. s.} \\ 10 \ 51 \ 41 \end{array}$
<i>Ast. T. ship.</i> 16 52 Sept. 3rd	Corr. Tab. (o.) I:	<i>Ast. T. ship</i> $\begin{array}{r} 16 \ 52 \ 0 \\ \hline 27 \ 43 \ 41 \end{array}$
Long. in time $\begin{array}{r} 4 \ 0 \ \text{E.} \\ \hline 12 \ 52 \ \text{Sept. 3rd} \end{array}$	For 12h. + 1 58' 28	Sub. $\begin{array}{r} 27 \ 43 \ 41 \\ 24 \\ \hline 3 \ 43 \ 41 \end{array}$
Green. date $\begin{array}{r} 12 \ 52 \ \text{Sept. 3rd} \\ \hline \end{array}$	For 52m. + 0 8' 54	R.A. merid. $\begin{array}{r} 3 \ 43 \ 41 \end{array}$
	R.A.M. sun $\begin{array}{r} 10 \ 51 \ 40 \cdot 96 \\ \hline \end{array}$	(Or sid. time of obs.)
	(Or, sid. time.)	(See Note to 3.)

Computation of the Latitude.

(4)	Obs. alt.	20° 10' 40''
	Index error	- 2 15
		<hr/>
	Dip, 15 feet	20 8 25
		- 3 49 (Tab. e.)
		<hr/>
	App. alt.	20 4 36
	Ref.	- 2 39 (Tab. n.)
		<hr/>
(5)	True alt.	20 1 57
	Subtract	- 1 0
		<hr/>
	Reduced alt.	20 0 57
(6)	{ Sid. time of obs. 3h. 40m. from Tab. I., N.A., page 499, gives <i>first</i> correction }	- 1 5 45
(7)	Approximate lat.	18 55 12
(8)	{ True alt. 20° 0' at top of Tab. II., N.A., page 500, and sid. time obs. 3h. 30m. at side, gives <i>second</i> correction }	+ 0 10
(9)	{ Month (Sept. 4th) taken from Tab. III., p. 501, as Sept. 1st, and 4h. 0m. sid. time of obs., gives <i>third</i> correction }	+ 0 28
(10)	Latitude*	<hr/> 18 55 50 North.

* By exact interpolation of the Tables, a difference of a few seconds in the latitude would have resulted: but at sea, for all practical purposes, seconds may be omitted.

VARIATION BY AMPLITUDE (SUN).

RULE.

1. WITH the time at ship and longitude, get a Greenwich date.
2. For this date take out the sun's declination from the Nautical Almanac (page I. or II. of the month), according as apparent or mean time is given, and correct it.

3. Then, $\sin. \text{true amp.} = \sin. \text{dec.} \times \sec. \text{lat.}$

or,

Add together the log. sine of the declin., and the log. secant of the lat., the sum (rejecting the tens in the index), will be the log. sine of the true amplitude.

4. Write E. *before* the true amplitude when the body is *rising*.

5. Write W. *before* the true amplitude when the body is *setting*.

6. Write N. *after* the true amplitude when the declin. is N.

7. Write S. *after* the true amplitude when the declin. is S.

8. Then, correct the compass amplitude for deviation, by allowing it to the *right* when easterly, and to the *left* when westerly. The result will be the *correct magnetic amplitude*. (See *Note*.)

9. Place the *correct magnetic* amplitude under the true amplitude.

10. When they are both reckoned *towards* N., or both reckoned *towards* S., their *difference* will be the *variation*.

11. When one is reckoned *towards* N., and the other *towards* S., their *sum* will be the *variation*.

12. When the *true* amplitude is to the *right* of the magnetic amplitude, the variation will be *easterly*.

(Supposing yourself at the CENTRE of the compass.)

13. When the true amplitude is to the *left* of the magnetic amplitude, the variation will be *westerly*.

(Supposing yourself at the CENTRE of the compass.)

Note.—Should the *correct magnetic amplitude* not be reckoned *from* the same point as the true amplitude (that is *from* E. or W.), take the correct magnetic amplitude from 90° to make it correspond to the true amplitude.

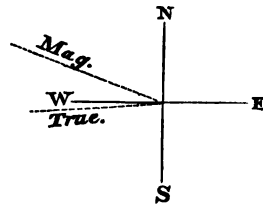
EXAMPLE.

September 23rd, at about 6h. om. P.M. *mean time*, at ship, in lat. $40^{\circ} 9' 15''$ S., and long. $52^{\circ} 30'$ E., the sun's setting amplitude* by compass was N. $60^{\circ} 10'$ W., deviation $8^{\circ} 40'$ W.; required the variation.

(1) Green. date.	(2) Declin. page II. N. A.	(2) Change, p. I. N. A.
Time at ship Long. in time Green. date	Dec. Sept. 23rd $0^{\circ} 5' 2''$ S. Correction for } 2h. 30m. } + 2 26 Declin. <u>0 7 28</u> S.	Hourly diff. $58'' \cdot 48$ × by G.D. <u>2 '5</u> 29240 11696 <hr/> 6,0)14,6'200 Corr. <u>2' 26'</u>

(3) Sin. true amp. = sin. dec. sec. lat.

Sin dec. $0^{\circ} 7' 28''$ S. ... 7'336852†
Sec. lat. $40^{\circ} 9' 15''$... 10'116729
Sin. true amp. 7'453581†



True amp. (5) W. $0^{\circ} 9' 46''$ S. (7) Compass amp. N. $60^{\circ} 10'$ W.
(9) Corr. mag. amp. W. 21 10 0 N. Deviation 8 40 left.

(13) Variation 21 19 46 W. N. 68 50 W.
90 (see Note)

(8) Corr. mag. amp. W. 21 10 N.

* To find the time of sunset or sunrise by Inspection. Table (e) Inman :-

By the time amplitude is meant the time the sun rises before or after 6 A.M., or sets before or after 6 P.M. When the lat. and declin. are both N. or both S., the sun rises so much *before* 6 A.M., or sets so much *after* 6 P.M. When the lat. and declin. are one N. and the other S., the sun rises so much *after* 6 A.M., and sets so much *before* 6 P.M.

Enter the Table with declin. at top and latitude at the side, and take out the corresponding time, which *add* to or *subtract* from 6 hours, apparent time, according to the above explanation.

† Using Table (s. Inman).

VARIATION BY AZIMUTH (SUN).

RULE.

1. WITH the ship mean time and longitude get the Green. date.

2. For this date take out the sun's declination from the Nautical Almanac (page II. of the month), and correct it. Take out also the sun's semid.

3. Find the sun's polar distance.

(a) By subtracting the declin. from 90° , when the latitude and declin. are of the *same* names.

(b) By adding 90° to the declin., when the latitude and declin. are of *contrary* names.

4. Correct the observed altitude for index error, dip, refraction, and semid., and thus get the true altitude of the sun's centre.

5. Under the latitude put the true altitude and take the *difference*. Under this difference put the polar distance, and take the *sum* and *difference*.

Then, take out the log. *secant* of the latitude.

log. *secant* of the true altitude.

$\frac{1}{2}$ log. *hav.* of the sum.

$\frac{1}{2}$ log. *hav.* of the difference.

Add these 4 logs. together (reject the tens in the index) and look it out as a *hav.* (in degrees).

This will give the *true bearing*, to be marked as follows:—

6. Write N. *before* the degrees of the true bearing when the latitude is N.

7. Write S. *before* the degrees of the true bearing when the latitude is S.

8. Write E. *after* the degrees of the true bearing when the body is east of meridian.

9. Write *W.* after the degrees of the true bearing when the body is west of meridian.

10. Correct the *compass* bearing for deviation, allowing it to the *right* when the deviation is *easterly*, and to the *left* when the deviation is *westerly*. Call the result the *correct magnetic* bearing.

11. Under the true bearing put the correct magnetic bearing, and take the *difference*.

This will be the *variation* of the compass.

12. *East*, when the *true* bearing is to the *right* of the magnetic bearing, supposing yourself at the CENTRE of the compass.

13. *West*, when the *true* bearing is to the *left* of the magnetic bearing, supposing yourself at the CENTRE of the compass.

(A) When observing the azimuth of a heavenly body with the compass, the bearing should always be reckoned from North or from South, never from East or from West; and remember that the *magnetic* azimuth must be reckoned from the *same* point (N. or S.) as the *true* azimuth. To effect this, the compass bearing will sometimes have to be subtracted from 180° .

(B) It may occasionally happen that a question is given in which the magnetic bearing is reckoned *towards* the opposite point to that *from* which the true bearing is reckoned, in which case 90° must be *added* to the magnetic bearing.

(C) When the magnetic bearing is reckoned *towards* the same point as that *from* which the true bearing is reckoned, the magnetic bearing must be *subtracted* from 90° .

Note.—By observing the sun's azimuth, and comparing the variation of the compass so determined with the *known* variation at the place of observation, the deviation on the course (at the time the azimuth is taken) can at once be ascertained; and therefore the deviations for *any given directions* of the ship's head.

EXAMPLE.

September 24th, at 3h. 10m. P.M. (mean time) in lat. $10^{\circ} 15' N.$, and long. $168^{\circ} 0' E.$, the obs. alt. sun's L.L. was $39^{\circ} 28' 0''$, index error $+ 1' 15''$, height of the eye 18 feet, when it bore by standard compass S. $84^{\circ} 20' W.$, dev. $13^{\circ} W.$; required the variation.

(1) Green. date.	(2) Declin. page II., N. A.	(2) Change, p. I., N. A.
Time at ship h. m. Sept. 24th 3 10 24	Dec. Sept. 23rd $0^{\circ} 5' 2'' S.$ Corr. $+ 15' 36''$	Hourly diff. $58'' \cdot 48$ x by G.D. 16
Long. in time 27 10 Sept. 23rd 11 12 E.	Declin. $0 20 38 S.$ 90	35088 5848
Green. date 15 58 Sept. 23rd	Pol. dist. (3b) $90 20 38$	6,093,568 Corr. $15' 36''$

(4) To find the True Altitude.

Obs. alt. L.L. $39^{\circ} 28' 0''$
 Index error $+ 1' 15''$

 Dip $39 29 15$
 $- 4 11$

 Ref. $39 25 4$
 $- 1 4$

 Semid. $39 24 0$
 $+ 15 59$

 True alt. centre $39 39 59$

(5) To compute the True Bearing.

Lat. $10^{\circ} 15' 0''$ Sec. $0^{\circ} 006987$
 True alt. $39 39 59$ Sec. $0^{\circ} 113638$

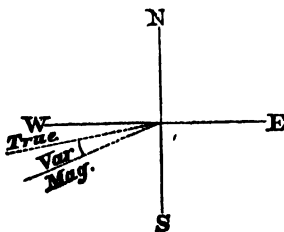
 Pol. dist. $90 20 38$

 Sum $119 45 37 \frac{1}{2}$ Hav. $4^{\circ} 937000$
 Diff. $60 55 39 \frac{1}{2}$ Hav. $4^{\circ} 705013$

 Haversine $9^{\circ} 762638$

True bearing (6) N. $99^{\circ} 5' 0'' W.$ (9)
 (11) Corr. mag. bearing N. $108 40 0 W.$

Variation $9 35 0 E.$ (12)



(10) Compass bearing S. $84^{\circ} 20' 0'' W.$
 Deviation $13 0 0 left.$

S. $71 20 0 W.$

 180

Corr. mag. bearing } N. $108 40 0 W.$
 ing, see (A)

TO CORRECT THE EQUATION OF TIME.

WITH the time at ship and longitude get a Greenwich date. For this date take out the equation of time from page II. of the month in the Nautical Almanac; also from page I. take out the *hourly difference*.

Multiply the *hourly difference* by the hours and decimal parts of an hour of the Greenwich date, which will give the *correction*.

Add this correction to the equation of time taken out, when the equation of time is *increasing*.

Subtract this correction from the equation of time taken out, when the equation of time is *decreasing*.

The sum or remainder will be the equation of time required.

Note 1.—When the *correction* exceeds the equation of time taken out from the Nautical Almanac, the *difference* between the correction and the equation of time will be the equation of time required.

Note 2.—Since the equation of time is used either for turning mean into *apparent* time, or for turning apparent into *mean* time, it will be necessary to attend to the precept given at the *head* of the column of equation of time.

Note 3.—When a line — is placed at the heading of the column, a similar one comes between two following days in the column itself; in which case, it means that *all* the equations of time *above* the line are to be applied according to the direction *above* the line in the heading; and *all* the equations of time *below* the line must be applied as directed *below* the line in the heading.

Note 4.—The line — is placed between those two days on which the equation of time *changes* from additive to subtractive, and *vice versâ*, showing that at some particular moment in the interval the equation of time has become om. os.

LONGITUDE BY CHRONOMETER (SUN).

RULE.

1. WITH the *estimated* mean time at ship and the longitude by *account*, get the approximate (or rough) Greenwich date.*

2. To the time *shown* by chronometer, apply its error on mean time at Greenwich.†

(a) When the error is *slow*, add it to the time shown by chronometer.

(b) When the error is *fast*, subtract it from the time shown by chronometer.‡

3. Then multiply the *daily rate* by the number of days and decimal parts of a day elapsed between the time at which the rate was *last* ascertained, and the *Greenwich* time.

Turn this product into minutes and seconds.

(a) *Adding* the result to the chronometer time (found in 2 above) when the rate is *losing*.

(b) *Subtracting* the result from the chronometer time (found in 2 above) when the rate is *gaining*.

(c) This will give the Greenwich *mean time by the chronometer*.§

(d) Affix to the mean time by chronometer the day of the month indicated by the approximate Greenwich date.||

4. Take out the sun's declination and equation of time from page II. of the month in the Nautical Almanac, and correct them for the mean time at Greenwich by the *chronometer*.

5. Mark the corrected equation of time "*added to*," or "*subtracted from*," apparent time, as directed by the heading of the column of Eq. Time in page I. of the month, in the Nautical Almanac.

* The approximate Greenwich time is not used in correcting the declination or equation of time.

† At sea, it is customary to *combine* the error and rate, and apply the result to the time *shown* by the chronometer.

‡ When the error is *fast*, and greater than the time shown by the chronometer, it will be necessary (before subtracting the error) to increase the chron. time by 12 or 24 hours, according as the *Greenwich* date shows the time to be near *midnight*, or near *noon*.

§ As the chronometer mean time is not always the proper time at *Greenwich*, it will occasionally require to be brought to that time by adding or subtracting 12 hours. When to do this will be known by comparing the chron. time with the *Greenwich* date, with which it must be made to agree.

|| When the *hour* shown by the chronometer happens to be on one side of *noon*, and the hour of the Greenwich date on the other, a consideration of the circumstances of the case will decide whether the chron. should be a day in *advance* of or *behind* the approx. Greenwich date.

6. Correct the observed altitude for index error, dip (ref. —par.) and semidiameter; get the true altitude, and find the *zenith distance*.*

7. To compute the *apparent time at ship* (sun's hour angle). Under the latitude place the declination, and mark each with its proper name, N. or S.

(a) When the names are *alike*, take the *difference*.

(b) When the names are *unlike*, take the *sum*.

Under the difference or sum, put the zenith distance, and take the *sum* and *difference*.

Then add together, the log. *secant* of the latitude.
 log. *secant* of the declination.
 $\frac{1}{2}$ log. *havens.* of the sum.
 $\frac{1}{2}$ log. *havens.* of the difference.

Reject the tens from the index of the sum of these four logs., and the remaining log. will be a *havering*.

8. Find the hours, minutes, and seconds corresponding to this *havering*.

(a) Taking them from the *top* of the page when the sun was observed *west* of meridian (or P.M. time at ship).

(b) And from the *bottom* of the table when the sun was observed *east* of meridian (or A.M. time at ship).

This will be the *apparent time at ship*.

9. To the apparent time at ship apply the equation of time, adding it, or subtracting it, according to the direction given in page I. of the month of the Nautical Almanac.

10. The result will be the *mean time at ship*, which must be marked with the proper *astronomical day* of the month.

11. Under the mean time at ship, place the mean time at Greenwich by *chronometer*. If the days of the month be different, add 24 hours to the one more advanced, and put the day *one back*. Take the *difference* of the times, which will give the longitude in *time*.

12. Turn this time into *degrees*, minutes, and seconds, which will give the *longitude* of the ship.

(a) West, when the Greenwich time is *greater* than the ship time.

(b) East, when the Greenwich time is *less* than the ship time.

Note.—To ascertain the *daily rate* when the errors on two different days are given, *see* (Additional Examples) at end of Book.

* When the altitude has been taken with an *artificial horizon*, the index error must be applied to it with its proper sign + or —, and the *result* divided by 2. Correct this for refraction and semidiameter as in ordinary cases, but not for *dip*.

EXAMPLE.

On September 30th, at about 8h. 20m. A.M., mean time, in lat. $30^{\circ} 10' N.$, and long., by account, $40^{\circ} 0' W.$, the obs. alt. of the sun's L.L. was $29^{\circ} 51' 50''$, index error $+ 2' 10''$, height of the eye 18 feet, when a chron. showed 9h. 44m. 44s.; find the true longitude of the ship.

At G. M. noon, September 20th, the chron. was slow 1h. 15m. 0s., and its daily rate was 4'5s. losing.

<p>(1) Approx. Green. date.</p> <p>Time at ship $\begin{matrix} h. & m. \\ 8 & 20 \end{matrix}$ A.M. Sep. 30th</p> <p>Ast. T. ship $\begin{matrix} 20 & 20 \end{matrix}$ Sept. 29th</p> <p>Long. in time $\begin{matrix} 2 & 40 \end{matrix}$ W.</p> <p>Green. date $\begin{matrix} 23 & 0 \end{matrix}$ Sept. 29th</p> <p><i>Note.</i>—This time is not to be used to correct the declin. or eq. time.</p>	<p>(4) Declin. p. II., N.A.</p> <p>Declin. 29th $\begin{matrix} 2^{\circ} & 25' & 26'' \\ & + & 22 & 23 \end{matrix}$ S.</p> <p>Declin. $\begin{matrix} 2 & 47 & 49 \end{matrix}$ S.</p>	<p>(4) Change, p. I., N.A.</p> <p>Hourly diff. $\begin{matrix} 58'' & 41 \\ & \times & \text{by G.M.T.} & 23 \end{matrix}$</p> <p>$\begin{matrix} 17523 \\ 11682 \\ \hline 60134,343 \end{matrix}$</p> <p>Corr. for 23h. $\begin{matrix} 22' & 23'' \end{matrix}$</p>
<p>(2) Green. Mean Time by Chron.</p> <p>Chron. showed $\begin{matrix} h. & m. & s. \\ 9 & 44 & 44 \end{matrix}$ Sept. 30th</p> <p>Error $\begin{matrix} 1 & 15 & 0 \end{matrix}$ slow</p> <p>$\begin{matrix} 10 & 59 & 44 \end{matrix}$</p> <p>Rate $\begin{matrix} s. \\ 4'5. \end{matrix}$ } $+ 0$ 45 losing</p> <p>Days $\begin{matrix} 10 \\ 45^0 \end{matrix}$ }</p> <p>Add $\begin{matrix} 11 & 0 & 29 \\ 12 \end{matrix}$</p> <p>Green. M. T. $\begin{matrix} 23 & 0 & 29 \end{matrix}$ Sept. 29th.</p>	<p>(4) Eq. of Time, p. II., N.A.</p> <p>Eq. time, 29th $\begin{matrix} m. & s. \\ 9 & 40'85 \end{matrix}$</p> <p>Corr. for 23h. $\begin{matrix} + & 18'86 \end{matrix}$</p> <p>Eq. time $\begin{matrix} 9 & 59'71 \end{matrix}$</p> <p>Sub. from App. Time. (5)</p>	<p>(4) Change, p. I., N.A.</p> <p>Hourly diff. $\begin{matrix} 820 \\ & \times & \text{by G.M.T.} & 23 \end{matrix}$</p> <p>$\begin{matrix} 2460 \\ 1640 \\ \hline 18860 \end{matrix}$</p> <p>Corr. for 23h. $\begin{matrix} 18'860 \end{matrix}$</p>

(6) *To find the Zen. Dist.*

Obs. alt.	$29^{\circ} 51' 50''$
In. error	$+ 2' 10''$
Dip	$\begin{matrix} 29 & 54 & 0 \\ & - & 4 & 11 \end{matrix}$
App. alt.	$29 49 49$
Ref.	$- 1 33$
Tr. alt.	$29 48 16$
Semid.	$+ 16 0$
Tr. alt. centre	$\begin{matrix} 30 & 4 & 16 \\ & 90 \end{matrix}$
Zen. dist.	$59 55 44$

(7) *To compute the App. Time at Ship.*

Lat.	$30^{\circ} 10' 0'' N.$	Sec.	$0^{\circ} 063201$
Declin.	$2 47 49 S.$	Sec.	$0^{\circ} 000517$
(7b) Sum	$32 57 49$	Zen. dist.	$59 55 44$
Sum	$92 53 33$	$\frac{1}{2}$ Hav.	$4^{\circ} 860172$
Diff.	$26 57 55$	$\frac{1}{2}$ Hav.	$4^{\circ} 367658$
		Haversine	$9^{\circ} 291548$
(8b) App. time at ship	$\begin{matrix} h. & m. & s. \\ 20 & 29 & 58 \end{matrix}$		
Eq. time	$\begin{matrix} - & 10 & 0 \end{matrix}$		
(10) Ast. mean time ship	$20 19 58$	Sept. 29th	
(3b) Green. mean time	$23 0 29$	Sept. 29th	
(11) Long. in time	$2 40 31$		
Longitude $40^{\circ} 7' 45'' W.$	(12a.)		

**LONGITUDE BY CHRONOMETER,
(MOON, PLANET, OR STAR.)**

RULE.

1. HAVING found the Greenwich mean time by chronometer, as directed in Rule for Chronometer, page 76.
2. Take from the N. A. the R. Asc. mean sun, and correct it ; also the Declin. and R. Asc. of the body, and correct them for Green. M. T. by chron. when necessary.
3. From the observed altitude get the true zenith distance.

To compute the hour angle.

4. Proceed as in (7) of Rule for Chron. page 77.

To find the mean time at ship.

5. To the hour angle add the R.A. of the body, and from the sum (increased if necessary by 24h.) subtract the R.A. mean sun. The remainder, diminished by 24h. if greater than 24h., will be the *mean time at ship*, which mark with the proper *astronomical* day of the month.

Then with the Green. mean time by chron. find the longitude, as in (11) and (12) of Rule, page 77.

EXAMPLE. (STAR).

Sept. 10th, at about 11h. om. P.M. mean time, in lat. $32^{\circ} 18' N.$, and long. D.R. $41^{\circ} 22' W.$, the obs. alt. of *a Lyrae* (west of mer.) was $44^{\circ} 7'$, index error $+ 1' 50''$, height of the eye 20 feet, when a chron. showed 1h. 5m. 20s. ; required the longitude of the ship.

On Sept. 1st, at G.M. noon, the chron. was 38m. 14s. slow, and its daily rate was 1.8 sec. losing.

(1) G.M.T. by chron.	h. m. s.	13 43 51	Sept. 10th.	R. A. mean sun	h. m. s.	11 19 25	}
(3) Zen. dist.		45° 56' 0''		R. A. <i>a Lyrae</i>	18 32 23		
				Dec. <i>a Lyrae</i>	38° 39' 39'' N.		
(4) Star's hour angle	h. m. s.	3 47 6					
Star's R. Asc.		18 32 23					
	Sum	22 19 29					
	R. A. mean sun	11 19 25					
(5) Mean time ship	h. m. s.	11 0 4	Sept. 10th.				
G.M.T. chron.		13 43 51	Sept. 10th.				
	Long. in time	2 43 47					
	Longitude	40° 56' 45'' W.					

LATITUDE BY DOUBLE ALTITUDE (SUN).

RULE.

1. GET a Greenwich date corresponding to *each* mean time at *ship*.

2. Take out the sun's declination from page II. of the month of the Nautical Almanac, and correct it separately for *each* Greenwich date. Take out also the sun's semidiameter.

3. Then with *each* declination find the corresponding polar distance.

4. Find the *polar angle in time*.

(a) When both times by chronometer are A.M. or both P.M., take their difference, which will give the polar angle *in time*.

(b) When one time by chronometer is A.M., and the other P.M., add 12 hours to the P.M. time, and take the A.M. time from it, the difference will be the polar angle *in time*.

Note.—The time at each observation should be taken by a *good* watch or chronometer, from which the *interval* between the two observations (polar angle) is to be obtained.

This should also be corrected for the *rate* of the watch or chronometer (if any of consequence) in the interval, adding it when *losing*, and subtracting it if *gaining*, which will give the polar angle in time.

5. Correct each observed altitude for index error, dip, semid., and (refraction — par.) then get the true altitudes.

6. Correct the *first* true altitude for the *run of the ship*, after which get the zenith distance. Also get zenith distance from second true altitude.

7. The first true altitude is corrected for *run*, as follows :—

Enter the Traverse Table (x. Inman) with the *distance* run in the *interval* in the *dist.* column, and the angle included between the true bearing of the sun at the *first* observation, and the true course of the ship (in the *interval*) as a *course*, and take out the corresponding *diff. lat.* in the same manner as when working the day's work.

(a) When the angle found above is *less* than 8 points, *add* the diff. lat. taken out of the Traverse Table, to the *first* true altitude.

(b) When the angle found above is *more* than 8 points, enter the Traverse Table with what it wants of 16 points, and *subtract* the diff. lat. from the *first* true altitude.

(c) When the angle found above is 0 points, *add the whole distance run in the interval* to the *first* true altitude.

(d) When the angle found above is exactly 16 points, *subtract the whole distance run in the interval* from the *first* true altitude.

(e) When the angle found above is exactly 8 points, there will be no correction.

Note 1.—The *tenths* of miles found in the Traverse Table can be turned into *seconds*, by multiplying them by 6.

Note 2.—By the *greater* bearing is meant the bearing of the body at the moment of observing the *lesser* altitude of the two; or its bearing when farthest from the meridian (Noon).

Note 3.—By the *less* bearing is meant the bearing of the body at the moment of observing the *greater* altitude of the two; or its bearing when nearest to the meridian (Noon).

Note 4.—When the bearings lie on opposite sides of the meridian, and are equal (that is, the same distance from N. or from S.), either may be taken as the less bearing.

TO COMPUTE THE LATITUDE.

To find *Arc* (1).

Add together

the log. sin. of the pol. dist. at G.B.

log. sin. of the pol. dist. at L.B.

log. havers. of the polar angle (in time),

reject the tens in the index of the sum; with this result as a log. *havervsine*, take out the arc in degrees, and mark it (*a*), which will be arc (1) when the difference of the polar distances is nothing, or *very small*. (See *Note*.)

Note.—When the diff. polar distances exceed 40', to the nat. versine (Tab. v. Inman) of arc (*a*) add the nat. versine of diff. pol. dists. With the sum as a nat. versine take out the arc and mark it (1).

To find *Arc* (2).

Under arc (1) place the pol. dist. at *greater* bearing, and take

their difference, under which diff. place the pol. dist. at *less* bearing, take the sum and diff.

Add together

the log. cosec. of arc. (1),
 log. cosec. of P.D.G.B.,
 $\frac{1}{2}$ log. havers. of the sum,
 $\frac{1}{2}$ log. havers. of the diff.,

reject the tens in the index of the sum of these four logs. and take out the result as a *haversine*, in degrees. Call this arc (2).

To find *Arc* (3).

Again, under arc (1) place the zen. dist. at *greater* bearing, and take their difference, under which difference place the zen. dist. at *less* bearing, take the sum and diff.

Add together

the log. cosec. of arc (1),
 log. cosec. of Z.D.G.B.,
 $\frac{1}{2}$ log. havers. of the sum,
 $\frac{1}{2}$ log. havers. of the diff.,

reject the tens in the index of the sum of these four logs. and take out the result as a log. *haversine*, in degrees. Call this arc (3).

To find *Arc* (4).

Take the *difference* between arcs (2) and (3), for arc (4), unless the distance arc joining the places of the sun at the two observations passes *between* the zenith and pole; in which case alone take the *sum* of arcs (2) and (3) for arc (4).

To find the *Latitude*.

Add together

the log. sin. of the P.D.G.B.,
 log. sin. of the Z.D.G.B.,
 log. havers. of arc (4),

reject the tens in the index of the sum, and take out the log. *haversine* of the remaining log. in degrees. Call this arc (5).

To the natural versine (Tab. v. Inman) of this arc (5), *add* the natural versine of the difference between the polar and zenith distances at the *greater* bearing. The sum will be natural versine of the co-latitude, which subtracted from 90° will give the latitude required.

Note.—When *near the equator* the arc may be greater than 90° , in which case subtract 90° from it, the remainder will then be the latitude of a *contrary* name to the latitude by *account*.

EXAMPLE.

September 24th, in lat. D.R. $10^{\circ} 25' N.$ and long. $168^{\circ} 15' E.$, the following observations were taken to find the latitude:—

M. T. Ship.	Chron. Times.	Obs. alt. Sun's L.L.	True Bearings.
11h. 15m. A.M.	oh. 5m. 15s.	$75^{\circ} 59' 20''$	S. $39^{\circ} E.$
3 10 P.M.	3 56 30	39 27 30	S. $82 W.$

Index error $+1' 30''$, height of the eye 18 feet; run of the ship in the interval S. $30^{\circ} E.$, 29 miles.

(1) Green. date at Gr. Bg.	(4a) Polar \angle (in Time)	(1) Green. date at Lr. Bg.																																										
<table border="0"> <tr><td>Time at ship</td><td><u>h. m.</u></td></tr> <tr><td></td><td>3 10 Sept. 24th</td></tr> <tr><td></td><td>24</td></tr> <tr><td>Ast. sh. T.</td><td>27 10 Sept. 23rd</td></tr> <tr><td>Long. in time</td><td>11 13 E.</td></tr> <tr><td>Gr. date</td><td>15 57 Sept. 23rd</td></tr> </table>	Time at ship	<u>h. m.</u>		3 10 Sept. 24th		24	Ast. sh. T.	27 10 Sept. 23rd	Long. in time	11 13 E.	Gr. date	15 57 Sept. 23rd	<table border="0"> <tr><td>Chronometer Times.</td></tr> <tr><td><u>h. m. s.</u></td></tr> <tr><td>0 5 15</td></tr> <tr><td>3 56 30</td></tr> <tr><td><u> </u></td></tr> <tr><td>3 51 15</td></tr> <tr><td>Polar \angle (in Time)*</td></tr> </table>	Chronometer Times.	<u>h. m. s.</u>	0 5 15	3 56 30	<u> </u>	3 51 15	Polar \angle (in Time)*	<table border="0"> <tr><td>Time at ship</td><td><u>h. m.</u></td></tr> <tr><td></td><td>11 15 A.M. Sept. 24th</td></tr> <tr><td></td><td>12</td></tr> <tr><td>Ast. sh. T.</td><td>23 15 Sept. 23rd</td></tr> <tr><td>Long. in time</td><td>11 13 E.</td></tr> <tr><td>Gr. date</td><td>12 2 Sept. 23rd</td></tr> </table>	Time at ship	<u>h. m.</u>		11 15 A.M. Sept. 24th		12	Ast. sh. T.	23 15 Sept. 23rd	Long. in time	11 13 E.	Gr. date	12 2 Sept. 23rd											
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(2) Declin. at Gr. Bg.	Change.	(2) Declin. at Lr. Bg.	Change.																																									
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* This is not necessarily the mean time at Greenwich, but the time shown by a good watch or chronometer, in order to obtain the polar angle as accurately as possible.

To find the Zenith Distances.

(5) *First obs. alt. (L.B.)*

Obs. alt.	79° 59' 20"
Index error	+ 1 30
<hr/>	
Dip	76 0 50
	- 4 11
<hr/>	
Ref.	75 56 39
	- 0 12
<hr/>	
Semid.	75 56 27
	+ 15 59
<hr/>	
Tr. alt. centre	76 12 26
(7a) {	Angle between bearing of sun at first obs. (L.B.) and run of ship, $\frac{1}{2}$ pts., dist. 29m.
<hr/>	
True alt.	71 41 8
	90
<hr/>	
(6) Zen. dist.	13 18 52

(5) *Second obs. alt. (G.B.)*

Obs. alt.	39° 27' 30"
Index error	+ 1 30
<hr/>	
Dip	39 29 0
	- 4 11
<hr/>	
Ref.	39 24 49
	- 1 4
<hr/>	
Semid.	39 23 45
	+ 15 59
<hr/>	
True alt. centre	39 39 44
	90
<hr/>	
Zen. dist.	50 20 16

To find Arc (1).

P.D.G.B.	90° 20' 38"	Sin.	9'999992
P.D.L.B.	90 16 44	Sin.	9'999995
Pol. angle	3h. 51m. 15s.	Hav.	9'368574

Havers. arc (a) 9'368561

Arc (1) 57° 48' 45".

To find Arc (2).

Arc (1)	57° 48' 45"	Cosec.	0'072471
P.D.G.B.	90 20 38	Cosec.	0'000008
<hr/>			
Diff.	32 31 53		
P.D.L.B.	90 16 44		
<hr/>			
Sum	122 48 37 $\frac{1}{2}$	Hav.	4'943503
Diff.	57 44 51 $\frac{1}{2}$	Hav.	4'683929
<hr/>			
Havers.	9'699811		
<hr/>			
Arc (2)	90° 6' 45".		

To find Arc (3).

Arc (1)	57° 48' 45"	Cosec.	0'072471
Z.D.G.B.	50 20 16	Cosec.	0'113612
<hr/>			
Diff.	7 28 29		
Z.D.L.B.	13 18 52		
<hr/>			
Sum	20 47 21 $\frac{1}{2}$	Hav.	4'256265
Diff.	5 50 23 $\frac{1}{2}$	Hav.	3'707196
<hr/>			
Havers.	8'149544		
<hr/>			
Arc (3)	13° 38' 30".		

To find the Latitude.

Arc (2)	90° 6' 45"		
Arc (3)	13 38 30		
<hr/>			
Arc (4)	76 28 15	Hav.	9'583233
<hr/>			
P.D.G.B.	90 20 38	Sin.	9'999992
Z.D.G.B.	50 20 16	Sin.	9'886388
<hr/>			
Diff.	40 0 22	Hav.	9'469613
<hr/>			
Arc (5)	65° 46' 45".		

Parts for 45"	199
Vers. Arc (5)	0'589546
Vers. diff. P. and Z.D.	0'233956
Parts for 22"	69
<hr/>	
Versin. (sum)	0'823770
<hr/>	
Arc	79° 51' 0" (co-latitude)
	90
<hr/>	
Latitude	10 9 0 N.

DOUBLE ALTITUDE,*By two altitudes of different bodies observed at the same time.*

(TWO STARS.)

RULE.

1. TAKE from the N. A. the declinations of the two bodies, and find the polar distances, as in (a) or (b) of 3, Rule for Double Altitude, and take their difference.

2. Also take from the N.A. the right ascensions of the two bodies, their difference (or if greater than 12h. what it wants of 24h.) will be the *polar angle* in time.

3. Correct the two observed altitudes, and get the zenith distances.

TO COMPUTE THE LATITUDE.

4. To find Arc (1).

Add together

the log. sin. of the pol. dist. at G.B.,
 log. sin. of the pol. dist. at L.B.,
 log. havens. of the polar angle (in time),

reject the tens in the index of the sum ; with this result as a log. haversine, take out the arc in degrees, and mark it (a). To the nat. versin. of (a) add the nat. versin. of the diff. pol. dist. With the sum as a nat. versin. take out the arc and mark it (1).

The remainder of the rule is the same as for the sun.

Note.—In finding arc (4), it will sometimes happen that the sum of arcs (2) and (3) has to be taken. This will be the case only when the distance arc joining the places of the two bodies passes *between* the zenith and pole. If the sum of (2) and (3) be greater than 180° , the supplement to 360° must be considered as the sum.

Arcs (1) and (2) may be taken by inspection from Rear-Admiral Shadwell's Star Tables, which afford a very easy and expeditious method of finding them.

**TO FIND THE TIME OF THE MOON'S PASSAGE
OVER A GIVEN MERIDIAN.**

RULES.

In West Longitude.

(A) TAKE from the Nautical Almanac (page IV. of the month) the moon's meridian passage on the given day; underneath which place the moon's passage on the *following* day, and take their difference. (Diff. moon's mer. pass.)

Then enter Table k. (Inman) with the diff. moon's mer. pass on the left hand of the page to *nearest minute*, and longitude of ship to *nearest degree* at the top, and take out the *correction*, which is to be added to the *first* passage; the sum will be the time of moon's passage at *ship*, of the day of first passage.

In East Longitude.

(B) Take from the Nautical Almanac (page IV. of the month) the moon's meridian passage for the given day; underneath which place the moon's passage on the *preceding* day, and take their difference. (Diff. moon's mer. pass.)

Then enter Table k. (Inman) with this diff. moon's mer. pass. on the left hand of the page to *nearest minute*, and longitude of ship to *nearest degree* at the top, and take out the *correction*, which *subtract* from the *first* passage; the remainder will be the time of moon's passage at *ship*, of the day of first passage.

Note 1.—In *west* longitude, should the sum of the moon's mer. pass. (from N.A.) and the *correction* from Table k. (together) exceed 24h., reject 24h., and put the *date* one day forward.

Note 2.—In *east* longitude, should the *correction* from Table k. be more than the time of the moon's mer. pass. (from N.A.), add 24h. to the passage before subtracting the *correction*, and put the *date* one day back.

EXAMPLE (A).—Longitude *West*.

On Sept. 19th, in long. 60° W., find the moon's meridian passage at ship.

Moon's passage at Gr. 19th Sept. (given day)	7h. 58m. (p. IV. N.A.)	
Ditto 20th Sept. (day following)	8 48	

Diff. moon's mer. pass.	0 50	
Enter Tab. k. with diff. pass. 50m. (taken as 48m.) at side, and long. 60° at top, gives	} + 8	{ added to first mer. pass.
<i>correction</i>	8 6	
Moon's meridian passage at <i>ship</i>		Sept. 19th

EXAMPLE (B).—Longitude *East*.

Find the moon's mer. pass. at a ship. in long. 79° E., on Sept. 20th.

Moon's passage at Gr. 20th Sept. (given day)	8h. 48m. (p. IV. N.A.)	
Ditto 19th Sept. (preceding day)	7 58	

Diff. moon's mer. pass.	0 50	
Enter Tab. k. with diff. pass. 50m. (taken as 48m.) at side, and long. 79° (taken as 80°) at top, which gives the	} - 11	{ subtracted from first mer. pass.
<i>correction</i>	8 37	
Moon's mer. pass. at <i>ship</i>		Sept. 20th

TO CORRECT THE MOON'S SEMIDIAMETER AND HORIZONTAL PARALLAX.

RULES.

1. Having found the Greenwich date, take out the moon's semidiameter from page III. of the Nautical Almanac, as follows:—

(A) When the Greenwich date is *less* than 12h., take out the semid. for *noon* of the day of the Greenwich date; underneath which place the semid. for *midnight* of the *same* day, and take their difference.

(B) When the Greenwich date is *more* than 12h., take out the semid. for *midnight* of the day of the Greenwich date; underneath which place the semid. for *noon* of the *following* day, and take their difference.

2. Then multiply this difference by the *hours* of the Greenwich date, and divide the product by 12; the quotient will be the *correction*.

Note.—If the G.D. exceeds 12h., use what it is *above* 12 hours.

3. If the semid. is *increasing*, add the correction to the semid. *first* taken out, the sum will be the semidiameter required.

4. If the semid. is *decreasing*, subtract the correction from the semid. *first* taken out; the remainder will be the semidiameter required.

Note.—The above rules are applied in precisely the same manner when correcting the *horizontal parallax*.

EXAMPLE (I).

Find the moon's semidiameter and horizontal parallax on Sept. 16th, at 9h. M.T.G.

Semid.	Change.	Hor. Par.	Change.
Semid. <i>noon</i> 16th Sept. $14^{\circ} 49' 5''$	$2'' 1$	Hor. Par. <i>noon</i> 16th Sept. $54^{\circ} 18' 7''$	$7'' 7$
Ditto <i>mid.</i> 16th Sept. $14^{\circ} 51' 6''$	\times by 9	Ditto <i>mid.</i> 16th Sept. $54^{\circ} 26' 4''$	\times by 9
Difference in 12h. $2' 1$	$12) 18' 9$	Difference in 12h. $7' 7$	$12) 69' 3$
Add to semid. at <i>noon</i> $+ 1' 6$	Corr. $1' 6$	Add to H.P. at <i>noon</i> $+ 5' 8$	Corr. $5' 8$
Semidiameter $14^{\circ} 51' 1''$	nearly	Hor. Par. $54^{\circ} 24' 5''$	nearly

EXAMPLE (II).

Green. date *more* than 12 hours.

Find the moon's semidiameter and horizontal parallax on Sept. 19th, at 15h. M.T.G.

Semid.	Change.	Hor. Par.	Change.
Semid. <i>mid.</i> 19th Sept. $15^{\circ} 18' 0''$	$6'' 2$	Hor. Par. <i>mid.</i> 19th Sept. $56^{\circ} 3' 3''$	$22' 9$
Semid. <i>noon</i> 20th Sept. $15^{\circ} 24' 2''$	\times by 3*	Hor. Par. <i>noon</i> 20th Sept. $55^{\circ} 26' 2''$	\times by 3*
Diff. in 12h. $6' 2$	$12) 18' 6$	Difference in 12h. $22' 9$	$12) 68' 7$
Add to semid. at <i>mid.</i> $1' 5$	Corr. $1' 5$	Add. to H.P. at <i>mid.</i> $5' 7$	Corr. $5' 7$
Semidiameter $15^{\circ} 19' 5''$	nearly	Hor. Par. $55^{\circ} 9' 0''$	nearly
* Here, the diff. in 12h. is multiplied by 3, or what 15h. (the G.D.) is over 12h. (see B).		* Here, the diff. in 12h. is multiplied by 3 (see B).	

TO CORRECT THE MOON'S DECLINATION.

RULE.

1. Find the time of the moon's passage at ship (*see* Rules, p. 87), and with the longitude in time, get the Greenwich date.

2. Take out the moon's declination from the N.A. for the *hours* of the Greenwich date; also take out from the adjoining column the diff. declin. for 10 *minutes*.

3. Multiply the diff. declin. by the *minutes* of the Greenwich date, divide the product by 10 and then by 60; the result will be the *correction* in minutes and seconds.

Or, remove the decimal point in the diff. dec. one place to the *left*, which will give the diff. dec. for *one* minute; then multiply by the *minutes* of the Greenwich date, and divide the product by 60; the result will be the *correction* in minutes and seconds.

4. When the declin. is *increasing*, add the correction, which will give the declin. required.

5. When the declin. is *decreasing*, subtract the correction; the remainder will be the declin. required.

Note 1.—When the correction is *subtractive* and is more than the declin. taken out, the difference between the correction and declination will give the declin. required, but which is to be marked with a *contrary* name to the declin. taken out.

EXAMPLE I.

Find the moon's declin. on Sept. 15th, at 5h. 35m. M.T.G.

(2) Declin.	(3) Diff. dec. in rom.
Declin. Mar. 8th, at 5h. $16^{\circ} 47' 5''S.$ Corr. for 35m. $+ 2 7$ <hr style="width: 50%; margin-left: auto; margin-right: 0;"/> Declination $16 49 12 S.$	Diff. dec. in rom. $36'' \cdot 44$ x by minutes of Green. date $\underline{35}$ 18280 $\underline{10932}$ $6,0)127,5 \cdot 40$ Corr. $2' 7''$

EXAMPLE II.

Find the moon's declin. on Sept. 23rd, at 22h. 52m. M.T.G.

(2) Declin.	(3) Diff. dec. in rom.
Declin. Sept. 23rd, at 22h. $0^{\circ} 7' 38''S.$ Corr. for 52m. (<i>Note 1</i>) $0 9 42$ <hr style="width: 50%; margin-left: auto; margin-right: 0;"/> Declination $0 2 4 N.$	Diff. dec. $111'' \cdot 84$ } decimal point in rom. $11' 184$ } removed one x by minutes } $\underline{52}$ } place to left. of Gr. date. } 22368 $\underline{55920}$ $6,0)58,1 \cdot 568$ Corr. $9' 42''$

MOON'S RIGHT ASCENSION.

To correct the moon's right ascension, take out the R.A. for the *hour* of the Green. date, and find the increase of R.A. between that hour and the next following. Then ascertain the proportion of increase for the *minutes* of the G.D., and *add* it to the R.A. taken out, which will be the R.A. required.

TO FIND THE LATITUDE BY THE MERIDIAN ALTITUDE OF THE MOON.

RULE.

1. Find the time of the moon's meridian passage at *ship* (by Rule, page 87), and with the given longitude get the Greenwich date.

2. For this date take from the Nautical Almanac the moon's declination, moon's semidiameter, and horizontal parallax, and correct them.

3. Then correct the observed altitude for index error, dip, and (augmented) semid.,* and get the apparent altitude.

4. Enter Table w. (Inman) with the degrees and minutes of *apparent altitude* at the side, and minutes of moon's hor. par. at the top, and take out the correction from the column marked $\left| \begin{array}{c} \text{corr.} \\ + \end{array} \right|$ and place it under the apparent altitude.

Note.—If the horizontal parallax is *less* than 54' (the limit of Table w.), take out the correction for 54', and also the parts for the seconds which the hor. par. is *less* than 54'; then *subtract* these parts from the "corr." given in the table, the remainder will be the correction required.

5. Then from the small table at the right-hand side of the same page, with the *seconds* of hor. par. in the column marked ("), take out the seconds given in the adjoining column marked *corr.*, which place under the corr. obtained in (4).

Add both these quantities to the app. alt., the sum will be the true altitude.

6. Subtract the true alt. from 90° for the zenith distance (Z.D.). Mark the Z.D. N. or S., according as the observer is N. or S. of the moon.

Under the Z.D. place the declin. also marked N. or S.

7. When the Z.D. and declin. are of the *same* name, their sum will be the latitude, of the name of either.

8. When the Z.D. and declin. are of contrary names, take the less from the greater; the difference will be the latitude, of the name of the *greater*.

* The *augmentation* of the moon's semid. is taken out from Table g, Inman, with alt. at side, and semid. at top.

Ex. to Note.—Suppose the app. alt. to be $30^\circ 20'$, and the hor. par. $53' 55''$. The "corr." corresponding to 54' of hor. par. and app. alt. $30^\circ 20'$ is $44' 58''$ and the parts (in the small table) for the 5'', by which the hor. par. is *less* than 54' are 4, these *subtracted* from $44' 58''$ give $44' 54''$, the correction required.

EXAMPLE.

September 15th, in long. 27° W., the obs. mer. alt. of moon's L.L. was 54° 15' 30" (Z.N.), index error - 3' 30", height of eye 19 feet; required the latitude.

(1) Gr. Date from pass. at Ship.	(2) Declination.	Change in 10m.
Moon's pass. Sept. 15th	Declin., 15th, at 6h. 16° 50' 42" S.	35"-68
Ditto, Sept. 16th	Corr. for 36m. 2 8	x by min. } 36
Diff. pass.	Declin. 16 52 50 S.	21408
Enter Tab. k. with diff. pass. 48, and long. 25°, gives		10704
Moon's pass. at ship		6,0)12,84.48
Long. in time		Corr. for } 2' 8"
Gr. date, Sept. 15th		

(2) Semid. p. III., N.A.	Change.	(2) Hor. par. p. III., N.A.	Change.
Semid. noon, 15th		Hor. par. noon, 15th	2'5
Ditto, mid., 15th	x by hours of } 7	Ditto, mid., 15th	x by min. of } 6'5
Diff. 12h.	G.D. } 6'5	Diff. in 12h.	2'5
Corr. for 6h. 36m.	12)4'55	Corr. for 6h. 36m.	+ 1'4
Semid.	Corr. 0'4	Hor. par.	54 12'4
Augmentation alt. 54° + 11'3			125
A.S.D.			150
			12)16'25
			Corr. 1'4

To find the Latitude.

(3) Obs. mer. alt.	54° 15' 30"
Index error	- 3 30
Dip	54 12 0
Aug. S. D.	- 4 17
App. alt. centre	54 7 43
With app. alt. 54° 22'	+ 14 59
(4) at side, and hor. par.	54 22 42
54 at top, gives	+ 30 46
(5) Small table, with 59'	
of hor. par. under	+ 7
Corr. gives	54 53 35
True alt.	90
(6) Zen. dist.	35 6 25 N.
Declin.	16 52 50 S.
(8) Latitude	18 13 35 N.

LONGITUDE BY LUNAR. (SUN.)

RULE.

1. With the mean time at ship and longitude get a Greenwich date.

2. For this date take out of the Nautical Almanac (page II. of the month) the sun's declin., the equation of time, and the sun's semid.

Also, from page III. of the month, take out the moon's semid. and horizontal parallax.

3. Correct the sun's declin. and the equation of time.

4. Correct the moon's semid. and hor. parallax.

5. Correct the *sun's* observed altitude

for $\left\{ \begin{array}{l} \text{Index error,} \\ \text{dip,} \\ \text{semid.} \end{array} \right\}$ and get the app. alt. (A. A.)

then apply the refraction (Tab. m.), and find the true alt. : also get the zenith dist. (Z.D.).

6. Correct the *moon's* observed altitude

for $\left\{ \begin{array}{l} \text{Index error,} \\ \text{dip,} \\ \text{augmented semid.} \end{array} \right\}$ and get the app. alt. (A. A.)

then apply the correction in alt. (Tab. w.), and find the true alt. : also get the zenith dist. (Z.D.)

7. To take out the auxiliary angle A. (Tab. w.)

Whilst Tab. w. is being used for the corr. in alt. look in the *adjoining* column to the *right* of that marked $\left| \begin{array}{c} \text{corr.} \\ + \end{array} \right|$, and under $\left| \begin{array}{c} A \\ 60^\circ \end{array} \right|$ in the *same* line with the "corr." just taken out, will be found the minutes and seconds of the auxiliary angle A.

Then, in the small table on the right-hand side of the page, and with the *seconds* of hor. par. take out the seconds in the column under $| A |$.

Next, go to the small table at the bottom of the page, and with the *nearest* degree to the app. alt. of the *sun*, take out the seconds given in the column under $\mid \odot \mid$.

Add these three quantities to 60° , which will be the auxiliary angle A required.

8. Correct the observed distance for index error (if any). To this result *add* the sun's semid. and also the moon's *augmented* semid. ; the sum will be the apparent distance (A.D.).

9. To compute the **true distance**.

Add together the two true zenith distances (Z.D.), and mark the sum *vers*.

Add together the two apparent altitudes (A.A.), under this sum put the auxiliary angle A., take the sum and difference, and mark each *vers*.

Under the app. dist. (A.D.) put the auxiliary angle (aux. \angle), take the sum and difference, and mark each *vers*.

10. From Table v. (Inman), take out the nat. versine of the degrees and minutes of each of the above arcs (marked *vers*.), using, however, only the right-hand *five* figures of each versine : at the same time take from the opposite page of the table the parts for *seconds* of each arc.

Note.—If any arc marked *vers*. be greater than 180° , take the *vers*. of what it wants of 360° .

11. Place these versines in one column, and their parts in another. Add up each column separately, and then make one sum of both columns, rejecting the tens in the fifth place to the *left*.

12. Then look in the column of versines marked with the *degree* of the *apparent* distance, or in the column preceding or following that degree, when the five figures in the above sum, or the next *less* to them, will give the degrees and minutes of the *true* distance (T.D.).

13. If there is any difference between the figures of the versine found from the table and those sought for, the *seconds* corresponding to this difference must be taken from the opposite page ("Parts for Seconds"), and added to the degrees and minutes of the true distance.

To compute the mean time at Greenwich.

14. Find in the Nautical Almanac, on the *given* day, two distances of the sun and moon *between* which the true distance lies—that is, two distances, one of which is *greater* and the other *less* than the true distance just found.

Under the true distance (T.D.), place that distance which comes *first* in order in the Nautical Almanac, which mark with the *hour* denoted at the head of its column; and at the same time take out the proportional logarithm | $\begin{array}{c} \text{P.L.} \\ \text{of} \\ \text{diff.} \end{array}$ | given in the *adjoining* column on the right.

Take the diff. between the T.D. and the one given in the Nautical Almanac, and find its corresponding proportional logarithm from Table r. (Inman).

15. Then take the less prop. log. from the greater: the result will be the prop. log. of a portion of time, to be added to the *hour* denoted at the head of the column of distances in the N.A.; the sum will be the *mean time at Greenwich* (expressed astronomically), which mark with the day given in the Nautical Almanac.

Note.—If great accuracy be required, such as when fixing the longitude of a place on shore, the moon's hor. par. and the latitude of the place should be corrected for the spheroidal figure of the Earth, by Table h. (Inman).

16. To compute the mean time at ship.

This is done precisely as in working a *sun* chronometer. (See 7 of that Rule.)

Note.—The sun is here the body from which the ship time is deduced.

Take the difference between the Greenwich mean time (obtained as above) and the ship mean time (reducing them, if necessary, to the *same* date by adding 24h. to the least of the two, and putting it back a day).

17. This difference will be the longitude in time, which, when turned into degrees, &c., will give the longitude of the ship.

18. (a) *East*, when the Greenwich time is the *less* of the two.
 (b) *West*, when the Greenwich time is the *greater* of the two.

EXAMPLE.

September 15th, at 3h. 20m. P.M. mean time, in lat. 35° N. and long. D.R. $25^{\circ} 30' W.$, the following lunar was taken :—

Obs. alt. sun's L.L. $32^{\circ} 39' 50''$	Obs. alt. moon's L.L. $33^{\circ} 25' 20''$	Obs. dist. $74^{\circ} 11' 10''$
Index error — 1 25	+ 1 40	+ 1 5

Height of eye 24 feet ; required the longitude.

(1) Green. date.	(3) Sun's declin. p. II., N.A.	(3) Change.	
Time at ship h. m. 3 20 Sept. 15th	Declin., 15th $\overset{\circ}{3} \overset{\prime}{1} \overset{\prime\prime}{19}$ N.	x G.D. $\overset{\prime\prime}{57} \overset{\prime}{81}$ <u>5</u>	
Long. in time 1 42 W.	Corr. —4 49		
Gr. date 5 2 Sept. 15th	Declin. 2 56 30 N.	$60)28,905$ Corr. $\overset{\prime}{4} \overset{\prime\prime}{49}$	
Sun's semid.	(3) Eq. time, p. II., N.A.	Change.	
Semid. $\overset{\prime}{15} \overset{\prime\prime}{57}$	Eq. time, 15th m. s. 4 50 0	x G.D. $\overset{\prime\prime}{879}$ <u>5</u>	
	Corr. +4 39		
	Eq. time 4 54 39	Corr. $\overset{\prime\prime}{4} \overset{\prime}{395}$	
	<i>Sub. from app. time.</i>		
(4) Moon's semid. p. III., N.A.	Change.	(4) Hor. par. p. III., N.A.	Change.
Semid. noon, 15th $\overset{\prime}{14} \overset{\prime\prime}{47} \overset{\prime\prime\prime}{3}$	x 5 <u>7</u>	Hor. par. noon, 15th $\overset{\prime}{54} \overset{\prime\prime}{11}$	x 5 <u>2 5</u>
Ditto, mid., 15th $\overset{\prime}{14} \overset{\prime\prime}{48} \overset{\prime\prime\prime}{0}$		Ditto, mid., 15th $\overset{\prime}{54} \overset{\prime\prime}{13} \overset{\prime\prime\prime}{5}$	
Diff. in 12h. 7	$12)3,5$ —	Diff. in 12h. 2 5	$12)12 5$ —
Corr. for 5h. + 3	Corr. 3	Corr. for 5h. + 1 0	Corr. $\overset{\prime}{1} \overset{\prime\prime}{0}$
Semid. $\overset{\prime}{14} \overset{\prime\prime}{47} \overset{\prime\prime\prime}{6}$		Hor. par. $\overset{\prime}{54} \overset{\prime\prime}{12}$	
Augmentation, }			
Tab. g. alt. 32° + 7 5			
Aug. semid. $\overset{\prime}{14} \overset{\prime\prime}{55} \overset{\prime\prime\prime}{1}$			

(3) To find Sun's Z.D.		(6) To find Moon's Z.D.		(8) To find App. Dist.	
Obs. alt.	32° 39' 50"	Obs. alt.	33° 25' 20"	Obs. dist.	74° 11' 10"
Ind. error	- 1 25	Ind. error	+ 1 40	Ind. error	+ 1 5
	<hr/>		<hr/>		<hr/>
Dip	32 38 25	Dip	33 27 0		74 12 15
	- 4 49		- 4 49	☉ Semid.	+ 15 57
	<hr/>		<hr/>	☽'s A.S.D.	+ 14 55
Semid.	32 33 36	Aug. S.D.	33 22 11	A.D.	74 43 7
	+ 15 57		+ 14 55		<hr/>
A.A.	32 49 33	A.A.	33 37 6	(7) Auxiliary angle A.	
Ref.	- 1 23	Corr.	+ 42 32		60° 0' 0"
	<hr/>	Tab. w. }	+ 0 10	Corr. Tab. w. }	+ 16 17
True alt.	32 48 10	True alt.	34 19 48	Seconds of }	+ 0 4
	90		90	par. }	+ 0 3
	<hr/>		<hr/>		<hr/>
Zen. dist.	57 11 50	Zen. dist.	55 40 12	Aug. angle	60 16 24
	<hr/>		<hr/>		<hr/>

(9) To compute the True Distance.
 Sun's Z.D. 57° 11' 50" (10) (11)
 Moon's Z.D. 55 40 12 (Parts for seconds.)

Sum 112 52 2 Vers. 88588 ... 13
 Sun's A.A. 32° 49' 33"
 Moon's A.A. 33 37 6

Sum 66 26 39
 Aux. angle 60 16 24

Sum 126 43 3 Vers. 97858 ... 12
 Diff. 6 10 15 Vers. 05786 ... 8

A. dist. 74° 43' 7"
 Aux. angle 60 16 24

Sum 134 59 31 Vers. 06901 ... 107
 Diff. 14 26 43 Vers. 31562 ... 50

30695 190
 Parts 190

Vers. 30885

(12) 30880 gives 74° 23' 0"
 85 parts give 18 (see 13)

True distance 74 23 18
 H

(14) To compute the Mean Time at Greenwich.

True dist.	74° 23' 18"	
{ Dist. at III. hours }	73 26 55	P.L. 34480
{ from Naut. Almc. }		
Diff.	0 56 23	P.L. 50412
(Table r. Inman)	P.L.	Diff. 15932

	h.	m.	s.	
	2	4	43	
Hour in Naut. Almc.	3	0	0	
M.T. Green.	5	4	43	Sept. 15th (see 15.)

(16) To compute the Mean Time at Ship.

Lat.	35° 0' 0" N.	Sec. 0°086636
Dec.	2 56 30 N.	Sec. 0°000573
Diff.	32 3 30	
Z. D.	57 11 50	
Sum	89 15 20	½ Hav. 4°846640
Diff.	25 8 20	½ Hav. 4°337680
		Havers. 9°271529

	h.	m.	s.	
App. time at ship	3	24	54	
Eq. time	—	4	54	
Mean time at ship	3	20	0	Sept. 15th
Mean time at Green.	5	4	43	Sept. 15th
(17) Longitude in time	1	44	43	
	Longitude 26° 10' 45" W. (18. b.)			

LONGITUDE BY LUNAR.

(STAR, OR PLANET, AND MOON.)

1. Get a Greenwich date as in rule for Sun Lunar.
2. For this date take from the N.A. the moon's semid. and hor. par., and correct them.
3. Take out the R.A. mean sun, and correct it.
4. Take out the R. Asc. and declin. of the body from whose altitude the mean time at ship is to be deduced, and, when necessary, correct them.

5. From the observed altitudes get the zenith distances.

Note.—When a planet is observed, take out its hor. par. from the N.A. and *add* the *parallax in altitude* (found in Tab. b., Inman) to the *app.* alt. after it has been corrected for refraction, and before finding the Z.D.

6. To the observed distance corrected for index error, *add* the moon's augmented semid. when the limb of the moon observed is *nearest* to the star (N.L.); but *subtract* it from the distance when the limb observed is *furthest* from the star (F.L.). The result will be the apparent distance (A.D.).

7. Take out the auxiliary angle from Tab. (w.) Inman, in the same manner as directed in (7) for Sun Lunar, p. 93; but in using the small table at the bottom, take out the corr. given in the column under | * | instead of that under | ⊙ |.

Note.—If a planet be observed, take the second corr. for A from Tab. (c.) Inman, instead of using the small table for | ⊙ | or | * |.

8. Compute the **true distance** by the Rule for Sun Lunar, see (9) to (13), p. 94.

NAUTICAL ASTRONOMY.

EXAMPLES.

TO FIND THE LATITUDE BY THE MERIDIAN ALTITUDE OF A STAR.

1. April 1st, the obs. mer. alt. of *Regulus* was $64^{\circ} 10'$ (Z.N.), index error $-1' 40''$, height of the eye 14 feet; required the latitude.

Ans. Latitude $38^{\circ} 31' N$.

2. May 10th, the obs. mer. alt. of *Spica* was $59^{\circ} 15'$ (Z.S.), index error $+2'$, height of the eye 20 feet; required the latitude.

Ans. Latitude $41^{\circ} 16' S$.

3. June 12th, the obs. mer. alt. of *Arcturus* was $50^{\circ} 10'$ (Z.S.), index error $-2' 20''$, height of the eye 18 feet; required the latitude.

Ans. Latitude $20^{\circ} 4' S$.

4. June 20th, the obs. mer. alt. of *Aldebaran* was $70^{\circ} 5'$ (Z.S.), index error $+1' 10''$, height of the eye 24 feet; required the latitude.

Ans. Latitude $3^{\circ} 45' S$.

5. July 15th, the obs. mer. alt. of *Castor* was $51^{\circ} 25'$ (Z.S.), index error $-2' 40''$, height of the eye 17 feet; required the latitude.

Ans. Latitude $6^{\circ} 32' S$.

6. August 27th, the obs. mer. alt. of *Vega* was $20^{\circ} 45'$ (Z.S.), index error $+1' 25''$, height of the eye 22 feet; required the latitude.

Ans. Latitude $30^{\circ} 41' S$.

LATITUDE BY MERIDIAN ALTITUDE OF A STAR
UNDER THE POLE.

1. January 2nd, the obs. mer. alt. of *Achernar* under the *South Pole* was $20^{\circ} 10'$, I.E. $-3'$, height of the eye 14 feet; required the latitude.

Ans. Latitude $52^{\circ} 5' S$.

2. June 10th, the obs. mer. alt. of *Canopus* under the *S. Pole* was $22^{\circ} 14'$, I.E. $+1' 40''$, height of the eye 18 feet; required the latitude.

Ans. Latitude $59^{\circ} 31' S$.

3. July 20th, the obs. mer. alt. of η *Argus* under the *S. Pole* was $30^{\circ} 12'$ I.E. $-2' 50''$, height of the eye 20 feet; required the latitude.

Ans. Latitude $61^{\circ} 4' S$.

4. August 14th, the obs. mer. alt. of α *Crucis* under the *S. Pole* was $32^{\circ} 22'$, I.E. $+1'$, height of the eye 17 feet; required the latitude.

Ans. Latitude $59^{\circ} 56' S$.

5. October 18th, the obs. mer. alt. of α *Ursæ Majoris* under the *N. Pole* was $18^{\circ} 50'$, I.E. $-1' 45''$, height of the eye 22 feet; required the latitude.

Ans. Latitude $46^{\circ} 13' N$.

6. November 14th, the obs. mer. alt. of α *Persei* under the *N. Pole* was $23^{\circ} 55'$, I.E. $+2' 30''$, height of the eye 16 feet; required the latitude.

Ans. Latitude $64^{\circ} 28' N$.

LATITUDE BY MERIDIAN ALTITUDE OF THE SUN.

1. September 20th, in long. $42^{\circ} W$., the obs. mer. alt. of the sun's L.L. was $50^{\circ} 13' 10''$ (Z.N.), I.E. $-2' 10''$, height of the eye 26 feet; required the latitude.

Declin. $1^{\circ} 2' 28'' N$.

Latitude $40^{\circ} 41' 14'' N$.

2. September 21st, in long. $70^{\circ} 30' E$., the obs. mer. alt. of the sun's L.L. was $60^{\circ} 15' 20''$ (Z.S.), I.E. 0, height of the eye 23 feet; required the latitude.

Declin. $0^{\circ} 46' 25'' N$.

Latitude $28^{\circ} 47' 29'' S$.

3. September 22nd, in long. $43^{\circ} 30'$ E., the obs. mer. alt. of the sun's L.L., by *artificial horizon*,* was $100^{\circ} 14' 20''$ (Z.N.), I.E. $+2' 40''$; required the latitude.

Declin. $0^{\circ} 21' 12''$ N. Latitude $39^{\circ} 57' 25''$ N.

4. September 23rd, in long. $70^{\circ} 14'$ W., the obs. mer. alt. of the sun's L. L., by *artificial horizon*, was $110^{\circ} 25' 10''$ (Z.S.), I.E. $-1' 20''$; required the latitude.

Declin. $0^{\circ} 9' 29''$ S. Latitude $34^{\circ} 42' 11''$ S.

5. September 23rd, in long. $80^{\circ} 20'$ E., the obs. mer. alt. of the sun's U.L., by *artificial horizon*, was $98^{\circ} 30' 40''$ (Z.N.), I.E. $+1' 10''$; required the latitude.

Declin. $0^{\circ} 0' 21''$ N. Latitude $41^{\circ} 1' 10''$ N.

6. September 24th, in long. $4^{\circ} 30'$ W., the obs. mer. alt. of the sun's L.L., by *artificial horizon*, was $84^{\circ} 14' 30''$ (Z.N.), I.E. $-2' 35''$; required the latitude.

Declin. $0^{\circ} 28' 36''$ S. Latitude $47^{\circ} 10' 26''$ N.

LATITUDE BY AN ALTITUDE OF THE POLE STAR.

1. Sept. 21st, at 8h. 40m. P.M. mean time, in long. 33° W., the obs. alt. of the *Pole Star* was $50^{\circ} 50'$, index error $+3'$, height of the eye 23 feet; required the latitude.

R.A. merid. 20h. 42m.
1st corr. $-0^{\circ} 32' 51''$ Latitude $50^{\circ} 18'$ N.

2. Sept. 23rd, at 2h. 45m. A.M. mean time, in long. $154^{\circ} 30'$ W., the obs. alt. *Polaris* was $49^{\circ} 32'$, index error $-2' 55''$, height of the eye 15 feet; required the latitude.

R.A. merid. 2h. 53m. 36s.
1st corr. $-1^{\circ} 15'$ Latitude $48^{\circ} 9'$ N.

3. September 12th, at 8h. 48m. P.M. mean time, in long. 45° W., the obs. alt. of the *Pole Star* was $45^{\circ} 25'$, I.E. $-2' 20''$, height of the eye 17 feet; required the latitude.

R. A. mer. 20h. 15m.
1st corr. $-23' 30''$. Latitude $44^{\circ} 55'$ N.

4. September 13th, at 4h. 0m. A.M. mean time, in long. $148^{\circ} 30'$ E., the obs. alt. of the *Pole Star* was $50^{\circ} 15'$, I.E. $+1' 30''$, height of the eye 19 feet; required the latitude.

R.A. mer. 3h. 26m.
1st corr. $-1^{\circ} 10'$. Latitude $49^{\circ} 2'$ N.

* See Note, page 77.

5. September 15th, at 9h. 40m. P.M. mean time, in long. $35^{\circ} 30' W.$, the obs. alt. of the *Pole Star* was $49^{\circ} 40'$, I.E. + $2' 20''$, height of the eye 24 feet; required the latitude.

R.A. mer. 21h. 19m.

1st corr. - $45'$.

Latitude $48^{\circ} 53' N.$

6. September 17th, at 3h. 35m. A.M. mean time, in long. $48^{\circ} W.$, the obs. alt. of *Polaris* was $35^{\circ} 45'$, I.E. - $3' 30''$, height of the eye 20 feet; required the latitude.

R.A. mer. 3h. 19m.

1st corr. - $1^{\circ} 11'$.

Latitude $34^{\circ} 24' N.$

VARIATION BY AMPLITUDE (SUN).

For Table of Deviations, see NAVIGATION, p. 47.

1. September 19th, at about 5h. 57m. P.M. (app. time), in lat. $22^{\circ} 50' S.$ and long. $38^{\circ} 15' W.$, the sun's setting amplitude by standard compass was $W. 8^{\circ} 15' S.$, deviation $8^{\circ} 50' E.$; required the variation.

Declin. $1^{\circ} 20' 9'' N.$

Corr. mag. amp. $W. 0^{\circ} 35' N.$

True amp. $W. 1^{\circ} 27' N.$

Variation $0^{\circ} 52' E.$

2. September 23rd, at 6h. 0m. A.M. (app. time), in lat. $39^{\circ} 33' N.$ and long. $160^{\circ} 0' W.$, the sun rose by standard compass $E. 22^{\circ} 0' N.$, deviation $6^{\circ} 0' E.$; required the variation.

Declin. $0^{\circ} 9' 24'' S.$

Corr. mag. amp. $E. 16^{\circ} 0' N.$

True amp. $E. 0^{\circ} 12' S.$

Variation $16^{\circ} 12' E.$

3. September 23rd, at about 6h. 0m. A.M. (mean time), in lat. $29^{\circ} 11' S.$ and long. $15^{\circ} 30' E.$, the sun rose by standard compass $S. 68^{\circ} 10' E.$, ship's head being $S.S.W. \frac{1}{4} W.$ by standard compass; required the variation.

Declin. $0^{\circ} 1' 50'' N.$

Corr. mag. amp. $E. 17^{\circ} 5' S.$

True amp. $E. 0^{\circ} 2' N.$

Variation $17^{\circ} 7' W.$

4. September 20th, at 6h. A.M. mean time, in lat. $12^{\circ} S.$ and long. $130^{\circ} 30' W.$, the sun's rising amplitude by standard compass was $E. 10^{\circ} S.$, ship's head west; find the variation.

Declin. $1^{\circ} 2' 28'' N.$

C.M.A. $E. 4^{\circ} N.$

True amp. $E. 1^{\circ} 14' N.$

Var. $2^{\circ} 56' E.$

5. September 22nd, at 6h. A.M. mean time, in lat. $30^{\circ} 15' S.$ and long. $62^{\circ} E.$, the sun rose by standard compass $N. 68^{\circ} E.$, ship's head $N.E.$ by $N.$; find the variation.

Declin. $0^{\circ} 28' N.$

C.M.A. $E. 17^{\circ} N.$

True amp. $E. 0^{\circ} 33' N.$

Var. $16^{\circ} 27' E.$

5. September 16th, at 4h. 25m. P.M. mean time, in lat. $24^{\circ} 40' S.$ and long. $50^{\circ} 25' W.$, the obs. alt. of the sun's L.L. was $17^{\circ} 51' 20''$, I.E. + $2' 30''$, height of the eye 23 feet, when the sun bore by standard compass S. $81^{\circ} 10' W.$, ship's head by standard compass N.E. by $E. \frac{1}{2} E.$; find the variation.

T.B. S. $101^{\circ} 33' W.$ Variation $8^{\circ} 23' E.$
C.M.B. S. $93^{\circ} 10' W.$

6. Sept. 17th, at 4h. om. P.M. mean time, in lat. $15^{\circ} N.$ and long. $21^{\circ} W.$, the obs. alt. of the sun's L.L. was $27^{\circ} 56' 20''$, I.E. + $2' 30''$, height of the eye 20 feet, when the sun bore by standard compass N. $93^{\circ} 31' W.$, the variation at the place of observation being $16^{\circ} 10' W.$; find the deviation for the direction of the ship's head,* which was E. by N. at the time of taking the bearing.

T.B. N. $95^{\circ} 41' W.$ Deviation for E. by N. $14^{\circ} E.$

7. Sept. 20th, at 5h. 30m. P.M. mean time, in lat. $48^{\circ} 45' N.$ and long. $170^{\circ} 20' E.$, the mean of a set of alts. of the sun's L.L. was $15^{\circ} 22' 30''$, index error - $2' 10''$, when the mean of a corresponding set of bearings of the sun's centre by azimuth compass was S. $49^{\circ} 40' W.$, height of the eye 20 feet, the ship's head being N. E. by standard compass, and the correct variation $10^{\circ} 0' E.$; find the deviation of the compass for the direction of the ship's head.

Pol. dist. $88^{\circ} 49' 11''$ True bearing N. $106^{\circ} 27' W.$
Deviation $15^{\circ} 53' E.$

LONGITUDE BY CHRONOMETER. (SUN.)

1. Sept. 15th, at about 4h. 15m. P.M. mean time, in lat. $19^{\circ} 50' N.$ and long. D.R. $39^{\circ} 20' W.$, the obs. alt. of the sun's L.L. was $25^{\circ} 27' 50''$, I.E. + $2'$, height of the eye 19 feet, when a chron. showed 6h. 39m. 45s.; find the longitude.

On August 23rd, at noon, the chron. was 8m. 15s. slow on G.M.T., and gaining daily 3'2s.

Declin. $2^{\circ} 54' 46'' N.$ M.T. ship 4h. 10m. 1s. Sept. 15th.
Eq. T. 4m. 56s. sub. M.T. chr. 6h. 46m. 46s. Sept. 15th.
Long. $39^{\circ} 11' 15'' W.$

2. Sept. 15th, at about 3h. 15m. P.M. mean time, in lat. $35^{\circ} N.$ and long. D.R. $24^{\circ} 30' W.$, the obs. alt. of the sun's L.L. was $32^{\circ} 39' 50''$, I.E. - $1' 25''$, height of the eye 24 feet, when a chron. showed oh. 59m. 10s.; find the longitude.

* See Note, page 73.

On Sept. 3d, at noon, the chron. was slow 3h. 57m. 50s. on G.M.T., and losing daily 4'4s.

Declin. $2^{\circ} 56' 32''$ N.	M.T. ship 3h. 20m. os. Sept. 15th.
Eq. T. 4m. 54s. sub.	M.T. chr. 4h. 57m. 53s. Sept. 15th.
Long. $24^{\circ} 28' 15''$ W.	

3. Sept. 16th, at about 4h. 25m. P.M. mean time, in lat. $24^{\circ} 40'$ S. and long. D.R. $49^{\circ} 25'$ W., the obs. alt. of the sun's L.L. was $17^{\circ} 51' 20''$, I.E. + $2' 30''$, height of the eye 23 feet, when a chron. showed 6h. 30m. 10s.; find the longitude.

Aug. 30th, at noon, the chron. was slow on G.M.T., 1h. 16m. 50s., and gaining daily 1'43s.

Declin. $2^{\circ} 30' 40''$ N.	M.T. ship 4h. 30m. 1s. Sept. 16th.
Eq. T. 5m. 18s. sub.	M.T. chr. 7h. 46m. 35s. Sept. 16th.
Long. $49^{\circ} 8' 30''$ W.	

4. Sept. 17th, at about 4h. 0m. P.M. mean time, in lat. 15° N. and long. D.R. 21° W., the obs. alt. of the sun's L.L. was $27^{\circ} 56' 20''$, I.E. + $2' 30''$, height of the eye 20 feet, when a chron. showed 5h. 10m. 15s.; find the longitude.

On Aug. 28th, at noon, the chron. was 14m. 4s. slow on G.M.T., and losing daily 1'7s.

Declin. $2^{\circ} 9' 43''$ N.	M.T. ship 4h. 0m. os. Sept. 17th.
Eq. T. 5m. 37s. sub.	M.T. chr. 5h. 24m. 53s. Sept. 17th.
Long. $21^{\circ} 13' 15''$ W.	

CHRONOMETER. (STAR, &c.)

1. Sept. 20th, at about 3h. 0m. A.M. mean time, in lat. $10^{\circ} 12'$ N. and long. D.R. $27^{\circ} 30'$ W., the obs. alt. of *Sirius* was $28^{\circ} 38'$ (E. of mer.), index error - $1' 10''$, height of eye 22 feet, when a chron. showed 0h. 44m. 30s.; required the longitude.

On Sept. 1st, at noon, the chron. was 4h. 5m. 12s. slow on G.M.T., and gaining 1'2s. daily.

R.A.M. sun	11h. 55m. 24s.	M.T. chr. 16h. 49m. 19s. Sept. 19th.
Star's hour ang.	20h. 16m. 1s.	Longitude $27^{\circ} 18' 45''$ W.
M.T. ship	15h. 0m. 4s.	Sept. 19th.

2. Sept. 15th, at about 8h. 25m. P.M. mean time, in lat. $35^{\circ} 20'$ S. and long. D.R. $50^{\circ} 30'$ E., the obs. alt. of the *Moon's* L.L. (W. of mer.) was $37^{\circ} 17' 30''$, I.E. - $40''$, height of the eye 26 feet, when a chron. showed 4h. 19m. 54s.; find the longitude.

On Aug. 29th, at noon, the chron. was 42m. 12s. slow on G.M.T., and losing 7·5s. daily.

Moon's hour angle 3h. 39m. 4s. M.T. chr. 5h. 4m. 14s. Sept. 15th.
M.T. ship 8h. 25m. 3s. Sept. 15th. Long. $50^{\circ} 12' 15''$ E.

3. Sept. 16th, at about 8h. 10m. P.M. mean time, in lat. $44^{\circ} 50'$ N. and long. D.R. $54^{\circ} 28'$ W., the obs. alt. of *a Pegasi* was $40^{\circ} 39'$ (E. of mer.), I.E. $+1' 10''$, height of the eye, 23 feet, when a chron. showed oh. 48m. 17s.; find the longitude.

On Sept. 2nd, at noon, the chron. was 1h. 0m. 10s. fast on G.M.T., and losing daily 2·3s.

Star's hour angle 2oh. 54m. 37s. M.T. chr. 11h. 48m. 39s. Sept. 16th.
M.T. ship 8h. 10m. 0s. Sept. 16th. Long. $54^{\circ} 29' 45''$ W.

LATITUDE BY DOUBLE ALTITUDE.

1. Sept. 24th, in lat. D.R. $10^{\circ} 25'$ N. and long. $168^{\circ} 15'$ E., the following observations were taken to find the latitude:—

M. T. Ship.	Chron. times.	Obs. Alts. Sun's L.L.	True bearings.
11h. 15m. A.M.	oh. 5m. 15s.	$75^{\circ} 59' 20''$	S. 39° E.
3 10 P.M.	3 56 30	$39 27 30$	S. 82 W.

Index error $+1' 30''$, height of the eye 18 feet. Run of the ship in the interval S. 30° E. 29 miles.

At G.B.			At L.B.		
P.D.	$90^{\circ} 20' 38''$		P.D.	$90^{\circ} 16' 44''$	
Z.D.	50 20 16		Z.D.	13 18 52	
Arc (1)	57 48 45				
(2)	90 6 45				
(3)	13 38 30				
(5)	65 46 45				
					Polar angle 3h. 51m. 15s.
					Latitude at last obs. $10^{\circ} 9'$ N.

2. September 17th, in lat. by D.R. $36^{\circ} 20'$ N., the following double altitude of two stars was observed at the same instant:—

Obs. Alt. <i>a Leonis</i> .	Bearing.	Obs. Alt. <i>a Orionis</i> .	Bearing.
$45^{\circ} 20' 0''$	S.E.	$55^{\circ} 11' 0''$	S.S.W.
Ind. error $+ 1 10$		$- 2 20$	

Height of the eye 20 feet; required the latitude.

Arc (1) $62^{\circ} 27' 0''$	Arc (3) $38^{\circ} 5' 15''$	Pol. ang. 4h. 13m. 18s.
Arc (2) 88 11 15	Arc (4) 50 6 0	Latitude $36^{\circ} 34'$ N.

3. September 15th, in lat. D.R. $23^{\circ} 20'$ S. and long. $32^{\circ} 19'$ W., the following double altitude was observed :—

M. T. ship, nearly.	Times by chron.	Obs. alt. sun's L. L.	True bearings.
h. m.	h. m. s.	$^{\circ}$ $'$ $''$	
10 12 A.M.	0 23 3	53 27 40	N. 27° E.
2 25 P.M.	4 35 59	45 40 0	N. 58° W.

Height of the eye 24 feet, I.E. $-2' 25''$, run of the ship in the interval W. by N. 6 miles; find the latitude at the last observation.

Arc (1) $63^{\circ} 8' 15''$ Arc (3) $39^{\circ} 24' 15''$ Latitude $22^{\circ} 47' S.$
 Arc (2) 91 53 0 Arc (4) 52 28 45

4. Sept. 18th, in lat. D.R. $9^{\circ} 40'$ N. and long. $24^{\circ} 16'$ W., the following double altitude was taken :—

M. T. ship, nearly.	Times by chron.	Obs. alt. sun's L. L.	True bearings.
h. m.	h. m. s.	$^{\circ}$ $'$ $''$	
8 52 A.M.	10 54 26	43 53 30	S. 83° E.
11 20 A.M.	1 22 38	78 17 20	—

The run of the ship in the interval was N.E. by E. 15 miles, I.E. $-40''$, height of the eye 15 feet; find the latitude at the last observation.

Arc (1) $37^{\circ} 1' 45''$ Arc (3) $11^{\circ} 28' 0''$ Latitude $9^{\circ} 55' N.$
 Arc (2) 89 26 15 Arc (4) 77 58 15

5. September 19th, in lat. by account $26^{\circ} 44'$ N., the following altitudes of two stars were observed at the same moment :—

Obs. Alt. β Orionis.	Bearing.	Obs. Alt. α Hydrae.	Bearing.
$39^{\circ} 40' 0''$	S.W. by S.	$47^{\circ} 20' 0''$	S.E.
Ind. er. $-2 10$		$+ 1 30$	

Height of the eye 15 feet; required the latitude.

Arc (1) $62^{\circ} 0' 0''$ Arc (3) $60^{\circ} 50' 45''$ Pol. ang. 4h. 12m. 53s.
 Arc (2) 95 12 15 Arc (4) 34 21 30 Latitude $26^{\circ} 49' N.$

6. June 10th, in lat. D.R. $40^{\circ} 22'$ S., the following simultaneous altitudes of two stars were taken :—

Obs. alt. <i>Sirius</i> .	Bearing.	Obs. alt. <i>Spica</i> .	Bearing.
$15^{\circ} 25'$	N. 99° W.	$57^{\circ} 1'$	N. 57° W.

Index error $+ 1' 35''$, height of the eye 22 feet; required the latitude.

Arc (1) $96^{\circ} 10' 40''$ Arc (3) $25^{\circ} 20' 15''$ Latitude $40^{\circ} 19' S.$
 Arc (2) 77 7 45 Arc (4) 51 47 30

7. September 16th, in lat. D.R. $25^{\circ} 30'$ N., the following double altitude of two stars was taken :—

Obs. alt. α Tauri.	Bearing.	Obs. alt. β Geminorum.	Bearing.
$74^{\circ} 59'$	S. 44° E.	$37^{\circ} 6'$	S. 107° E.

Index error $- 1' 40''$, height of the eye 30 feet; required the latitude.

Arc (1) $45^{\circ} 2' 0''$ Arc (3) $17^{\circ} 5' 30''$ Latitude $25^{\circ} 31' 52'' N.$
 Arc (2) 95 8 45 Arc (4) 78 3 15

8. September 17th, in lat. D.R. $54^{\circ} 20' N.$, the following simultaneous altitudes of two stars were taken:—

True alt. α <i>Andromeda</i> .	Bearing.	True alt. α <i>Tauri</i> .	Bearing.
$51^{\circ} 18'$	S. $24^{\circ} E.$	$46^{\circ} 46'$	S. $33^{\circ} E.$
Arc (1) $62^{\circ} 9'$	Arc (3) $43^{\circ} 22'$	Latitude $54^{\circ} 8' N.$	
Arc (2) $66\ 3$	Arc (4) $22\ 41$		

LATITUDE BY MERIDIAN ALTITUDE OF THE MOON.

1. September 15th, in long. $27^{\circ} W.$, the obs. mer. alt. of the moon's L.L. was $54^{\circ} 15' 30''$ (Z.N.), I.E. $-3' 30''$, height of the eye 19 feet; required the latitude.

G.D. 6h. 36m. Sept. 15th.

Declin. $16^{\circ} 52' 50'' S.$

Latitude $18^{\circ} 13' 35'' N.$

2. September 16th, in long. $35^{\circ} W.$, the obs. mer. alt. of the moon's L.L. was $83^{\circ} 10' 20''$ (Z.S.), I.E. $+2' 10''$, height of the eye 29 feet; required the latitude.

G.D. 7h. 57m. Sept. 16th.

Declin. $17^{\circ} 58' 51'' S.$

Latitude $24^{\circ} 30' S.$

3. September 17th, in long. $32^{\circ} E.$, the obs. mer. alt. of the moon's L.L. was $38^{\circ} 17' 40''$ (Z.S.), I.E. $-50''$, height of the eye 30 feet; required the latitude.

G.D. 4h. 8m. Sept. 17th.

Declin. $18^{\circ} 10' 25'' S.$

Latitude $69^{\circ} 2' 18'' S.$

4. September 18th, in long. $45^{\circ} W.$, the obs. mer. alt. of the moon's U.L. was $23^{\circ} 18' 20''$ (Z.N.), I.E. $+1' 40''$, height of the eye 20 feet; required the latitude.

G.D. 10h. 15m. Sept. 18th.

Declin. $17^{\circ} 24' 47'' S.$

Latitude $48^{\circ} 46' 13'' N.$

5. September 19th, in long. $54^{\circ} E.$, the obs. mer. alt. of the moon's U.L. was $40^{\circ} 10' 30''$ (Z.N.), I.E. $-1' 15''$, height of the eye 20 feet; required the latitude.

G.D. 4h. 15m. Sept. 19th.

Declin. $16^{\circ} 19' 13'' S.$

Latitude $33^{\circ} 9' 38'' N.$

LONGITUDE BY LUNAR.

1. September 20th, at 11h. 20m. P.M. mean time, in lat. $18^{\circ} 50'$ S. and long. by D.R. 90° E., the following lunar was taken :—

Obs. alt. <i>α Arietis</i> (E. of mer.)	Obs. alt. moon's L.L.	Obs. dist. F.L.
$32^{\circ} 36' 30''$	$51^{\circ} 0' 0''$	$87^{\circ} 24' 25''$
I.E. + 1 10	— 2 10	+ 0 20

height of the eye 28 feet ; required the longitude.

Star's declin.	$22^{\circ} 49' 50''$ N.	True dist.	$86^{\circ} 42' 47''$
Star's R.A.	1h. 59m. 41s.	M.T. Green.	5h. 19m. 15s. Sept. 20th.
R.A.M. sun	11h. 57m. 28s.	M.T. ship	11h. 20m. os. Sept. 20th.
	Longitude $90^{\circ} 11' 15''$ E.		

2. September 23rd, at oh. 40m. A.M. mean time, in lat. $50^{\circ} 15'$ N. and long. by account $45^{\circ} 8'$ W., the following lunar was taken :—

Obs. alt. <i>Aldebaran</i> (E. of mer.)	Obs. alt. moon's L.L.	Obs. dist. F.L.
$34^{\circ} 28' 40''$	$27^{\circ} 22' 50''$	$87^{\circ} 51' 10''$
I.E. — 1 20	+ 0 30	+ 1 5

height of the eye 20 feet ; required the longitude.

Star's R.A.	4h. 28m. 17s.	True dist.	$87^{\circ} 6' 5''$
Star's dec.	$16^{\circ} 14' 14''$ N.	M.T. Green.	15h. 40m. 15s. Sept. 22nd.
R.A.M. sun	12h. 7m. 3s.	M.T. ship	12h. 39m. 59s. Sept. 22nd.
	Longitude $45^{\circ} 4' 0''$ W.		

3. September 15th, at 3h. 20m. P.M. mean time, in lat. 35° N. and long. D.R. $25^{\circ} 30'$ W., the following lunar was taken :—

Obs. alt. sun's L.L.	Obs. alt. moon's L.L.	Obs. dist.
$32^{\circ} 39' 50''$	$33^{\circ} 25' 20''$	$74^{\circ} 11' 10''$
I.E. — 1 25	+ 1 40	+ 1 5

height of the eye 24 feet ; required the longitude.

Declin.	$2^{\circ} 56' 30''$ N.	M.T. ship	3h. 20m. os. Sept. 15th.
Eq. T.	4m. 54s. sub.	M.T. Green.	5h. 4m. 43s. Sept. 15th.
	True dist. $74^{\circ} 23' 18''$.	Longitude $26^{\circ} 10' 15''$ W.	

4. September 17th, at 4h. om. P.M. mean time, in lat. 15° N. and long. D.R. 21° W., the following lunar was taken :—

Obs. alt. sun's L.L.	Obs. alt. moon's L.L.	Obs. dist.
$27^{\circ} 56' 20''$	$41^{\circ} 33' 40''$	$96^{\circ} 20' 10''$
I.E. + 2 30	— 1 15	+ 1 0

height of the eye 20 feet ; required the longitude.

Declin.	$2^{\circ} 9' 43''$ N.	M.T. ship	4h. om. 1s. Sept. 17th.
Eq. T.	5m. 37s. sub.	M.T. Green.	5h. 24m. 56s. Sept. 17th.
	True dist. $96^{\circ} 24' 2''$	Longitude $21^{\circ} 13' 45''$ W.	

5. September 16th, at 8h. 10m. P.M. mean time, in lat. $44^{\circ} 50' N.$
and long. D.R. $54^{\circ} 30' W.$, the following lunar was taken :—

Obs. alt. α <i>Pegasi</i> (E. of mer.)	Obs. alt. moon's L.L.	Obs. dist. F.L.
$40^{\circ} 38' 50''$	$18^{\circ} 0' 10''$	$88^{\circ} 35' 30''$
I.E. + 1 10	- 0 55	- 1 10

height of the eye 23 feet ; required the longitude.

M.T. ship	8h. 9m. 58s. Sept. 16th.	True dist. $87^{\circ} 7' 47''$
M.T. Green.	11h. 47m. 18s. Sept. 16th.	Long. $54^{\circ} 20' 0'' W.$

6. September 17th, at 8h. 20m. P.M. mean time, in lat. $33^{\circ} 55' S.$
and long. D.R. $62^{\circ} 30' E.$, the following lunar was taken :—

Obs. alt. α <i>Pegasi</i> (E. of mer.)	Obs. alt. moon's L.L.	Obs. dist. F.L.
$26^{\circ} 35' 10''$	$57^{\circ} 36' 0''$	$80^{\circ} 32' 30''$
I.E. + 1 15	- 1 40	- 0 30

height of the eye 18 feet ; required the longitude.

M.T. ship	8h. 20m. os. Sept. 17th.	True dist. $80^{\circ} 2' 1''$
M.T. Green.	4h. 10m. 4s. Sept. 17th.	Long. $62^{\circ} 29' E.$

FIXED STARS.

MEAN PLACES FOR JANUARY, 1866.

Star's Name.	Mag.	Right Ascension.			Annual Var.	Declination.			Annual Var.	
		h.	m.	s.		°	'	"		
α Andromedæ	2	0	1	28	+3'1	N.	28	21	2	+19'9
γ Pegasi (<i>Algenib</i>)	3'2	0	6	20	3'1	N.	14	26	18	20'0
β Ceti	2	0	36	52	3'0	S.	18	43	22	19'8
α Ursæ Min. (<i>Polaris</i>)...	2	1	9	58	19'6	N.	88	35	42	19'1
α Eridani (<i>Achernar</i>) ...	1	1	32	43	2'2	S.	57	55	5	18'4
α Arietis	2	1	59	37	3'4	N.	22	49	38	17'2
α Ceti	2'3	2	55	17	3'1	N.	3	33	43	14'4
α Persei	2	3	14	46	4'2	N.	49	22	52	13'2
α Tauri (<i>Aldebaran</i>) ...	1	4	28	14	3'4	N.	16	14	14	7'6
α Aurigæ (<i>Capella</i>)	1	5	6	48	4'4	N.	45	51	28	4'2
β Orionis (<i>Rigel</i>)	1	5	8	6	2'9	S.	8	21	33	4'5
β Tauri	2	5	17	49	3'8	N.	28	29	27	3'5
α Argus (<i>Canopus</i>)	1	6	20	59	1'3	S.	52	37	25	- 1'8
α Canis Maj. (<i>Sirius</i>) ...	1	6	39	15	2'6	S.	16	32	6	4'6
α^2 Gemin. (<i>Castor</i>)	2'1	7	26	3	3'8	N.	32	10	45	7'4
α Canis. Min. (<i>Procyon</i>)	1	7	32	17	3'1	N.	5	33	58	8'9
β Gemin. (<i>Pollux</i>)	1'2	7	37	7	3'7	N.	28	20	49	8'3
α Hydræ	2	9	21	0	2'9	S.	8	4	46	15'4
α Leonis (<i>Regulus</i>)	1'2	10	1	14	3'2	N.	12	37	15	17'4
γ Argûs	2	10	39	52	2'3	S.	58	58	48	18'7
α Ursæ Majoris	2	10	55	26	3'8	N.	62	28	25	19'3
β Leonis	2	11	42	13	3'1	N.	15	19	16	20'1
α^1 Crucis	1	12	19	10	3'3	S.	62	21	18	19'9
α Virginis (<i>Spica</i>)	1	13	18	8	3'1	S.	10	27	40	18'9
γ Ursæ Majoris	2	13	42	15	2'4	N.	49	58	59	18'1
β Centauri	1	13	54	24	4'1	S.	59	43	29	17'7
α Boötis (<i>Arcturus</i>)	1	14	9	33	2'7	N.	19	52	53	18'9
α^2 Centauri	1	14	30	32	4'0	S.	60	16	39	15'0
β Ursæ Minoris	2	14	51	0	-0'2	N.	74	42	10	14'7
α Coronæ Borealis	2	15	29	1	+2'5	N.	27	10	3	12'3
α Serpentis	2'3	15	37	40	2'9	N.	6	50	58	11'6
α Scorpii (<i>Antares</i>)	1'2	16	21	12	3'7	S.	26	7	54	8'4
α Trianguli Australis ...	2	16	34	30	6'3	S.	68	46	34	7'4
α Ophiuchi	2	17	28	43	2'8	N.	12	39	36	- 2'9
α Lyræ (<i>Vega</i>)	1	18	32	24	2'0	N.	38	39	39	+ 3'1
α Aquilæ (<i>Altair</i>)	1'2	19	44	15	2'9	N.	8	31	0	9'1
α Pavonis	2	20	15	2	4'8	S.	57	9	38	11'1
α Cygni	2'1	20	36	52	2'0	N.	44	48	10	12'7
α Gruis	2	21	59	46	3'8	S.	47	36	29	17'2
α Pis. Aus. (<i>Fomalhaut</i>)	1'2	22	50	14	3'3	S.	30	19	54	19'0
α Pegasi (<i>Markab</i>)	2	22	58	5	+3'0	N.	14	29	6	+19'3

AT APPARENT NOON, SEPTEMBER, 1866.

Day of the Month.	THE SUN'S				Equation of Time, to be subtracted from Apparent Time.	Diff. for 1 hour.
	Apparent Right Ascension.	Diff. for 1 hour.	Apparent Declination.	Diff. for 1 hour.		
1	h. m. s. 10 41 33 ⁶⁴	s. 9 ⁰⁷²	N. 8 17 3 ⁸	" 54 ⁴⁶	m. s. 0 7 ³⁷	s. 0 ⁷⁸²
2	10 45 11 ²⁴	9 ⁰⁶²	7 55 12 ⁶	54 ⁸⁰	0 26 ²⁶	0 ⁷⁹³
3	10 48 48 ⁶⁰	9 ⁰⁵¹	7 33 13 ⁶	55 ¹¹	0 45 ⁴²	0 ⁸⁰³
4	10 52 25 ⁷¹	9 ⁰⁴¹	7 11 7 ²	55 ⁴¹	1 4 ⁸⁰	0 ⁸¹³
5	10 56 2 ⁵⁹	9 ⁰³²	6 48 53 ⁷	55 ⁷⁰	1 24 ⁴³	0 ⁸²²
6	10 59 39 ²⁵	9 ⁰²³	6 26 33 ⁴	55 ⁹⁸	1 44 ²⁶	0 ⁸³⁰
7	11 3 15 ⁷¹	9 ⁰¹⁵	6 4 6 ⁷	56 ²⁴	2 4 ²⁹	0 ⁸³⁸
8	11 6 51 ⁹⁸	9 ⁰⁰⁸	5 41 34 ⁰	56 ⁴⁹	2 24 ⁵²	0 ⁸⁴⁶
9	11 10 28 ⁰⁸	9 ⁰⁰¹	5 18 55 ⁵	56 ⁷²	2 44 ⁹²	0 ⁸⁵³
10	11 14 4 ⁰³	8 ⁹⁹⁵	4 56 11 ⁶	56 ⁹³	3 5 ⁴⁷	0 ⁸⁵⁹
11	11 17 39 ⁸³	8 ⁹⁸⁹	4 33 22 ⁷	57 ¹⁴	3 26 ¹⁶	0 ⁸⁶⁴
12	11 21 15 ⁵²	8 ⁹⁸⁵	4 10 29 ⁰	57 ³³	3 46 ⁹⁶	0 ⁸⁶⁹
13	11 24 51 ¹¹	8 ⁹⁸¹	3 47 31 ⁰	57 ⁵⁰	4 7 ⁸⁷	0 ⁸⁷³
14	11 28 26 ⁶¹	8 ⁹⁷⁸	3 24 29 ⁰	57 ⁶⁶	4 28 ⁸⁷	0 ⁸⁷⁶
15	11 32 2 ⁰⁴	8 ⁹⁷⁵	3 1 23 ³	57 ⁸¹	4 49 ⁹³	0 ⁸⁷⁹
16	11 35 37 ⁴¹	8 ⁹⁷³	2 38 14 ²	57 ⁹⁴	5 11 ⁰⁴	0 ⁸⁸¹
17	11 39 12 ⁷⁶	8 ⁹⁷²	2 15 2 ¹	58 ⁰⁶	5 32 ²⁰	0 ⁸⁸²
18	11 42 48 ⁰⁹	8 ⁹⁷²	1 51 47 ³	58 ¹⁶	5 53 ³⁶	0 ⁸⁸²
19	11 46 23 ⁴³	8 ⁹⁷³	1 28 30 ³	58 ²⁵	6 14 ⁵²	0 ⁸⁸¹
20	11 49 58 ⁷⁹	8 ⁹⁷⁵	1 5 11 ²	58 ³³	6 35 ⁶⁴	0 ⁸⁷⁹
21	11 53 34 ²¹	8 ⁹⁷⁷	0 41 50 ⁵	58 ³⁹	6 56 ⁷²	0 ⁸⁷⁷
22	11 57 9 ⁷⁰	8 ⁹⁸⁰	N. 0 18 28 ⁴	58 ⁴⁴	7 17 ⁷³	0 ⁸⁷³
23	12 0 45 ²⁷	8 ⁹⁸⁵	S. 0 4 54 ⁶	58 ⁴⁸	7 38 ⁶⁵	0 ⁸⁶⁹
24	12 4 20 ⁹⁷	8 ⁹⁹⁰	0 28 18 ⁴	58 ⁵⁰	7 59 ⁴⁴	0 ⁸⁶⁴
25	12 7 56 ⁸¹	8 ⁹⁹⁷	0 51 42 ⁶	58 ⁵¹	8 20 ¹⁰	0 ⁸⁵⁷
26	12 11 32 ⁸³	9 ⁰⁰⁵	1 15 6 ⁷	58 ⁵⁰	8 40 ⁵⁸	0 ⁸⁴⁹
27	12 15 9 ⁰⁴	9 ⁰¹⁴	1 38 30 ⁷	58 ⁴⁹	9 0 ⁸⁵	0 ⁸⁴⁰
28	12 18 45 ⁴⁹	9 ⁰²⁴	2 1 54 ¹	58 ⁴⁶	9 20 ⁹⁰	0 ⁸³⁰
29	12 22 22 ¹⁸	9 ⁰³⁴	2 25 16 ⁶	58 ⁴¹	9 40 ⁷²	0 ⁸²⁰
30	12 25 59 ¹⁴	9 ⁰⁴⁶	2 48 37 ⁸	58 ³⁵	10 0 ²⁵	0 ⁸⁰⁸
31	12 29 36 ³⁹	9 ⁰⁵⁹	S. 3 11 57 ⁴	58 ²⁸	10 19 ⁵⁰	0 ⁷⁹⁶

AT MEAN NOON, SEPTEMBER, 1866.

Day of the Month:	THE SUN'S			Equation of Time, to be added to Mean Time.	Sidereal Time.
	Apparent Right Ascension.	Apparent Declination.	Semidiam.*		
	h. m. s.	N. ° ' "	' " "	m. s.	h. m. s.
1	10 41 33.66	N. 8 17 3.7	15 53.6	0 7.37	10 41 41.03
2	10 45 11.31	7 55 12.2	15 53.8	0 26.27	10 45 37.59
3	10 48 48.71	7 33 12.9	15 54.0	0 45.43	10 49 34.14
4	10 52 25.87	7 11 6.2	15 54.3	1 4.82	10 53 30.69
5	10 56 2.80	6 48 52.4	15 54.5	1 24.45	10 57 27.24
6	10 59 39.51	6 26 31.8	15 54.7	1 44.28	11 1 23.79
7	11 3 16.02	6 4 4.8	15 54.9	2 4.32	11 5 20.34
8	11 6 52.34	5 41 31.7	15 55.2	2 24.55	11 9 16.90
9	11 10 28.49	5 18 52.9	15 55.4	2 44.96	11 13 13.45
10	11 14 4.49	4 56 8.7	15 55.7	3 5.51	11 17 10.00
11	11 17 40.35	4 33 19.4	15 55.9	3 26.21	11 21 6.56
12	11 21 16.09	4 10 25.4	15 56.2	3 47.02	11 25 3.11
13	11 24 51.73	3 47 27.1	15 56.4	4 7.93	11 28 59.66
14	11 28 27.28	3 24 24.7	15 56.7	4 28.93	11 32 56.21
15	11 32 2.76	3 1 18.6	15 57.0	4 50.00	11 36 52.76
16	11 35 38.19	2 38 9.2	15 57.3	5 11.12	11 40 49.31
17	11 39 13.59	2 14 56.7	15 57.5	5 32.28	11 44 45.87
18	11 42 48.97	1 51 41.6	15 57.8	5 53.45	11 48 42.42
19	11 46 24.36	1 28 24.2	15 58.1	6 14.61	11 52 38.97
20	11 49 59.78	1 5 4.8	15 58.4	6 35.74	11 56 35.52
21	11 53 35.25	0 41 43.7	15 58.6	6 56.82	12 0 32.07
22	11 57 10.79	N. 0 18 21.3	15 58.9	7 17.84	12 4 28.62
23	12 0 46.42	S. 0 5 2.1	15 59.2	7 38.76	12 8 25.18
24	12 4 22.17	0 28 26.2	15 59.5	7 59.56	12 12 21.73
25	12 7 58.06	0 51 50.7	15 59.7	8 20.22	12 16 18.28
26	12 11 34.13	1 15 15.2	16 0.0	8 40.70	12 20 14.83
27	12 15 10.40	1 38 39.5	16 0.3	9 0.98	12 24 11.38
28	12 18 46.90	2 2 3.2	16 0.6	9 21.03	12 28 7.93
29	12 22 23.64	2 25 26.0	16 0.8	9 40.85	12 32 4.49
30	12 26 0.65	2 48 47.5	16 1.1	10 0.39	12 36 1.04
31	12 29 37.95	S. 3 12 7.4	16 1.4	10 19.64	12 39 57.59

* The Semidiameter for *Apparent Noon* may be assumed the same as that for *Mean Noon*.

ELEMENTS from the NAUTICAL ALMANAC, 1866.
MEAN TIME.

SEPTEMBER.

Day of the Month.	THE MOON'S				Meridian Passage.
	Semidiameter.		Horizontal Parallax.		
	<i>Noon.</i>	<i>Midnight.</i>	<i>Noon.</i>	<i>Midnight.</i>	
15	' "	' "	' "	' "	h. m.
15	14 47'3	14 48'0	54 11'0	54 13'5	4 45
16	14 49'5	14 51'6	54 18'7	54 26'4	5 32
17	14 54'4	14 57'9	54 36'7	54 49'5	6 20
18	15 2'0	15 6'8	55 4'8	55 22'3	7 9
19	15 12'2	15 18'0	55 41'9	56 3'3	7 58
20	15 24'2	15 30'8	56 26'2	56 50'2	8 48
21	15 37'5	15 44'3	57 14'9	57 39'8	9 39
22	15 51'0	15 57'5	58 4'4	58 28'2	10 29

ELEMENTS from the NAUTICAL ALMANAC, 1866.
MEAN TIME.

LUNAR DISTANCES—SEPTEMBER.

Day of the Month.	Star's Name and Position.	<i>Noon.</i>	P.L. of diff.	III ^b .	P.L. of diff.	VI ^b .	P.L. of diff.	IX ^b .	P.L. of diff.
15	Sun, W.....	72 5 33	3448	73 26 55	3448	74 48 17	3447	76 9 41	3445
16	Sun, W.....	82 57 23	3428	84 19 8	3423	85 40 58	3418	87 2 54	3412
16	α Pegasi, E.	93 19 52	3282	91 55 20	3277	90 30 42	3272	89 5 58	3267
17	Sun, W.....	93 54 29	3374	95 17 15	3365	96 40 11	3356	98 3 18	3345
17	α Pegasi, E.	82 0 43	3238	80 35 19	3231	79 9 47	3225	77 44 8	3219
20	α Arietis, E.	89 30 11	2811	87 55 57	2795	86 21 22	2779	84 46 26	2763
22	Antares, W.	84 23 12	2495	86 4 32	2479	87 46 14	2464	89 28 18	2449
22	Aldebaran, E.	96 6 36	2437	94 23 54	2422	92 40 51	2406	90 57 25	2391

**ELEMENTS from the NAUTICAL ALMANAC, 1866.
MEAN TIME.**

THE MOON'S RIGHT ASCENSION AND DECLINATION.

Hour.	Right Ascension.	Declination.	Diff. Dec. for 10m.	Hour.	Right Ascension.	Declination.	Diff. Dec. for 10m.
SEPTEMBER 15TH.				SEPTEMBER 18TH.			
	h. m. s.				h. m. s.		
5	16 23 32.6	16 47 53 S.	36.44	10	19 4 59.6	17 25 29.0 S.	28.25
6	16 25 35.2	16 50 41.7	35.68	11	19 7 8.2	17 22 36.9	29.13
7	16 27 37.9	16 54 13.5	34.92	SEPTEMBER 19TH.			
SEPTEMBER 16TH.				4	19 43 44.0	16 20 19.4 S.	44.12
7	17 17 9.1	17 55 32.6 S.	15.90	5	19 45 53.6	16 15 52.1	44.99
8	17 19 14.0	17 57 5.5	15.08	SEPTEMBER 23RD.			
SEPTEMBER 17TH.				22	23 55 9.0	0 7 37.6 S.	111.84
4	18 1 9.7	18 10 27.2 S.	1.85	23	23 57 25.3	0 3 33.7 N.	111.93
5	18 3 16.3	18 10 13.5	2.72				

**ELEMENTS from the NAUTICAL ALMANAC, 1866.
MEAN TIME.**

LUNAR DISTANCES—SEPTEMBER.

Day of the Month	Star's Name and Position.	Midnight.	P.L. of diff.	XV ^b .	P.L. of diff.	XVIII ^b .	P.L. of diff.	XXI ^b .	P.L. of diff.
15	Sun, W.....	77 31 7	3442	78 52 36	3440	80 14 8	3437	81 35 43	3432
16	Sun, W.....	88 24 57	3405	89 47 8	3399	91 9 26	3391	92 31 53	3383
16	α Pegasi, E.	87 41 8	3262	86 16 12	3256	84 51 9	3251	83 26 0	3244
17	Sun, W.....	99 26 37	3336	100 50 7	3325	102 13 50	3313	103 37 47	3301
17	α Pegasi, E.	76 18 21	3213	74 52 27	3206	73 26 25	3199	72 0 14	3193
20	α Arietis, E.	83 11 9	2747	81 35 31	2731	79 59 32	2716	78 23 13	2700
22	Antares, W.	91 10 43	2434	92 53 30	2419	94 36 38	2405	96 20 6	2391
22	Aldebaran, E.	89 13 37	2376	87 29 27	2362	85 44 57	2347	84 0 5	2333

TABLES FOR DETERMINING THE LATITUDE BY OBSERVATIONS OF THE POLE STAR OUT OF THE MERIDIAN.

TABLE I. (A.)

Containing the *First* Correction.

Argument : Sidereal Time of Observation.

Sidereal Time of Obs.	Correction.	Sidereal Time of Obs.	Sidereal Time of Obs.	Correction.	Sidereal Time of Obs.
h. m.		h. m.	h. m.		h. m.
2 0	- 1 22 1 +	14 0	8 0	+ 0 18 11 -	20 0
10	1 21 8	10	10	0 21 44	10
20	1 20 7	20	20	0 25 16	20
30	1 18 56	30	30	0 28 44	30
40	1 17 36	40	40	0 32 9	40
50	1 16 8	50	50	0 35 30	50
3 0	1 14 31	15 0	9 0	0 38 47	21 0
10	1 12 45	10	10	0 42 0	10
20	1 10 51	20	20	0 45 8	20
30	1 8 49	30	30	0 48 11	30
40	1 6 39	40	40	0 51 8	40
50	1 4 21	50	50	0 54 0	50

TABLE II. (A.)

Containing the *Second* Correction (*always to be added*).

Arguments :—Sidereal Time and Altitude.

Sidereal Time of Obs.	ALTITUDE.								Sidereal Time of Obs.
	15°	20°	25°	30°	35°	40°	45°	50°	
h. m.	' "	' "	' "	' "	' "	' "	' "	' "	h. m.
2 0	0 1	0 1	0 1	0 2	0 2	0 2	0 3	0 3	14 0
3 0	0 4	0 5	0 6	0 8	0 9	0 11	0 13	0 16	15 0
4 0	0 8	0 10	0 13	0 16	0 20	0 24	0 28	0 33	16 0
5 0	0 12	0 16	0 20	0 25	0 31	0 37	0 44	0 52	17 0
6 0	0 15	0 20	0 26	0 32	0 39	0 47	0 56	1 7	18 0
7 0	0 16	0 22	0 29	0 36	0 43	0 52	1 1	1 13	19 0
8 0	0 16	0 21	0 27	0 34	0 41	0 49	0 59	1 10	20 0
9 0	0 13	0 18	0 23	0 28	0 34	0 41	0 48	0 58	21 0
10 0	0 9	0 12	0 16	0 19	0 23	0 28	0 33	0 40	22 0

TABLE III. (A.)

Containing the *Third* Correction (*always to be added*).

Arguments :—Sidereal Time and Date.

Sidereal Time of Obs.	March 1.	April 1.	June 1.	Sept. 1.
h.	' "	' "	' "	' "
2	0 58	0 49	0 32	0 36
4	1 1	0 57	0 39	0 28
14	1 2	1 11	1 28	1 24
18	0 56	0 55	1 9	1 32
20	0 54	0 48	0 54	1 23
22	0 53	0 44	0 41	1 7

ADDITIONAL EXAMPLES.

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LONGITUDE BY CHRONOMETER.

1. September 10th, at about 8h. 10m. A.M., in lat. $30^{\circ} 10' S.$ and long. D.R. $45^{\circ} 30' W.$, when a chron. showed oh. 30m. 10s., the obs. alt. of the sun's L.L. was $24^{\circ} 40' 30''$, index error $+1' 30''$, height of the eye 20 feet: required the longitude.

August 20th, the chron. was fast on G.M.T. 1h. 20m. 10s., and on August 30th, it was fast on G.M.T. 1h. 20m. 30s.

See (Note).—The rate to be ascertained from these two errors.

Ans. Long. $45^{\circ} 31' 30'' W.$

2. September 23rd, at about 8h. 10m. A.M., in lat. $39^{\circ} 50' N.$, and long. D.R. $152^{\circ} E.$, when a chron. showed 5h. 30m. 25s., the obs. alt. of the Sun's L.L. was $21^{\circ} 52' 15''$, index error $+2' 10''$, height of the eye 24 feet: required the longitude.

August 15th, the chron. was slow on G.M.T. 4h. 12m. 30.5s., and on August 25th, it was slow on G.M.T. 4h. 12m. 5.5s.

Ans. Long. $151^{\circ} 57' E.$

3. September 27th, at 3h. 30m. P.M., in lat. $49^{\circ} 51' N.$, and long. D.R. $96^{\circ} 30' W.$, when a chron. showed 1h. 30m. 20s., the obs. alt. of Sun's L.L. was $19^{\circ} 50' 30''$, index error $+2' 30''$, height of the eye 18 feet: required the longitude.

July 30th, at noon, the chron. was fast on G.M.T. 3h. 30m. 40s., and on August 29th, it was fast on G.M.T. 3h. 31m. 10.5s.

Ans. Long. $96^{\circ} 38' 30'' W.$

4. September 17th, at about 7h. 0m. A.M., in lat. $20^{\circ} 10' N.$, and long. D.R. $37^{\circ} 30' E.$, the following observation for longitude was taken:—

Time by Chron.	Obs. Alt. Sun's L.L.
----------------	----------------------

3h. 7m. 10s.	$16^{\circ} 10' 10''$
--------------	-----------------------

Index error $+1' 10''$.	Dip, 20 feet.
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On September 2nd, at noon, the chron. was slow on G.M.T. 1h. 22m. 19s., and on September 10th, it was slow on G.M.T. 1h. 22m. 59s.

Ans. Long. $37^{\circ} 30' E.$

5. September 23rd, at about 7h. 0m. A.M., in lat. $34^{\circ} 0' N.$, and long. D.R. $1^{\circ} 0' W.$, the following observation was taken to determine the longitude:—

Time by Chron.	Obs. Alt. Sun's L.L.
----------------	----------------------

5h. 56m. 55s.	$15^{\circ} 53' 0''$
---------------	----------------------

Index error $-1' 10''$	Dip, 18 feet.
------------------------	---------------

On August 30th, at noon, the chron. was slow on G.M.T. 1h. 11m. 48s., and on September 15th, it was slow on G.M.T. 1h. 10m. 6s.

Ans. Long. $1^{\circ} 1' W.$

6. September 10th, at about 7h. 30m. A.M., in lat. $32^{\circ} S.$, and long. D.R. $68^{\circ} 15' W.$, the following observation was taken to determine the longitude:—

Time by Chron.	Obs. Alt. Sun's L.L.
----------------	----------------------

11h. 57m. 48s.	$15^{\circ} 51' 10''$
----------------	-----------------------

Index error $+0' 30''$.	Dip, 19 feet.
--------------------------	---------------

On August 10th, at noon, the chron. was fast on G.M.T. oh. 0m. 54s., and on August 30th, it was slow on G.M.T. oh. 1m. 6s.

Ans. Long. $68^{\circ} 25' 15'' W.$

ERROR OF CHRONOMETER.

1. September 14th, A.M., in lat. $50^{\circ} 21' 36''$ N., and long. $3^{\circ} 35'$ W. the following sights were taken to find the error of chronometer:—

Time by Chron.	Obs. Alt.	Sun's L.L. (art. hor.)
8h. 47m. os.		$53^{\circ} 48' 50''$
8 47 20		53 53 40
8 47 39		53 58 40
8 47 54		54 3 30
8 48 10		54 7 30

Index error $+0' 45''$.

Ans. Error, 3m. 7.6s. slow on G.M.T.

2. September 13th, A.M., in lat. $50^{\circ} 21' 36''$ N., and long. $3^{\circ} 35'$ W., the following sights were taken to find the error of chronometer:—

Time by Chron.	Obs. Alt.	Sun's L.L. (art. hor.)
9h. 21m. 14s.		$63^{\circ} 13' 20''$
9 21 25		63 16 10
9 21 35		63 18 10
9 21 48		63 21 20
9 21 59		63 24 0

Index error 0.

Ans. Error, 2m. 44.2s. slow on G.M.T.

3. September 25th, in lat. $53^{\circ} 21' 36''$ N., and long. $3^{\circ} 35'$ W., the following observation was taken to find the error of the chronometer:—

Time by Watch.	Obs. Alt.	Sun's L.L. (art. hor.)
9h. 31m. 27s. (A.M.)		$57^{\circ} 39' 40''$

Index error $-0' 30''$.

Watch fast of chron. 1h. 42m. 18s. On September 13th, chron. was slow on G.M.T. 1h. 38m. 41s.

Find the error and also the rate.

Ans. Error, 1h. 41m. 33s. slow on G.M.T.

Rate, 14.33 seconds, losing.

Note.—When the errors on two different days are given, from which the daily rate has to be determined, if the errors are of the *same* name, that is, both *fast*, or both *slow*, it is merely necessary to take one error from the other, and divide the *difference* by the number of days elapsed between the dates of the two errors, the quotient will be the daily rate.

(a) If both errors are *fast*, the rate will be *gaining* when the *last* error was the greater; and *losing* when the *last* error was the lesser of the two.

(b) If both errors are *slow*, the rate will be *losing* when the *last* error was the greater; and *gaining* when the *last* error was the lesser of the two.

(c) When the errors on the two given days are of *contrary* names, that is, when one is fast and the other slow, or when one is slow and the other fast, the "rate" will be ascertained by dividing the *sum* of the errors by the number of days elapsed between the dates of the two errors.

(d) The "rate" will be *gaining* if the chronometer has changed from slow to fast; and *losing* if the chronometer has changed from fast to slow.



